

Possible UV Completion through Classicalization Mechanism

Luke Keltner

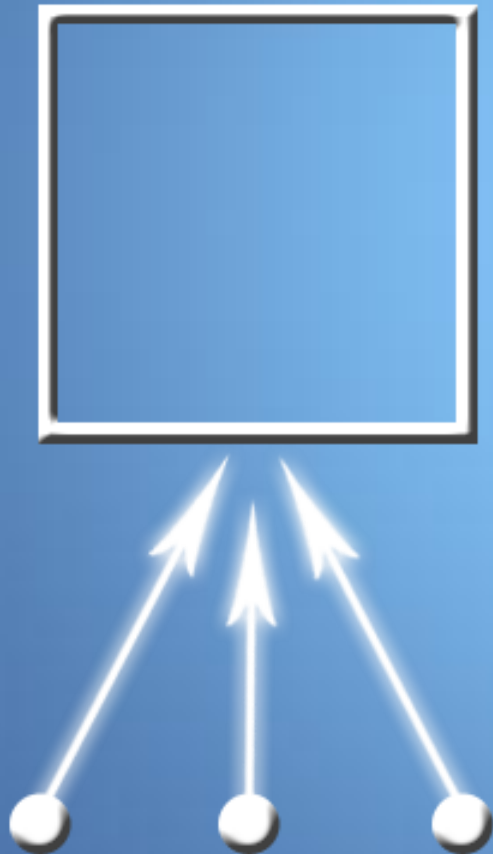


Andrew Tolley



Some things
come in

$$a_{k_1}^\dagger a_{k_2}^\dagger \dots |0\rangle$$

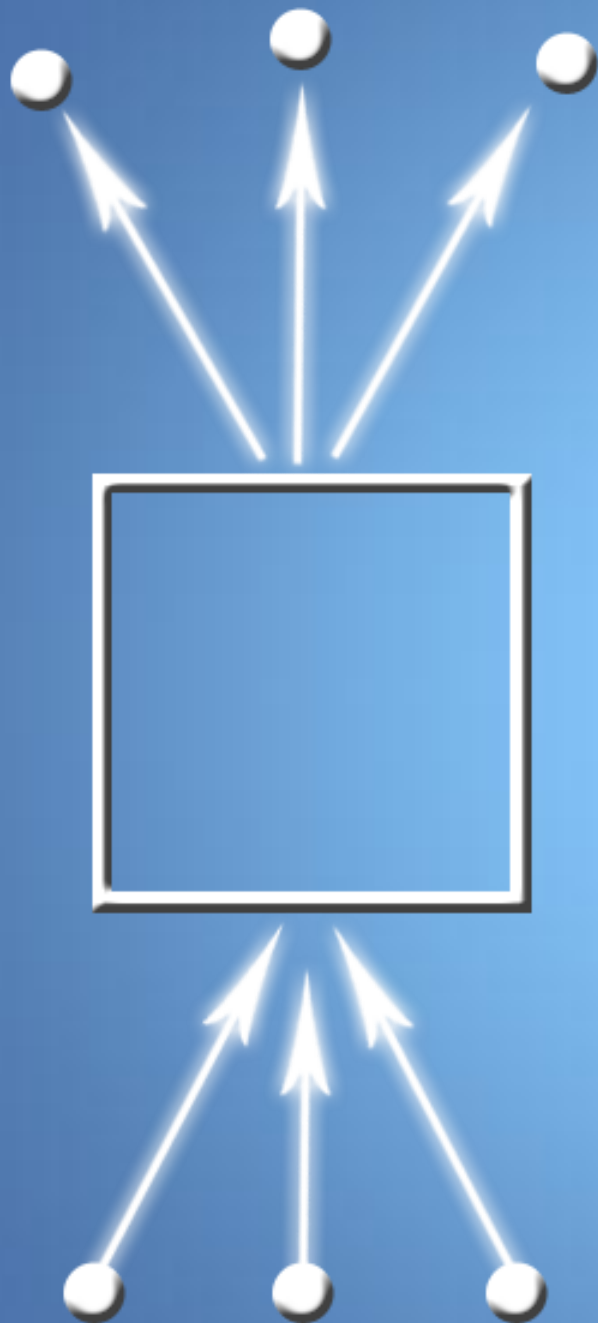


Some things
happen

$$T^* e^{i \int d^4 x \mathcal{L}_{int}}$$

Some things
come in

$$a_{k_1}^\dagger a_{k_2}^\dagger \dots |0\rangle$$



Some things
come out

$$a_{k_3}^\dagger a_{k_4}^\dagger \dots |0\rangle$$

Some things
happen

$$T^* e^{i \int d^4 x \mathcal{L}_{int}}$$

Some things
come in

$$a_{k_1}^\dagger a_{k_2}^\dagger \dots |0\rangle$$

\hat{S} Matrix

$$\langle 0 | a_{k_4} a_{k_3} \hat{S} a_{k_1}^\dagger a_{k_2}^\dagger | 0 \rangle = \langle 0 | a_{k_4} a_{k_3} T^* e^{i \int \mathcal{L}_I(x) d^4x} a_{k_1}^\dagger a_{k_2}^\dagger | 0 \rangle$$

$$= \langle 0 | a_{k_4} a_{k_3} \left[1 + i \int d^4x \mathcal{L}_I(x) + \frac{(i)^2}{2} \int \int d^4x d^4y T^* \mathcal{L}_I(x) \mathcal{L}_I(y) + \dots \right] a_{k_1}^\dagger a_{k_2}^\dagger | 0 \rangle$$

The \hat{S} matrix is often used for scattering and is usually treated perturbatively.

The Problem

$$S = \int d^4x \quad -\frac{1}{2}(\partial\phi)^2 - \frac{1}{4\Lambda^4}(\partial\phi)^4$$

The Problem

Treat perturbatively

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The Problem

Treat perturbatively

$$S = \int d^4x \quad -\frac{1}{2}(\partial\phi)^2 - \frac{1}{4\Lambda^4}(\partial\phi)^4$$



$$\langle 0 | a_{k_4} a_{k_3} \hat{S} a_{k_1}^\dagger a_{k_2}^\dagger | 0 \rangle \sim \frac{s^2}{\Lambda^4} \quad s \sim E_{\text{CM}}^2$$

Tree level calculation breaks unitarity
once $s^2 \sim \Lambda^4$.

Wilsonian Completion

Wilsonian Completion

$$S[l, H] = \int d^4x \quad -\frac{1}{2}(\partial l)^2 - \frac{1}{2}(\partial H)^2 - V(l, H)$$

$$V(l, H) = \frac{1}{2}m^2 l^2 + \frac{1}{2}M^2 H^2 + \frac{1}{4!}g_l l^4 + \frac{1}{4!}g_H H^4 + \frac{1}{4}g_{lH} l^2 H^2 + \frac{1}{2}\tilde{m} l^2 H + \frac{1}{3!}\tilde{g}_H M H^3$$

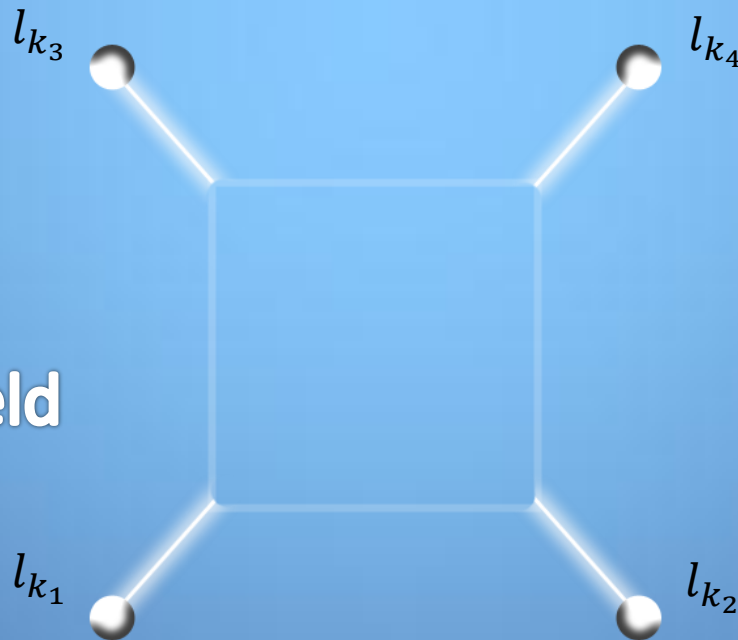
Wilsonian Completion

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$l \rightarrow$ light field

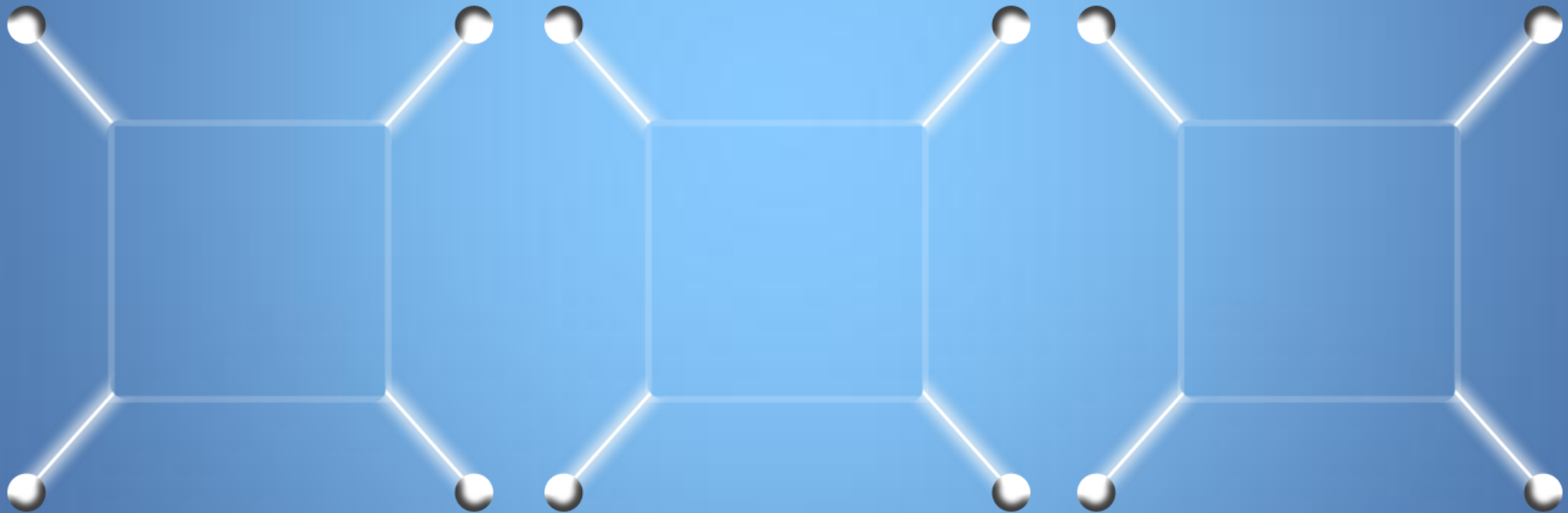
$H \rightarrow$ heavy field



$$\langle 2 | \hat{S} | 2 \rangle$$

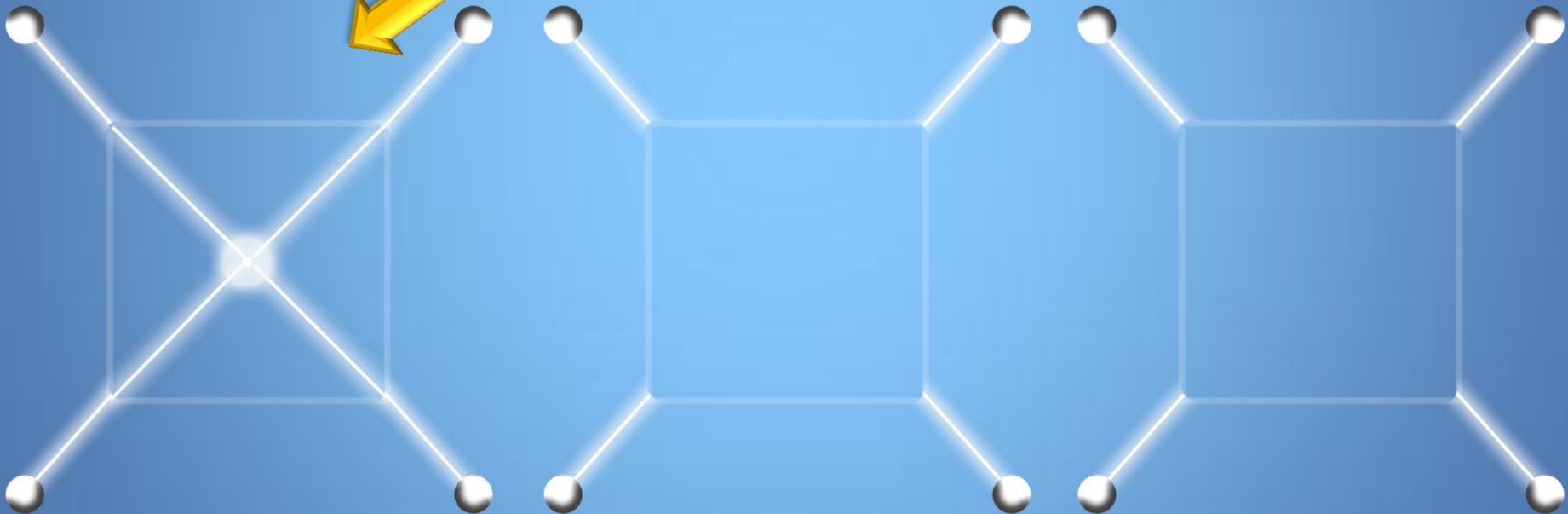
Wilsonian Completion

$$V(l, H) = \frac{1}{2}m^2 l^2 + \frac{1}{2}M^2 H^2 + \frac{1}{4!}g_l l^4 + \frac{1}{4!}g_H H^4 + \frac{1}{4}g_{lH} l^2 H^2 + \frac{1}{2}\tilde{m}l^2 H + \frac{1}{3!}\tilde{g}_H M H^3$$



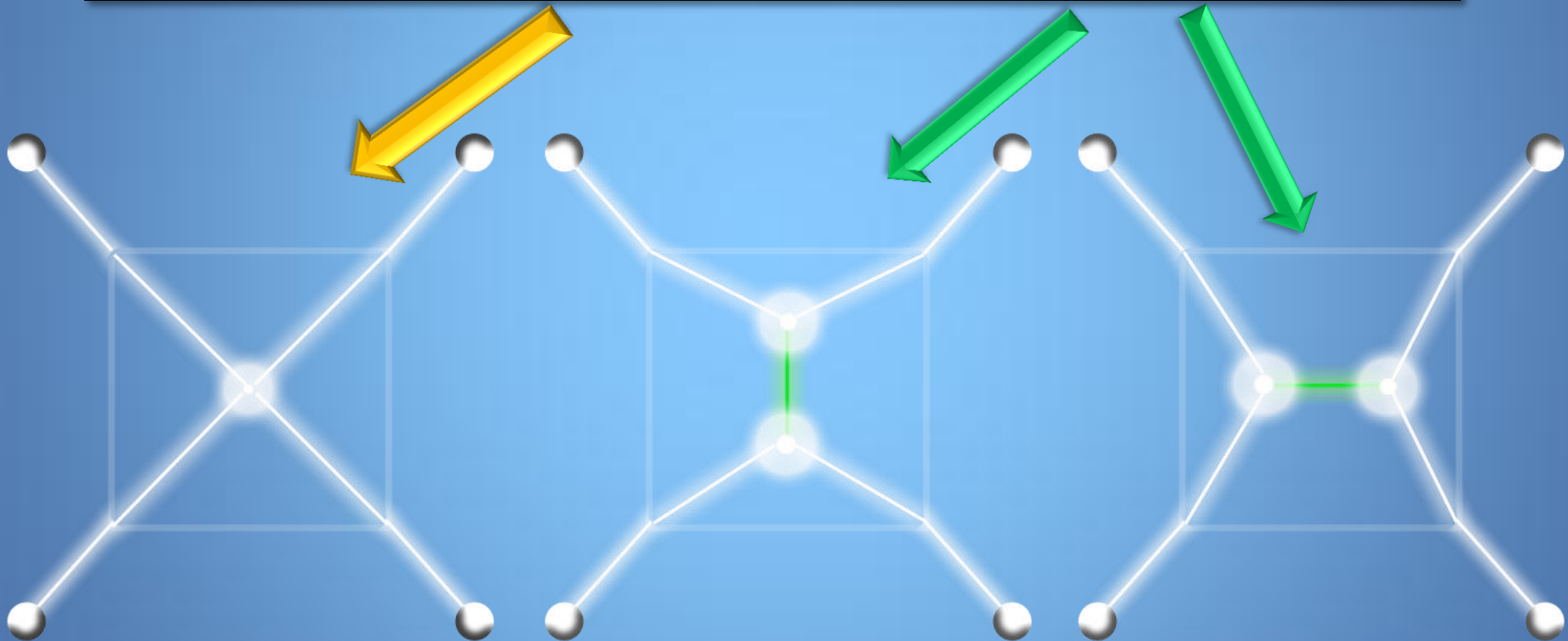
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$$\langle 2 | \hat{S} | 2 \rangle$$

$$\langle 0 | a_{k_4} a_{k_3} \hat{S} a_{k_1}^\dagger a_{k_2}^\dagger | 0 \rangle = i(2\pi)^4 \delta^{(4)}(k_1 + k_2 - k_3 - k_4) \mathcal{A}(k_1, k_2, k_3, k_4)$$

$$\langle 2 | \hat{S} | 2 \rangle$$

$$\langle 0 | a_{k_4} a_{k_3} \hat{S} a_{k_1}^\dagger a_{k_2}^\dagger | 0 \rangle = i(2\pi)^4 \delta^{(4)}(k_1 + k_2 - k_3 - k_4) \mathcal{A}(k_1, k_2, k_3, k_4)$$

$$\mathcal{A}(k_1, k_2, k_3, k_4) = -g_l + \tilde{m}^2 \left[\frac{1}{s + M^2} + \frac{1}{t + M^2} + \frac{1}{u + M^2} \right]$$

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$$\mathcal{A}(k_1, k_2, k_3, k_4) \sim -g_I + \frac{3\tilde{m}^2}{M^2} + \frac{4\tilde{m}^2 m^2}{M^4} + \frac{\tilde{m}^2 s^2}{M^6} + \dots$$

$$\langle 2 | \hat{S} | 2 \rangle$$

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Breaks unitarity when $s^2 \sim M^4$

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Breaks unitarity when $s^2 \sim M^4$

$$S[l, H]$$



$$S_{\text{eff,NR}}[l]$$

Energy



Below energy scale M ,
 $S_{\text{eff,NR}}[l]$ is *effectively*
correct.

Renormalizable

$$S[l, H]$$



M

Non-renormalizable

$$S_{\text{eff,NR}}[l]$$



Solution

Solution

Wilsonian Completion

Solution

Wilsonian Completion

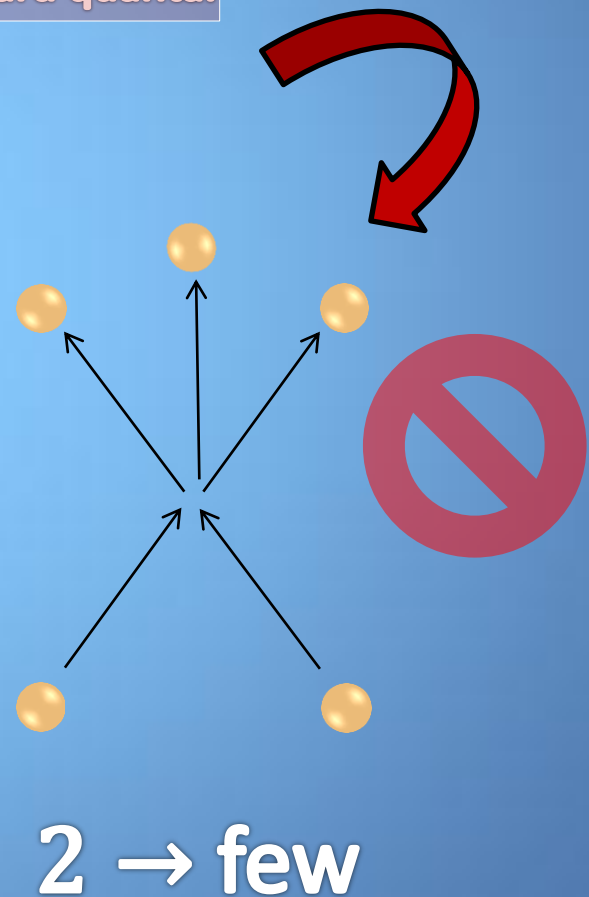
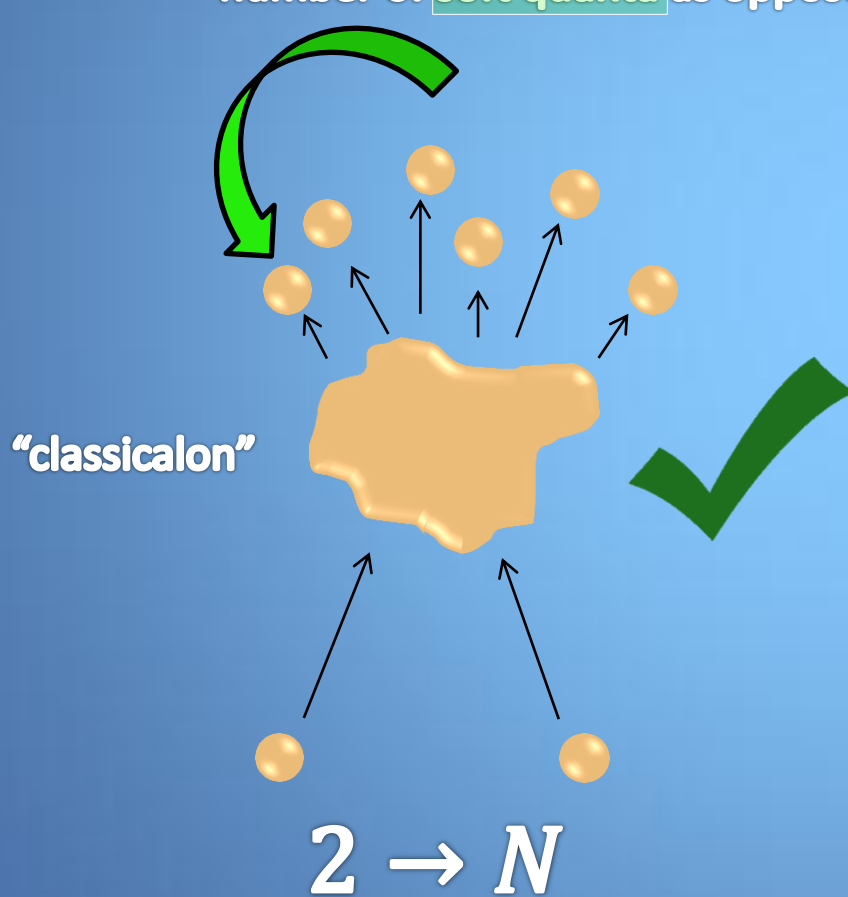


Hypothesis (Dvali et al.): [arXiv:1010.1415v2 \[hep-ph\]](https://arxiv.org/abs/1010.1415v2)

Classicalization

Classicalization

At large energies, scattering prefers processes involving a large number of **soft quanta** as oppose to a few **hard quanta**.



Why Classical?

Why Classical?

$$S = \int d^4x \quad -\frac{1}{2}(\partial\phi)^2 - \frac{1}{4\Lambda^4}(\partial\phi)^4$$

Goldstone

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$$S = \int d^4x \quad -\frac{1}{2}(\partial\phi)^2 - \frac{1}{4\Lambda^4}(\partial\phi)^4$$

Goldstone

$$[S] = \hbar$$

$$[\phi] = \frac{\sqrt{\hbar}}{L}$$

$$[\Lambda^4] = \frac{\hbar}{L^4}$$

$$[E] = \frac{\hbar}{L}$$

Why Classical?

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Goldstone

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$$[E] = \frac{\hbar}{L}$$

$\lim_{\hbar \rightarrow 0}$



Classical length
increases with
energy

$$r_* \sim \frac{1}{\Lambda} \left(\frac{E}{\Lambda} \right)^{\frac{1}{3}}$$

Why Classical?

$$S = \int d^4x \left[-\frac{1}{2}(\partial\phi)^2 - \frac{1}{\Lambda^3} \square\phi(\partial\phi)^2 \right]$$

Galileon

Why Classical?

$$S = \int d^4x \quad -\frac{1}{2}(\partial\phi)^2 - \frac{1}{\Lambda^3}\Box\phi(\partial\phi)^2$$

Galileon

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$$[\Lambda^3] = \frac{\sqrt{\hbar}}{L^3}$$

$$[E] = \frac{\hbar}{L}$$

$\lim_{\hbar \rightarrow 0}$



$$r_* \sim \frac{1}{\Lambda} \left(\frac{E}{\Lambda} \right)^{\frac{1}{5}}$$

Why Classical?

$$S = \int d^4x \quad -\frac{1}{2}(\partial\phi)^2 - \frac{1}{\Lambda^6}(\Box\phi)^2(\partial\phi)^2$$

(another)
Galileon

Why Classical?

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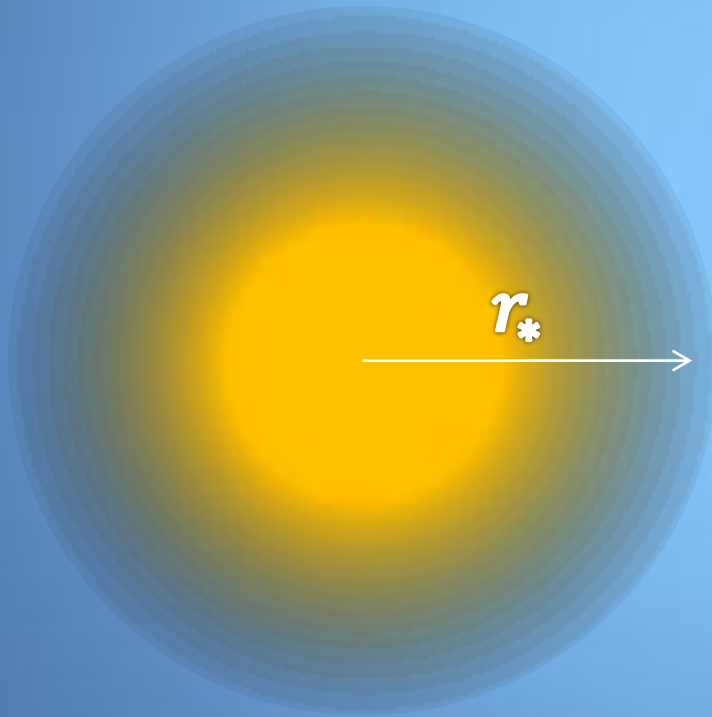
(another)
Galileon

$$r_* \sim \frac{1}{\Lambda} \left(\frac{E}{\Lambda} \right)^{\frac{1}{5}}$$

What is r_* ?

What is r_* ?

Goldstone



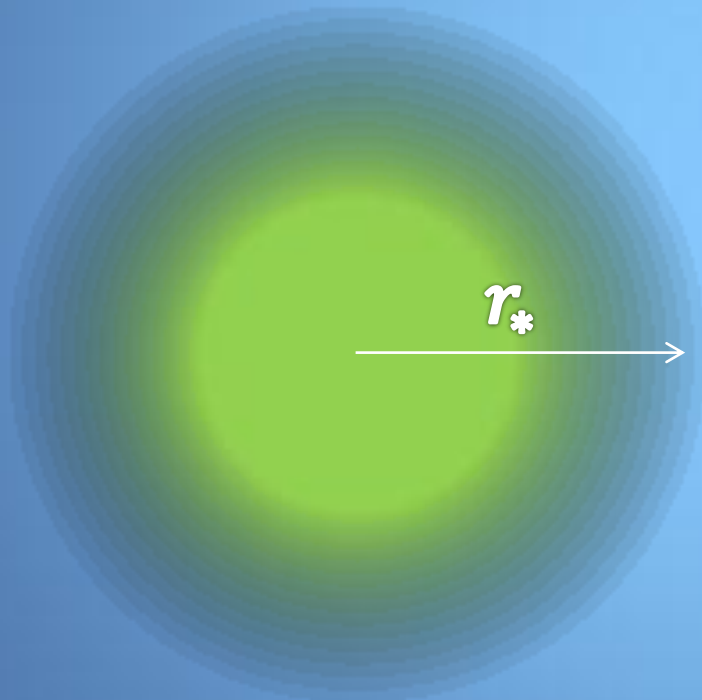
$$E \sim \int_{r < r_*} dr \, r^2 (\partial_r \phi)^2 \sim r_* \phi^2$$

$$(\partial \phi)^2 \sim \frac{1}{\Lambda^4} (\partial \phi)^4$$

$$r_* \sim \frac{1}{\Lambda} \left(\frac{E}{\Lambda} \right)^{\frac{1}{3}}$$

What is r_* ?

Galileon



$$E \sim \int_{r < r_*} dr \, r^2 (\partial_r \phi)^2 \sim r_* \phi^2$$

$$(\partial \phi)^2 \sim \frac{1}{\Lambda^3} \square \phi (\partial \phi)^2$$

$$r_* \sim \frac{1}{\Lambda} \left(\frac{E}{\Lambda} \right)^{\frac{1}{5}}$$

Region I

S_{weak}

Perturbation theory **works**

Region II

S_{strong}

Perturbation theory
doesn't work

r_*



Self Energy

- **Divergent** classical and quantum **self energies** arise with the notion of a point particle.
- Regularization and renormalization help to **artificially** make the quantum self energy finite.

Self Energy

Goldstone

Region I

$$S_I = \int d^4x \quad -\frac{1}{2}(\partial\phi)^2 - \phi J$$

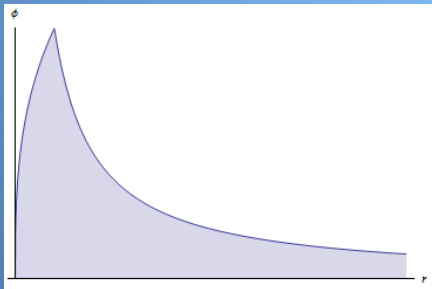
$$\phi_I \sim \frac{M}{r}$$

Region II

$$S_{II} = \int d^4x \quad -\frac{1}{\Lambda^4}(\partial\phi)^4 - \phi J$$

$$\phi_{II} \sim (M\Lambda^4)^{\frac{1}{3}} r^{\frac{1}{3}}$$

$$E_{tot} \sim \frac{M^2}{r_*} + (M\Lambda)^{(4/3)} r_*^{1/3}$$



Finite self energy

Self Energy

Galileon

Region I

$$S_I = \int d^4x \quad -\frac{1}{2}(\partial\phi)^2 - \phi J$$

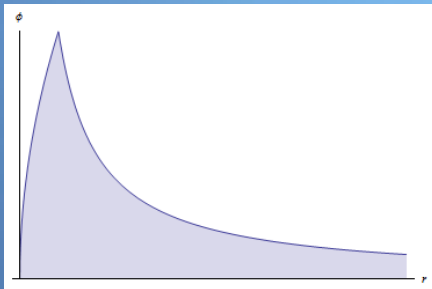
$$\phi_I \sim \frac{M}{r}$$

Region II

$$S_{II} = \int d^4x \quad -\frac{1}{\Lambda^3} \square\phi(\partial\phi)^2 - \phi J$$

$$\phi_{II} \sim (M\Lambda^3)^{\frac{1}{2}} r^{\frac{1}{2}}$$

$$E_{tot} \sim \frac{M^2}{r_*} + (M\Lambda)^{(3/2)} r_*^{1/2}$$



Finite self energy

What is a “Classicalon”?

What is a “Classicalon”?

Claim:

Coherent State

$$|\alpha\rangle \sim e^{\alpha a^\dagger} |0\rangle$$

Black Holes: coherent state of gravitons

$$\Delta x_\alpha \Delta p_\alpha = \frac{\hbar}{2}$$

Electric Field: coherent state of photons

Semiclassical Calculation

Lasma, Bezrukov arXiv:1206.5311v1 [hep-th]

$$\sigma(E, N) = \sum_f |\langle f | P_E P_N \hat{S} \hat{A} | 0 \rangle|^2$$

Semiclassical Calculation

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$$\sigma(E, N) = \sum_f |\langle f | P_E P_N \hat{S} \hat{A} | 0 \rangle|^2$$

$$\sigma(E, N) = \int db_{\mathbf{k}}^* db_{\mathbf{k}} d\xi d\eta \mathcal{D}\phi \mathcal{D}\phi' \exp \left(- \int dk b_{\mathbf{k}}^* b_{\mathbf{k}} e^{i\omega_{\mathbf{k}}\xi + i\eta} + iE\xi + iN\eta + B_i(0, \phi_i) + B_f(b_{\mathbf{k}}^*, \phi_f) \right. \\ \left. + B_i^*(0, \phi'_i) + B_f^*(b_{\mathbf{k}}, \phi'_f) + iS[\phi] - iS[\phi'] + J\phi(0) + J\phi(0)' \right)$$

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$$B_i(0, \phi) = -\frac{1}{2} \int d\mathbf{k} \omega_{\mathbf{k}} \phi_i(\mathbf{k}) \phi_i(-\mathbf{k})$$

$$B_f(b_{\mathbf{k}}^*, \phi_f) = -\frac{1}{2} \int d\mathbf{k} b_{\mathbf{k}}^* b_{-\mathbf{k}}^* e^{2i\omega_{\mathbf{k}}T_f} + \int d\mathbf{k} \sqrt{2\omega_{\mathbf{k}}} b_{\mathbf{k}}^* \phi_f(\mathbf{k}) e^{i\omega_{\mathbf{k}}T_f} - \frac{1}{2} \int d\mathbf{k} \omega_{\mathbf{k}} \phi_f(\mathbf{k}) \phi_f(-\mathbf{k})$$

Saddle Point Approximation



$$\begin{aligned}\phi &\rightarrow \Phi/g \\ \phi' &\rightarrow \Phi'/g \\ J &\rightarrow j/g \\ b &\rightarrow \beta/g \\ b^* &\rightarrow \beta^*/g \\ E &\rightarrow \epsilon/g^2 \\ N &\rightarrow n/g^2\end{aligned}$$

$$S(\phi, m, g) \rightarrow S(\Phi/g, m, g) = \frac{1}{g^2} s(\Phi, m)$$

$$g^2 = \hbar$$

Saddle Point Approximation



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 \end{aligned}$$

$$S(\phi, m, g) \rightarrow S(\Phi/g, m, g) = \frac{1}{g^2} s(\Phi, m)$$

$$\phi_i(\mathbf{k}) = \frac{a_{-\mathbf{k}}^*}{\sqrt{2\omega_{\mathbf{k}}}} e^{i\omega_{\mathbf{k}}t}$$

$$\phi_f(\mathbf{k}) = \frac{1}{\sqrt{2\omega_{\mathbf{k}}}} (b_{\mathbf{k}} e^{\omega_{\mathbf{k}}T - \theta - i\omega_{\mathbf{k}}t} + b_{-\mathbf{k}}^* e^{i\omega_{\mathbf{k}}t})$$

$$g^2 = \hbar$$

Saddle Point Approximation



$$g^2 = \hbar$$

$$\begin{aligned}\phi &\rightarrow \Phi/g \\ \phi' &\rightarrow \Phi'/g \\ J &\rightarrow j/g \\ b &\rightarrow \beta/g \\ b^* &\rightarrow \beta^*/g \\ E &\rightarrow \epsilon/g^2 \\ N &\rightarrow n/g^2\end{aligned}$$

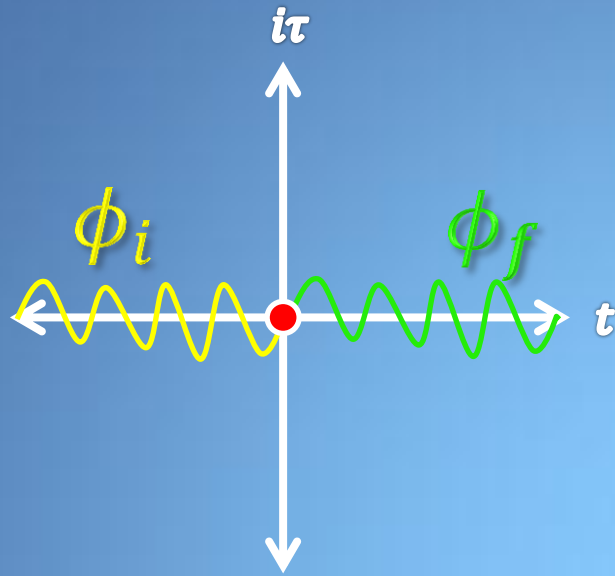
$$S(\phi, m, g) \rightarrow S(\Phi/g, m, g) = \frac{1}{g^2} s(\Phi, m)$$

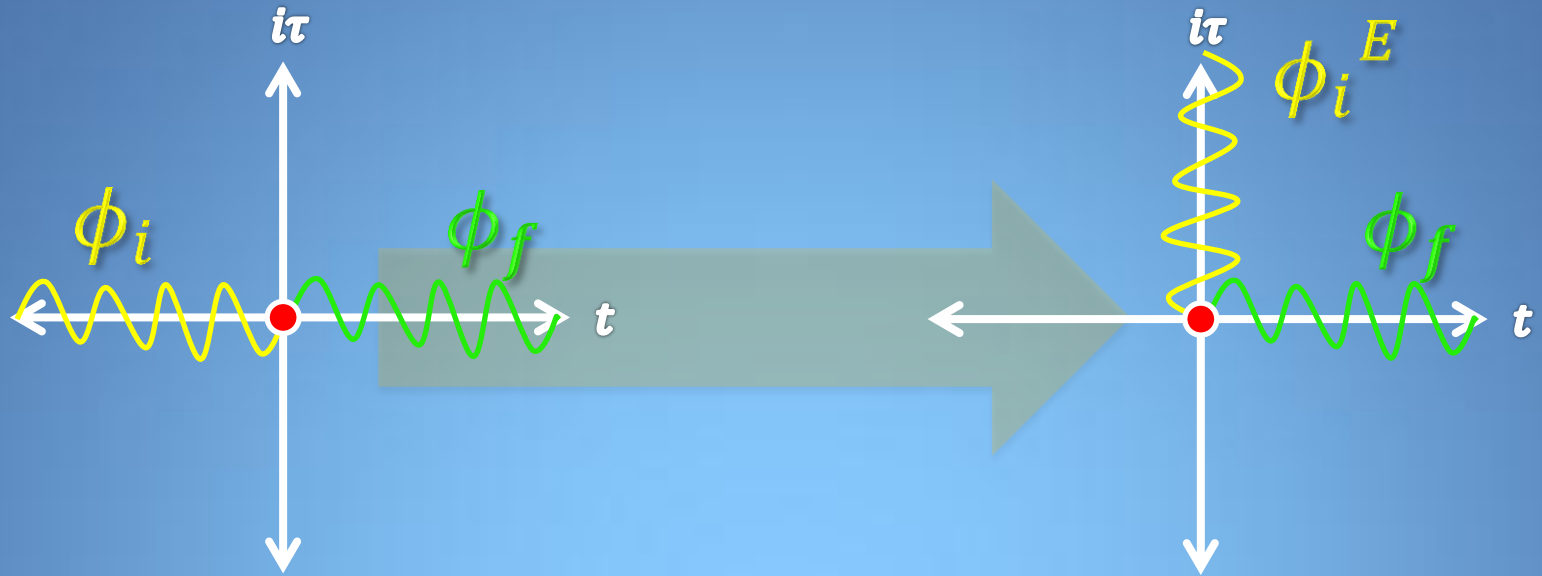
$$\phi_i(\mathbf{k}) = \frac{a_{-\mathbf{k}}^*}{\sqrt{2\omega_{\mathbf{k}}}} e^{i\omega_{\mathbf{k}}t}$$

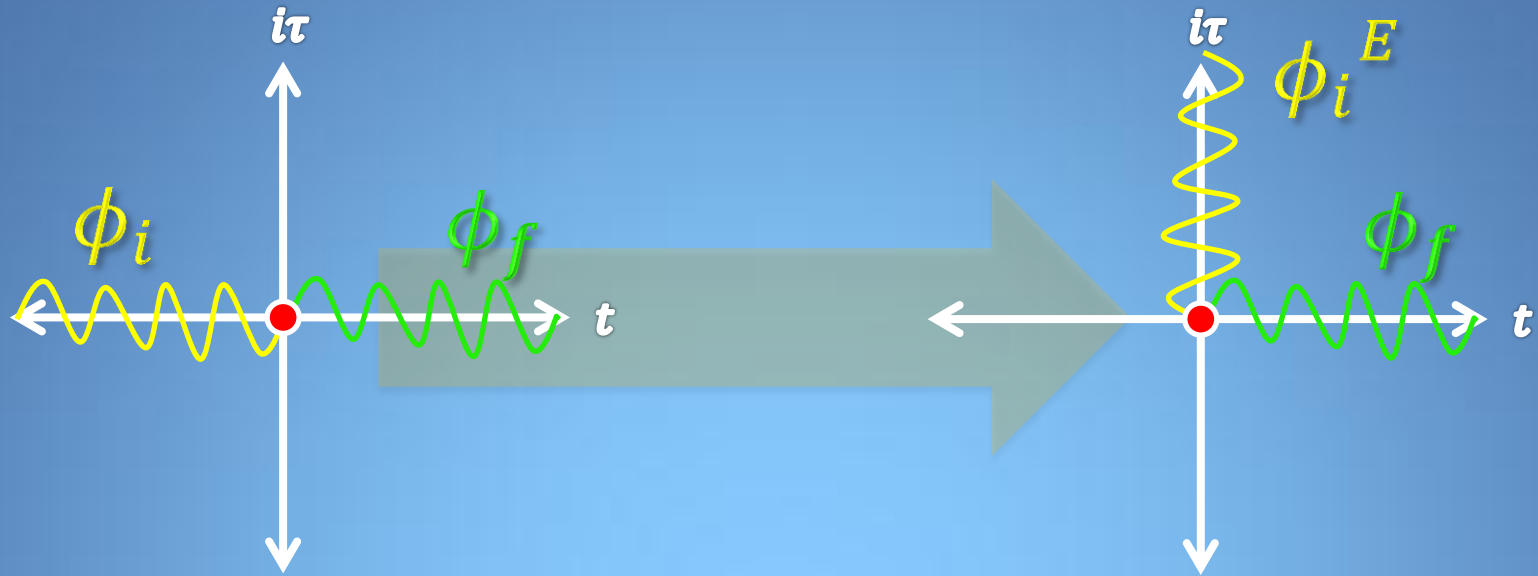
$$\phi_f(\mathbf{k}) = \frac{1}{\sqrt{2\omega_{\mathbf{k}}}} (b_{\mathbf{k}} e^{\omega_{\mathbf{k}}T - \theta - i\omega_{\mathbf{k}}t} + b_{-\mathbf{k}}^* e^{i\omega_{\mathbf{k}}t})$$



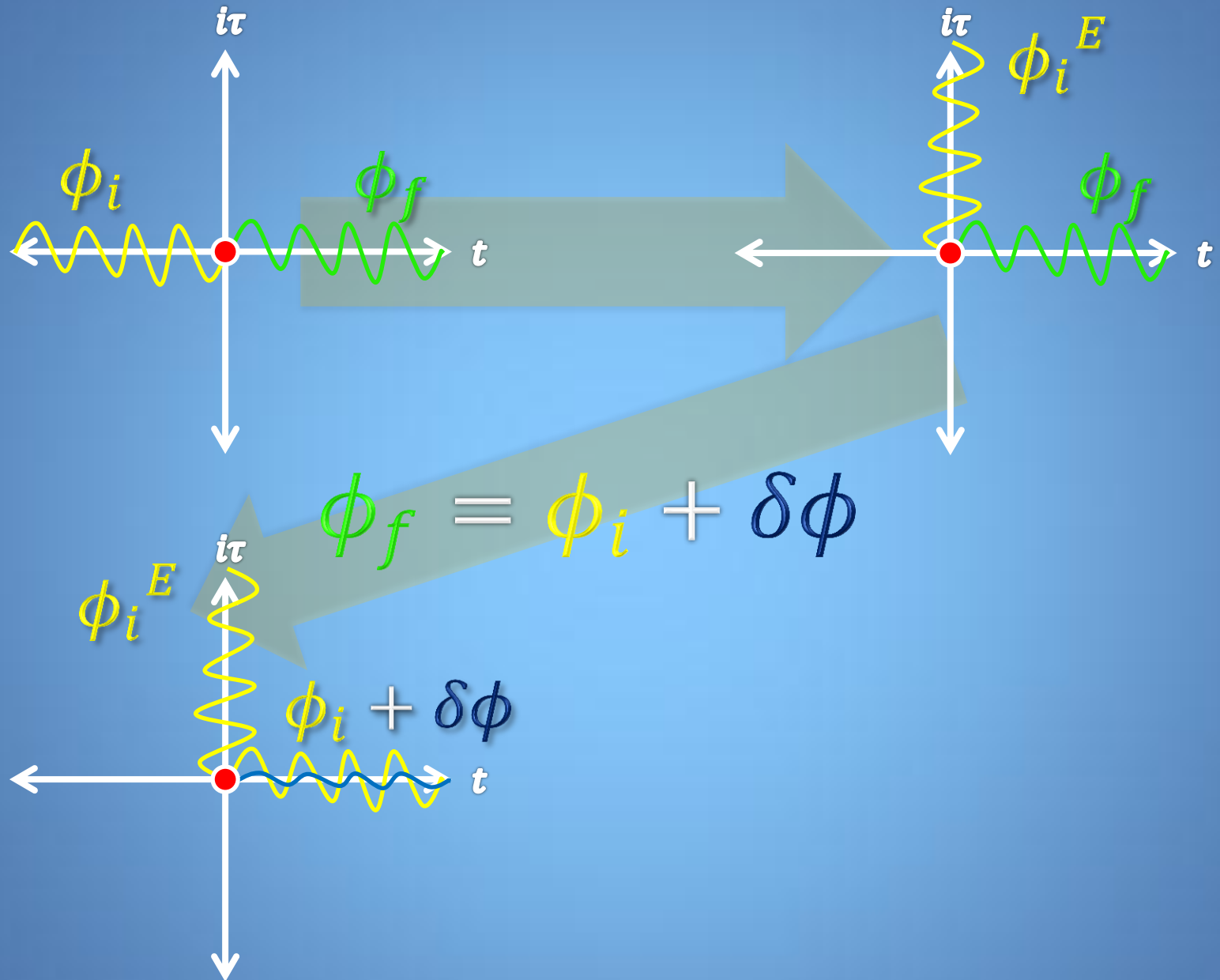
$$\sigma(E, N) \sim e^{-2\Im(S[\phi])/g^2}$$

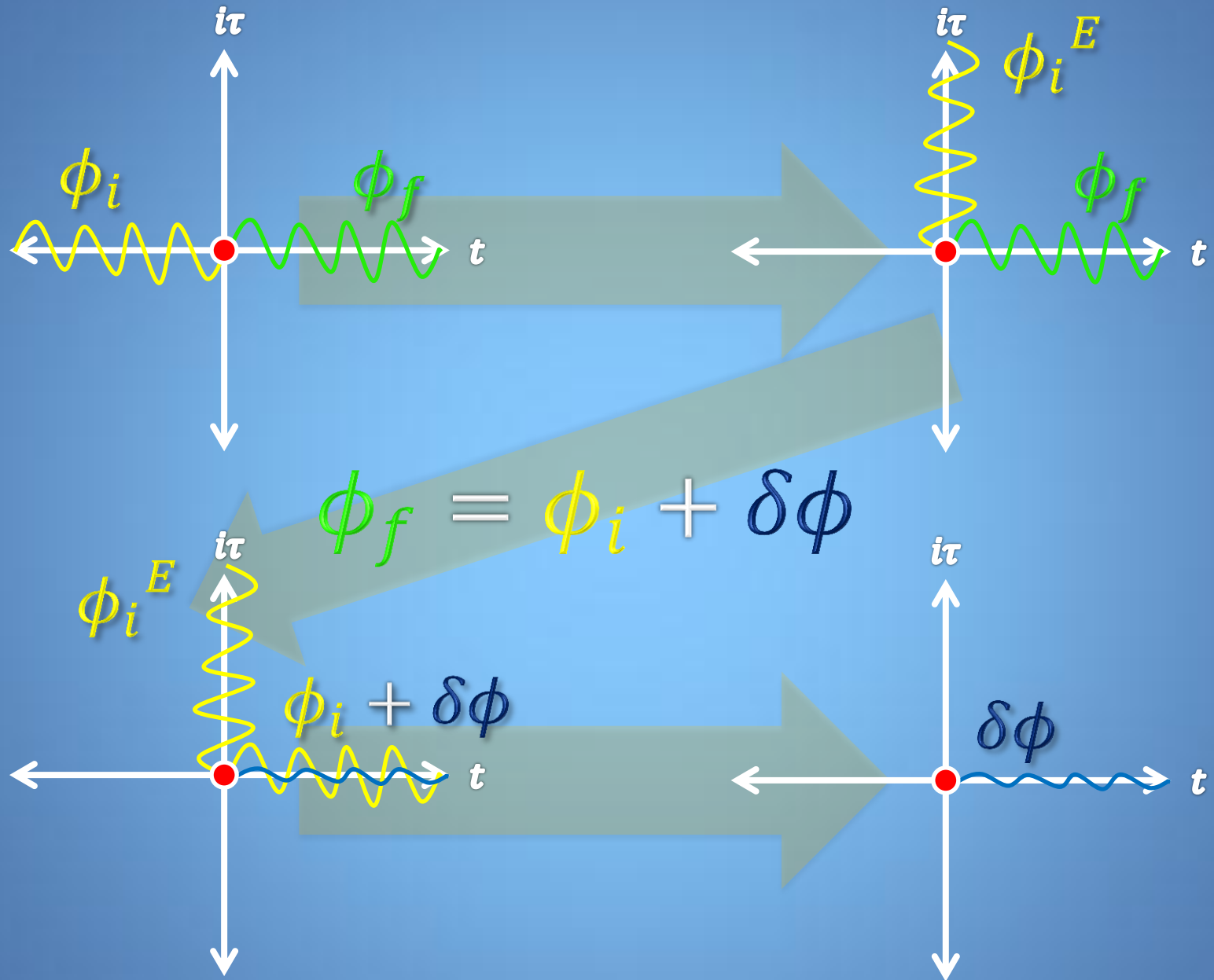






$$\phi_f = \phi_i + \delta\phi$$





Goldstone

$$\sigma(E, N) = \left(\frac{N_{\text{crit}}}{N} \right)^{3N}$$

$$N_{\text{crit}}^3 = c^3 \left(\frac{E}{\Lambda} \right)^4$$

Goldstone

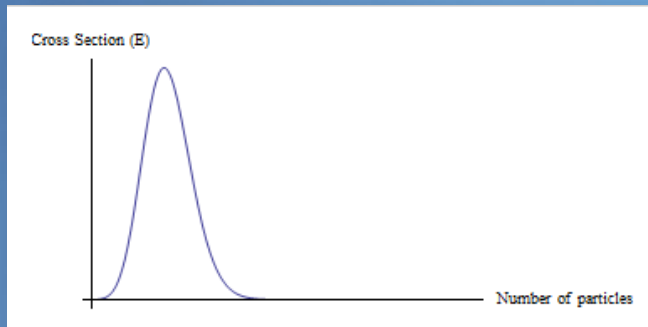
$$\sigma(E, N) = \left(\frac{N_{\text{crit}}}{N} \right)^{3N}$$

Classical length
scale r_*
dominates
behavior of
quantum
scattering
cross section.

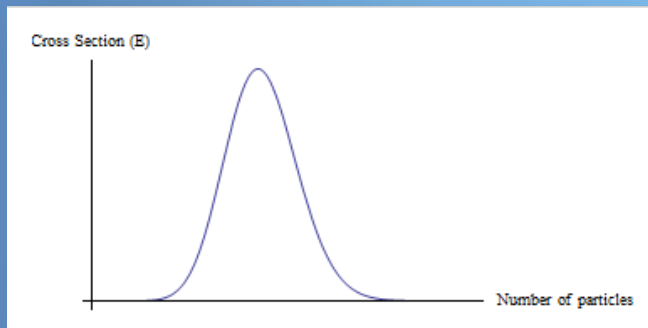
$$N_{\text{crit}}^3 = c^3 \left(\frac{E}{\Lambda} \right)^4$$

$$\frac{1}{r_{\text{crit}}} = \frac{E}{N_{\text{crit}}}$$

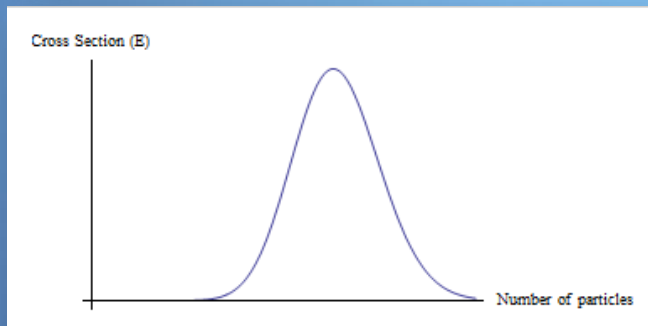
$$r_{\text{crit}} = \frac{c}{\Lambda} \left(\frac{E}{\Lambda} \right)^{1/3} = r_*$$



E



E



E

As energy increases, the preferred cross section is the one that has more particles, each with low energy, as oppose to a few particles with large energy.

Galileon

$$\sigma(E, N) = \left(\frac{N_{\text{crit}}}{N} \right)^{5N}$$

$$N_{\text{crit}}^3 = c^5 \left(\frac{E}{\Lambda} \right)^6$$

$$\frac{1}{r_{\text{crit}}} = \frac{E}{N_{\text{crit}}}$$

$$r_{\text{crit}} = \frac{c}{\Lambda} \left(\frac{E}{\Lambda} \right)^{1/5} = r_*$$

Summary

- r_* phenomenon is important for **Classicalization** and needs to be taken into account for scattering process. Perturbation theory can't see this effect.
- These models admit positive **finite self energy**.
- Many **soft quanta are preferred** over few hard quanta.

Future Work

- Construct **path integral formalism** to capture classicalization effect.