# Background model for a Nal(TI) detector in the frame of the ANAIS experiment

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(On behalf of the ANAIS Collaboration)

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# OUTLINE

• Brief Introduction and Present Status of the ANAIS Experiment

• Filtering of low energy events

• BACKGROUND MODEL FOR THE ANAIS-0 PROTOTYPE

CONCLUSIONS AND PROSPECTS

#### The ANAIS Experiment

to cm neutron shielding anti-Rin box 2 mm Cd Vibration isolator

250 kg of ultrapure Nal(TI) detectors at Canfranc Underground Lab.

Goal: study the annual modulation in the dark matter signal

SAME TARGET AND TECHNIQUE AS DAMA/LIBRA

#### The ANAIS Experiment

Long effort in R+D at the University of Zaragoza

HPGe radiopurity test bench for material selection
 PMT Testing
 Background understanding
 Optimization of data analysis and readout





## New radiopure Nal(TI) crystals

- HPGe spectrometry screening of K content in Nal powder from different providers.
- Two 12.5 kg cylindrical crystals have been grown by AS with the best raw powder and are being encapsulated.
  - They should be received this summer for final background
    - assessment at LSC. We require no more than 0.02 ppm Potassium.



#### New radiopure Nal(TI) crystals

#### EVERYTHING AT LSC IS READY FOR THE MOUNTING





# At the Canfranc Underground Laboratory – HALL B

Lead and polyethylene ready for the shielding of ANAIS



Access to ANAIS control room

ANAIS control room

Removable walls allow for convenient access to the experimental space

# At the Canfranc Underground Laboratory – SHIELDING of ANAIS-0



- 10 cm archaeological lead
- 20 cm low activity lead
- Tightly closed box kept with boil-off
  nitrogen overpressure
- Three plastic scintillators as muon vetoes



#### **ANAIS-0** Prototype





ANAIS-0 module consists of a 9,6 kg Nal(TI) made by Saint Gobain 4"x4"x10" Encapsulated at the UZ using ETP Copper

Choice of using LG and test different PMTs



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## ANAIS-ELECTRONIC CHAIN

#### ALMOST FULLY COMMISSIONED

BUS VME/NIM LINUX-ROOT



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		PMT Testing		
HP Ge spe at Canfrar	ectrometry	Ham LB	Ham ULB	Ham VLB
	Model	<sup>40</sup> K (mBq/PMT)	<sup>232T</sup> h (mBq/PMT)	<sup>2328</sup> U (mBq/PMT)
- <b>P</b> -	HAM - R6233-100 HAM-LB	678 ± 42	68 ± 3	100 ± 3
	HAM - R11065SEL	12 ± 7	3.6 ± 1.2	<sup>238</sup> U: 47 ± 28
	HAM-ULB			<sup>226</sup> Ra: 8.0 ± 1.2
	HAM - R6956MOD	97± 18	20 ± 2	<sup>238</sup> U: 128 ± 38
The second secon	HAM-VLB			<sup>226</sup> Ra: 84 ± 3
Participantes .	M.L. Sarsa	, IDM 2012		13

## **PMT** Testing

#### We would like to avoid LG ...

# Direct measurement in ANAIS-0 prototype



Ham ULB Ham VLB



#### **PMT** Testing

Ham VLB

We would like to avoid LG ...

We need FILTER events below 20 keV (no electronic noise)

Rely on SIMULATIONS



Ham ULB



We want to select real bulk Nal scintillation events

CUTS have been developed for:

 Periods of high rate (f.i. periods after calibrations or electronically noisy)
 Events coincident with a muon on plastic vetoes.
 Events after a muon (Very High Energy event) in the Nal crystal in a 0.5 s-window.

4. Events having in total no more than 4 photoelectrons (n0s≤4 p.e.)

5. Anomalous fast events

#### Muon related events in ANAIS-0

We have studied events coincident with a VETO SIGNAL in the plastic scintillators (partially covering the shielding)



#### Muon related events in ANAIS-0

We have studied events coincident with a VETO SIGNAL in the plastic scintillators (partially covering the shielding)



Events at low energy coincident with a veto signal are present in our data and scintillation in the LG is confirmed

#### Muon related events in the crystal

High energy depositions excite very slow scintillation component in NaI(TI), producing a strong increase in rate after each muon event (very low energy events):



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# Determination of the very slow scintillation time constants

We have determined in a precise way the time constants of the Nal(TI) phosphorescence

M.L. Sarsa, IDM 2012

Very difficult to measure at sea level



We count the total number of photoelectrons in every low energy event



We count the total number of photoelectrons in every low energy event



#### We filter the fast anomalous event population





#### We filter the fast anomalous event population



M.L. Sarsa, IDM 2012

We filter the fast anomalous event population



We have built a background model for the ANAIS-0 module, based on radiopurity input data for the different detector and shielding components + Géant 4 code

NaI crystal bulk activity		
mBq/kg		
$12.7\pm0.5$		
$0.96 \pm 0.06$		
$0.075 \pm 0.005$		
$0.023\pm0.007$		
$0.098 \pm 0.004$		
$0.188\pm0.005$		
$0.013 \pm 0.005$		
$0.035\pm0.003$		



Background contribution from bulk crystal contaminations

Simulated component	Isotope	e Activity	
	<sup>40</sup> K	< 11 mBq	
	<sup>232</sup> Th	< 4.1  mBq	
Copper encapsulation	<sup>238</sup> U	< 140  mBq	
	$^{226}$ Ra	< 2  mBq	
	<sup>60</sup> Co	< 0.94  mBq	
	<sup>40</sup> K	< 12  mBq/kg	
	$^{232}$ Th	< 2.2  mBq/kg	
Quartz optical window	<sup>238</sup> U	< 100  mBq/kg	
	$^{226}$ Ra	< 1.9  mBq/kg	
Tillerile	<sup>40</sup> K	< 21  mBq/guide	
	<sup>232</sup> Th	< 4.1  mBq/guide	
Light guides	<sup>238</sup> U	< 120  mBq/guide	
	$^{226}$ Ra	< 4.7 mBq/guide	
	<sup>40</sup> K	< 200  mBq/kg	
Optical coupling grease	<sup>232</sup> Th	< 200  mBq/kg	
	<sup>238</sup> U	$< 2000 \mathrm{~mBq/kg}$	
	$^{226}$ Ra	< 30  mBq/kg	
	<sup>210</sup> Pb	< 20  mBq/kg	
Archaeological Lead	$^{232}$ Th	< 0.3  mBq/kg	
	<sup>238</sup> U	< 0.2  mBq/kg	
Inner volume air	$^{222}$ Rn	$< 0.6 \ \mathrm{Bq/m^3}$	
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nd model for the ANAIS-0 module, based on data for the different detector and shielding



Background contribution from other detector components, shielding and

components

We have built a background model for the ANAIS-0 module, based on radiopurity input data for the different detector and shielding components



**TO APPEAR IN Astroparticle Physics** 

M.L. Sarsa, IDM 2012

Introducing some small modifications: reducing some contaminations below the upper limits, adding some surface components, ...



Simulated	Background contribution			
component	150 - 3000  keV			
	counts/kg/day	%		
Crystal	1000.2 (1002.7)	77.5 (77.7)		
Copper encapsulation	67.8 (47.5)	5.3 (3.7)		
Quartz window	13.5(6.6)	1.0(0.5)		
Optical coupling grease	21.4 (7.2)	1.7(0.6)		
Ham ULB PMT	47.2	3.7		
Lead shielding	36.0 (145.0)	2.8 (11.2)		
Inner volume air	9.1	0.7		
Total	1195.2 (1265.3)	92.6 (98.1)		

#### Table 6

Contribution to the background of ANAIS-0 module in set-up 4 from the different simulated set-up components: total rate derived from simulation from 150 keV to 3000 keV and corresponding percentage of the measured background. Rates and percentages shown in parentheses correspond to the simulation including extra hypotheses (see text).

#### **TO APPEAR IN Astroparticle Physics**

Simulated	Background contribution				
component	5 - 150 keV		2 - 20  keV		
	counts/kg/day	%	counts/kg/day	%	
Crystal	196.0 (255.0)	53.0 (68.8)	49.6 (65.0)	64.8 (84.8)	
Copper encapsulation	46.1 (31.4)	12.5(8.5)	1.4 (8.0)	1.8(10.5)	
Quartz window	15.5(2.5)	4.2 (0.7)	1.8 (0.2)	2.4(0.3)	
Optical coupling grease	15.7 (3.3)	4.2 (0.9)	5.8 (0.9)	7.6 (1.2)	
Ham ULB PMT	15.4	4.2	0.5	0.7	
Lead shielding	3.2 (33.7)	0.9 (9.1)	0.3 (1.3)	0.3(1.8)	
Inner volume air	1.2	0.3	0.1	0.1	
Total	293.2 (342.1)	79.3 (92.5)	59.5 (76.0)	77.7 (99.3)	

#### Table 7

Contribution to the background of ANAIS-0 module in set-up 4 from the different simulated set-up components: total rate derived from simulation from 5 keV to 150 keV and corresponding percentage of the measured background, total rate derived from simulation from 2 keV to 20 keV and corresponding percentage of the measured background. Rates and percentages shown in parentheses correspond to the simulation including extra hypotheses (see text).

#### **TO APPEAR IN Astroparticle Physics**

# Background model for the new crystals

Applying the same background model, assuming the only improvement is the <sup>40</sup>K reduction down to 20 ppb level:



And without profit from coincidence between modules rejection factor ...

#### **Conclusions and Prospects**

- Radiopure Nal powder (<90 ppb potassium) has been found and two 12.5 kg prototypes are almost ready to final background asessment at LSC
- If potassium content is below 20 ppb, 250 kg Nal (20 x 12,5 kg) will be mounted at LSC along 2013
- Electronic chain and readout is almost ready and the lead and polyethylene for the shielding are at LSC waiting to be mounted
- We understand quite well our present backgrounds: simulations and filtering protocols seem to work well