

SuperCDMS-SNOLAB: an active neutron veto shield design

IDM 2012 Chicago
July, 23 2012

Silvia Scorza
Southern Methodist University
SuperCDMS collaboration

SuperCDMS

Dark Matter Search

Goal: direct detection of WIMP elastic scattering off nuclei few WIMPS/year/ton

Signature: nuclear recoil with $E < 100$ KeV

Shielding (Pb, polyethylene, Cu)

Reduce backgrounds from radioactivity

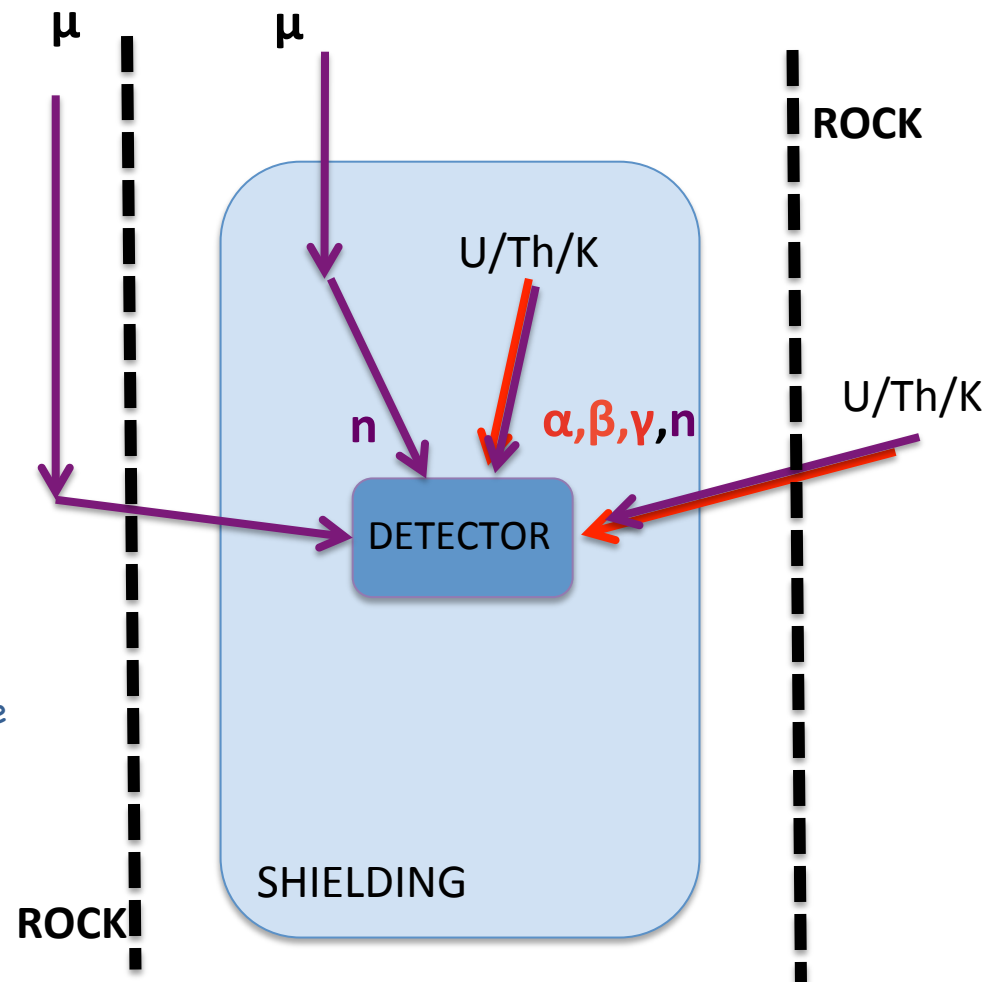
Active Background Rejection

Distinguish between nuclear recoils (WIMPS, neutrons) and electron recoils (backgrounds) \rightarrow by recording both ionization and heat (phonon) signals

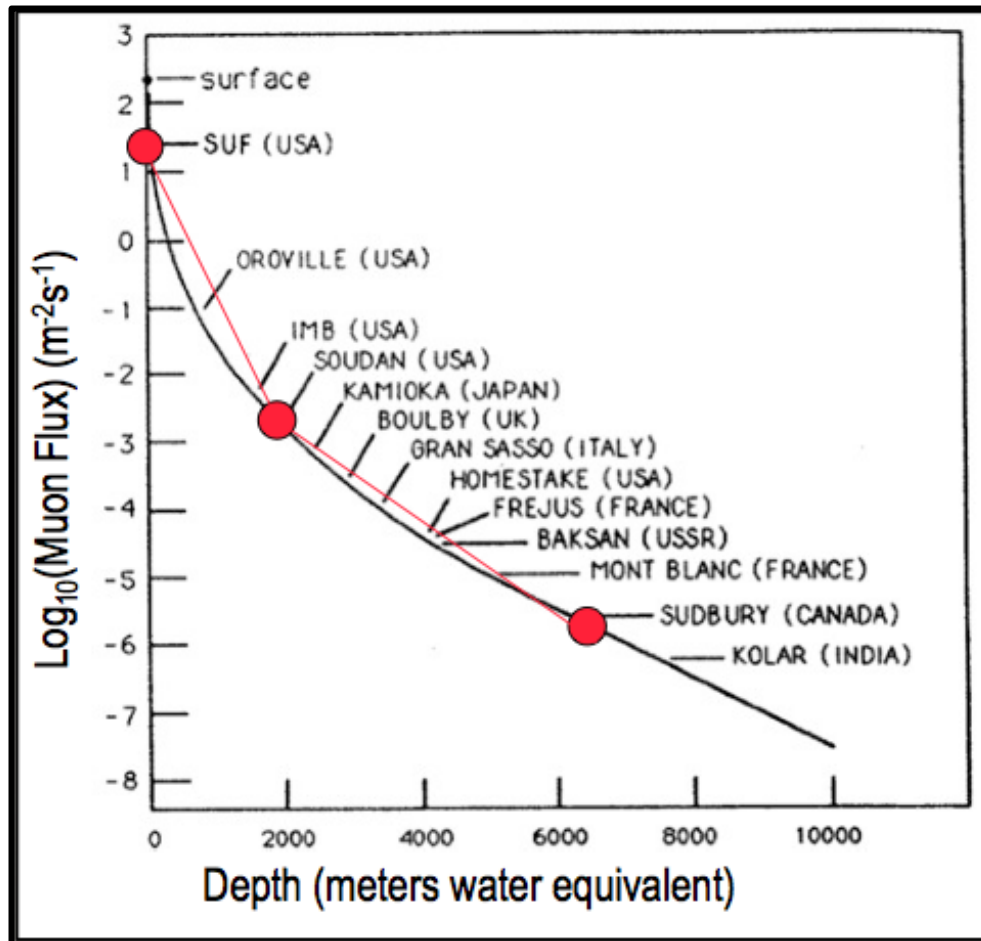
Surface events tag \rightarrow interdigitized electrode scheme – phonon sensor

Deep Underground (SNOLAB)

Fewer cosmic rays to produce neutrons
Neutrons produce nuclear recoils



Depth is Important !



SUF

17 mwe

0.5 n/d/kg
(182.5 n/y/kg)

Soudan

2090 mwe

0.05 n/y/kg

SNOLab

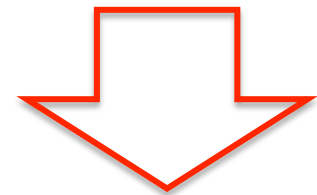
6060 mwe

0.2 n/y/ton
(0.0002 n/y/kg)

Moving from
Soudan to
SNOLab:

Reduce muon
flux by **500x**

Reduce high-
energy neutron
flux by **100x**



Worry about
neutrons from
**residual
radioactivity
only**

Which Neutrons?

- **Cosmogenic** muon-induced: expect <0.1 in 100kg- years @SNOLab
- **External radiogenics** (Fission and (α, n) from U,Th in cavern rock): expected to be negligible with passive shielding
- **Internal radiogenics** (Fission and (α, n) from U,Th in Cu cans and supports): expect order of 1 in 100 kg - depending on screening and material cleanliness



**Active
Neutron
Veto**

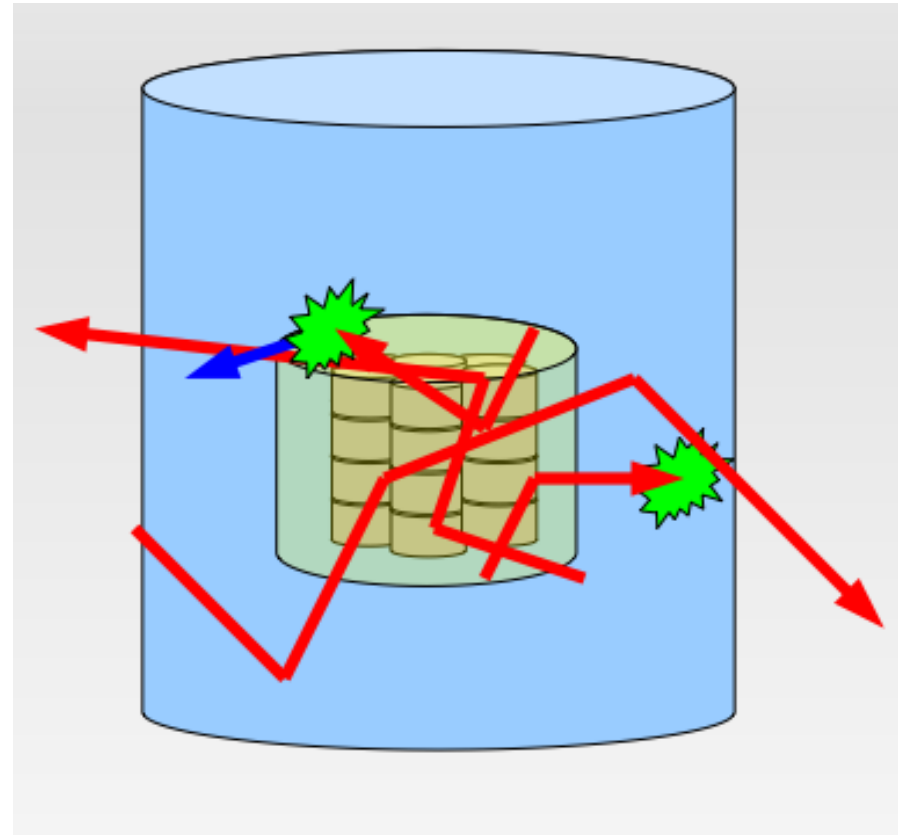
- Identification background events especially neutrons that can produce nuclear recoils similar to WIMPS.
- The veto will indirectly act as a diagnostic device.

Neutron veto - how to

- n capture in the veto
- fast scatters in the veto
- gamma from captures outside the veto

WIMP search strategy

Any WIMP candidate in coincidence with a veto energy deposited of the n-capture process will be rejected



Physics Requirements

- Total unvetoes background in Ge $\ll 1$ counts in 100kg SNOlab phase
- Total background rate (neutron & gamma) must be the same: neutron veto must not generate excessive backgrounds in the zips
 - > Implies radioclean construction
- Negligible contribution to dead-time
 - > Implies low ($< \text{kHz}$) non-coincident trigger rate
- High ($\sim 90\%$ or better) efficiency
(a modest efficiency of $\sim 80\%$ would reduce neutron background to < 1 event)

Neutron Flux Monitor

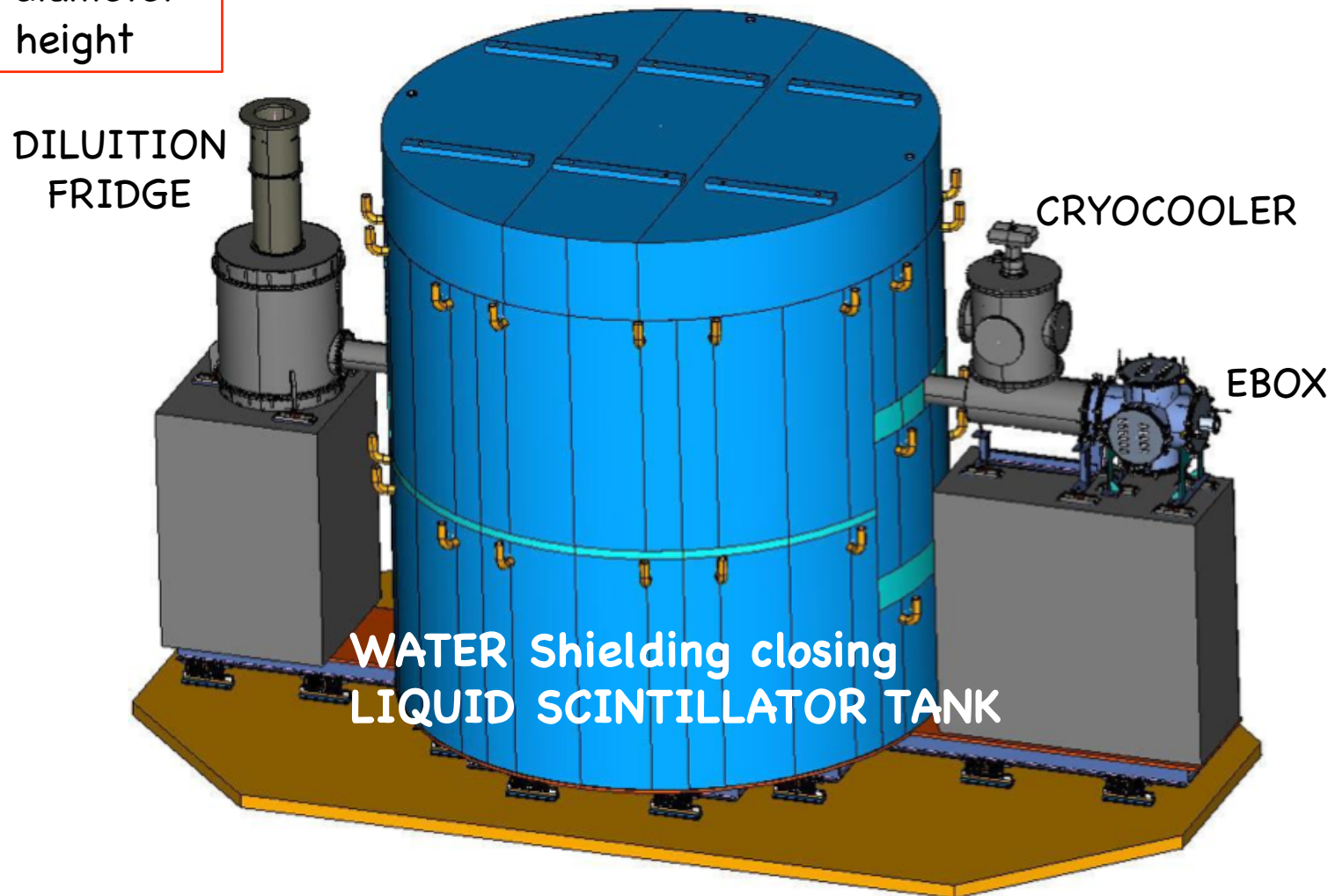
Additional requirement:

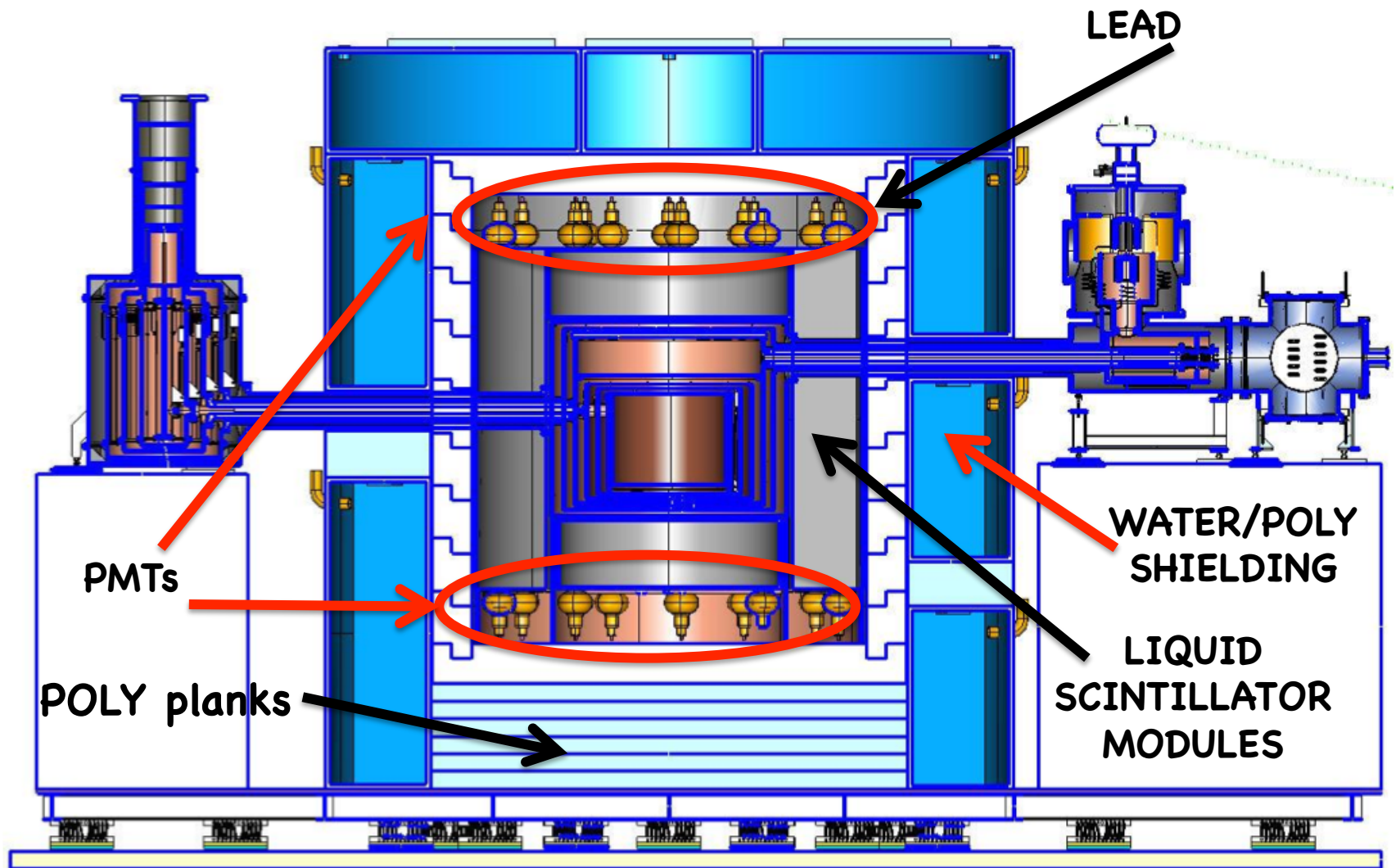
good ability to discriminate neutrons from the gamma background

- *In situ* measurement of radiogenic neutron rates – better precision than the multiple scattering measurement
- Evaluation of Monte Carlo systematics
- Monitoring/tuning of Ge nuclear recoil acceptance using tagged neutrons

Current Design

Dimensions:
13 feet diameter
14 feet height

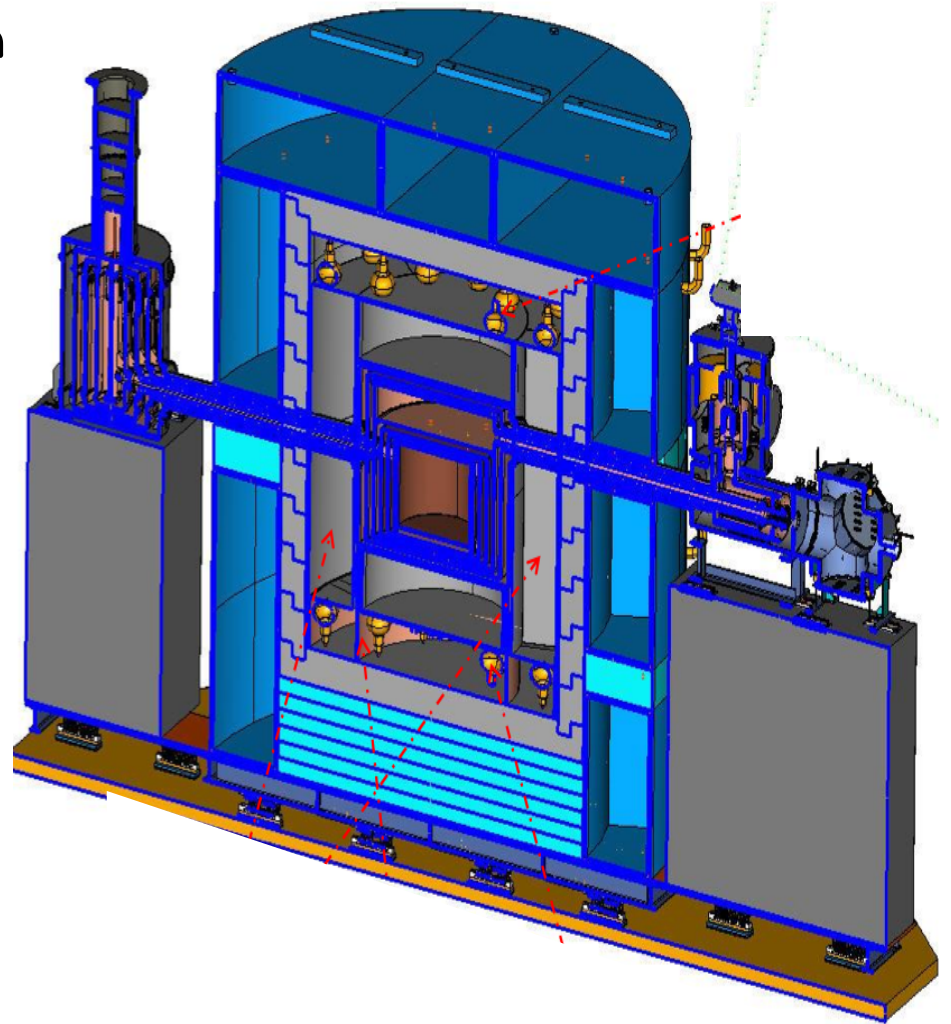




Neutron veto modules filled with Linear Alkylbenzene (LAB), read out by PMTs

Doped with high cross-section isotopes (B, Gd, Li)

- decreases capture time/distance (ex. 5% Boron doping reduces capture time from 250 μ s to 3 μ s)
- affects design due to need to contain+detect capture products



LAB Doping

Boron

Gadolinium

-> **ALPHA** (~3 MeV) + **GAMMA** (500keV)
-> observed light may be as low as 50keVee

Challenges:

- ✓ minimize environmental radioactivity by
 - constructing the detector out of radiopure materials,
 - developing a clean boron-loaded scintillator,
 - utilizing adequate shielding for the neutron veto.
- ✓ energy threshold

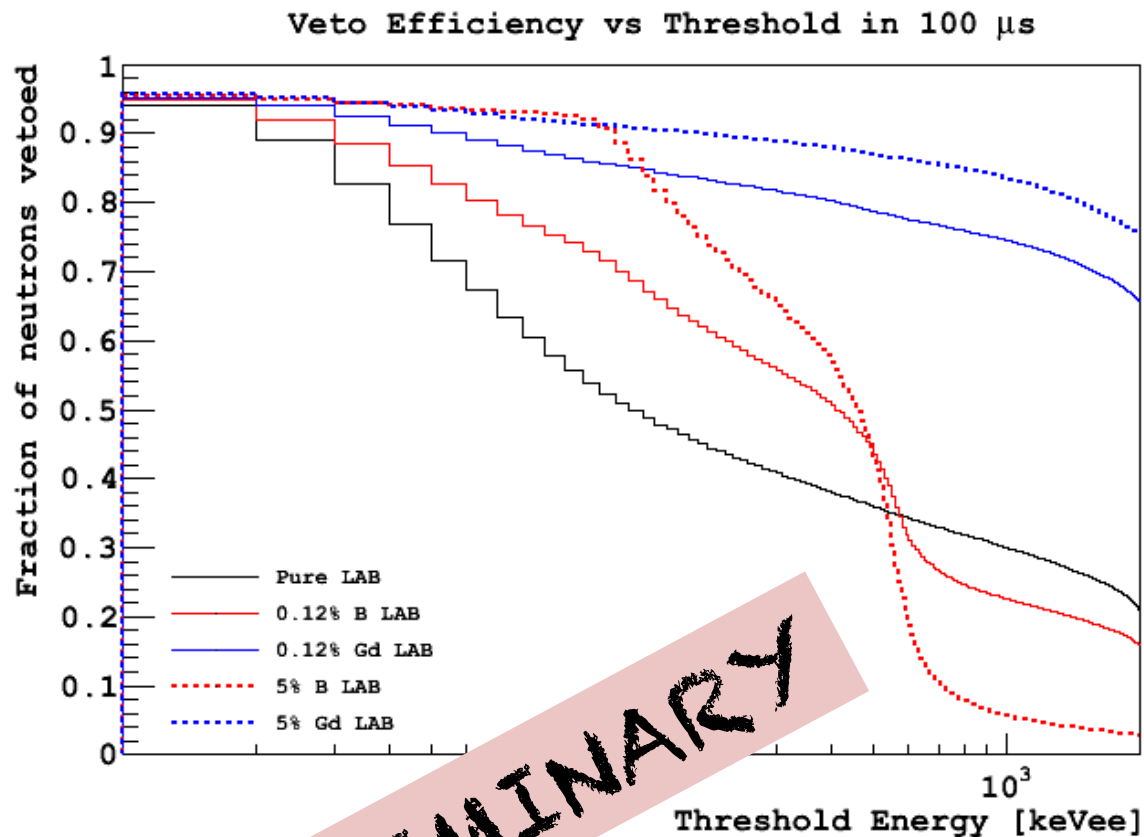
-> **GAMMA** cascade 8MeV
(> ^{208}Tl line ~2.7MeV)

- Reduction outer shielding
- It has been demonstrated by Daya Bay experiment

BUT

- decreased efficiency for detecting internal neutrons
- possible introduction of radio contaminant (Gd is less pure than B)

Efficiency



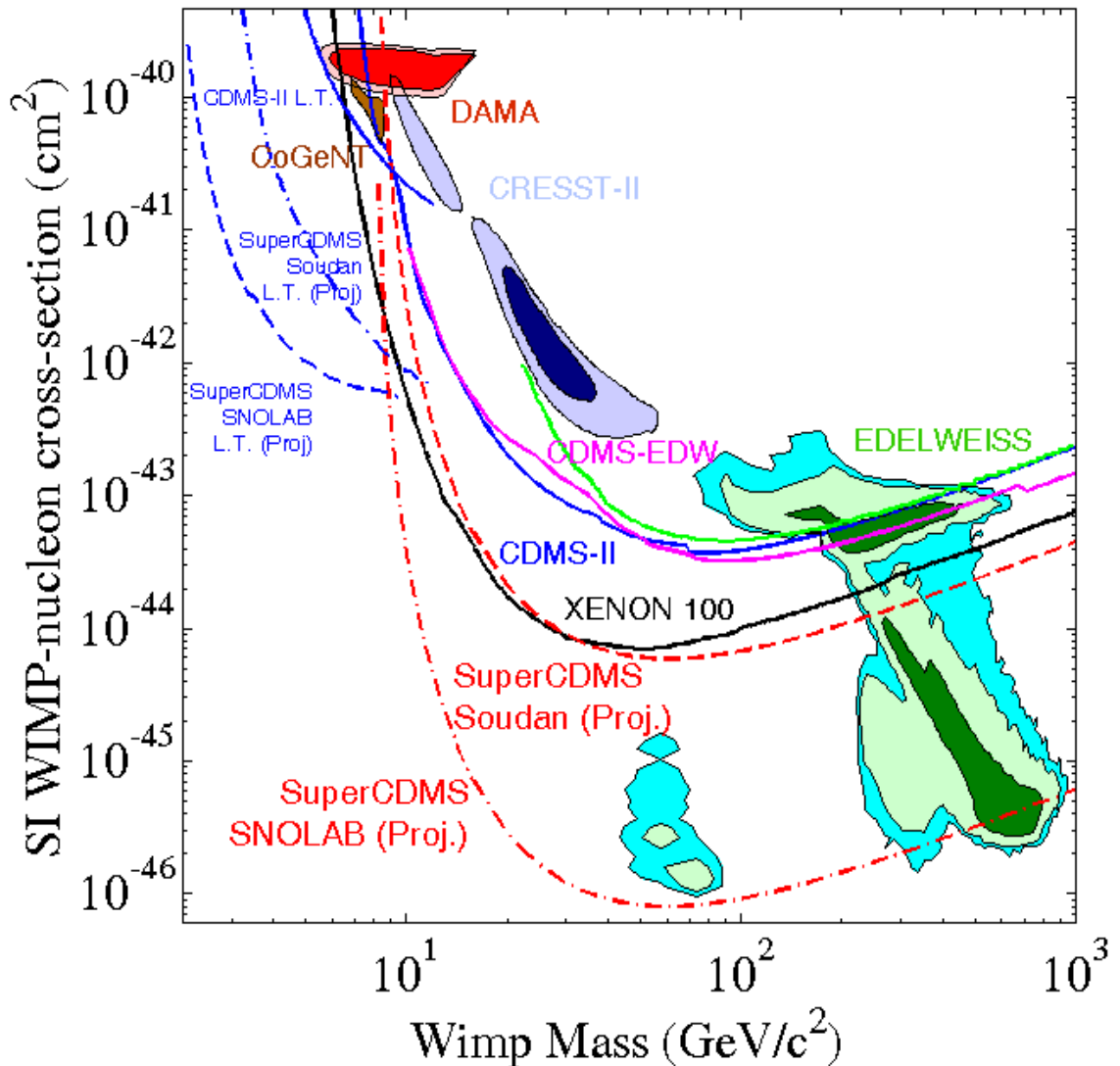
Veto efficiency vs threshold for 100 μ s veto times for recoil events.

Recoil events: passing the energy deposition criteria (10 -100keV) and >10% of the deposited energy must have come via recoils

LAB Gd doped shows higher efficiency than LAB B doped

Sensitivity and Timeline

- Scintillator and optical test starting soon
- Final design by 2014
- SuperCDMS SNOLab construction 2014



CDMS/SuperCDMS Collaborations



Caltech

Instituto de Fisica Teorica, Universidad Autonoma de Madrid Fermilab

MIT

NIST

Queens University

Santa Clara University

SLAC/KIPAC

Southern Methodist University

Stanford University

Syracuse University

University of British Columbia

University of California , Berkeley

University of California , Santa Barbara

University of Colorado, Denver

University of Florida

University of Minnesota

University of Texas A&M

THANKS