



# Sunyaev Zel'dovich Array Observations of Strong Lensing Galaxy Clusters

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SZA Collaboration

Sloan Giant Arcs Collaboration, RCS Collaboration

# Strong Lensing Galaxy Clusters

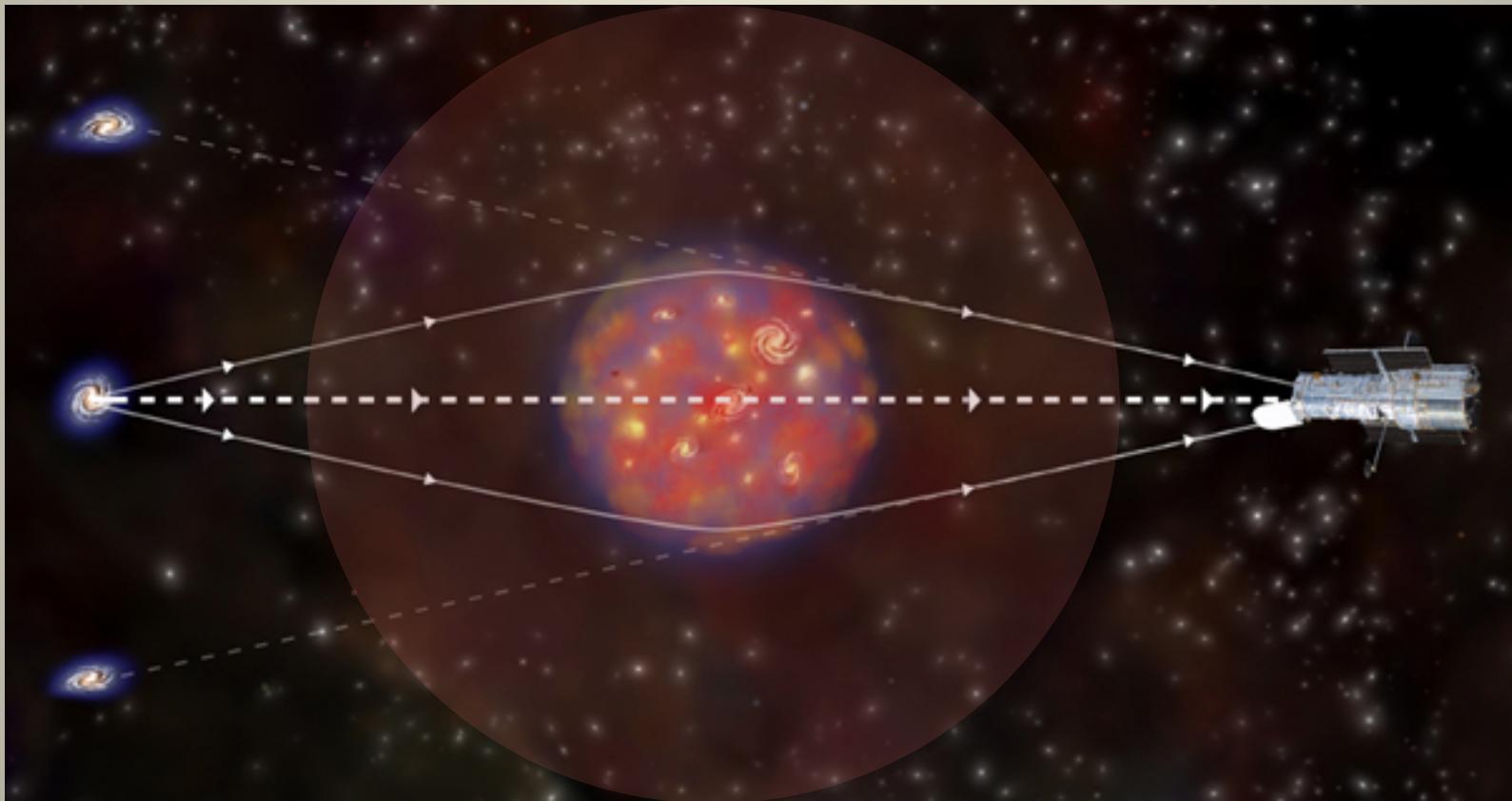


Illustration: NASA/CXC/M.Weiss

# Strong lensing selection

- Ingredients that increase your strong lensing chances:
    - Mass
    - Concentration
    - Line of sight elongation
    - Substructure
    - BCG mass
    - Baryonic physics
    - Surrounding large scale structure
- 
- Hennawi, et al. 2007
- Meneghetti et al. 2007
- Meneghetti et al. 2003
- Rozo et al. 2008
- Puchwein & Ewald 2009

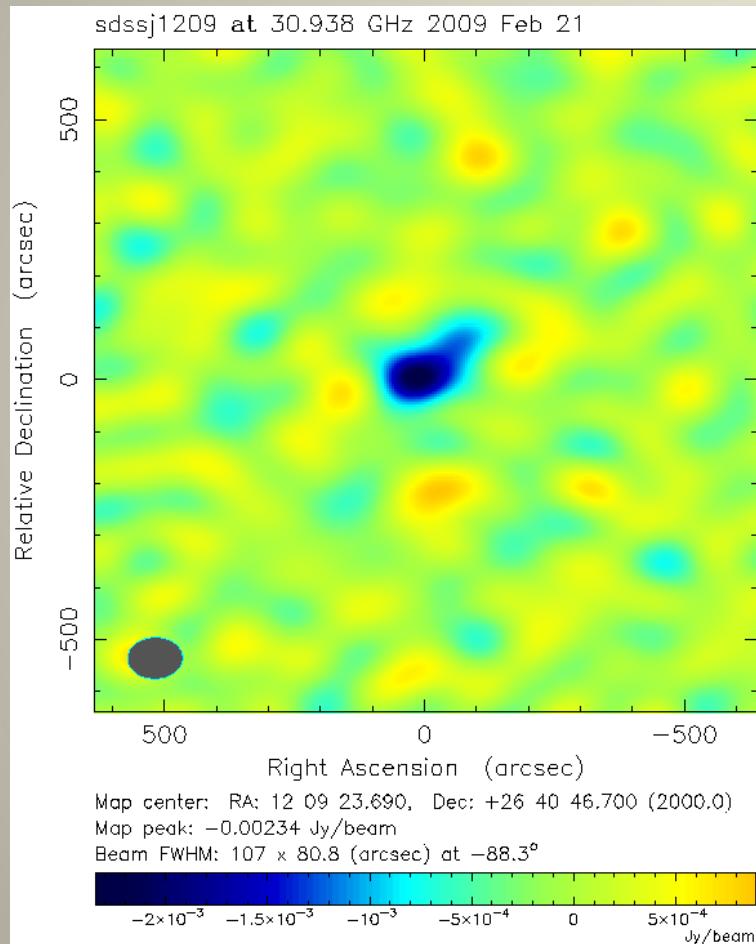
# The problem: A mismatch between theory and observations?

- Overconcentration problem:
  - Oguri et al. (2005) compared CDM predictions for halo triaxiality with A1689 lensing measurements: found 6% of cluster halos can match A1689 at  $2\sigma$
  - Hennawi et al. (2007) used simulations to predict average concentrations of lensing clusters. Observations of lenses indicated higher concentrations than predicted
  - Problem persists with larger samples: eg., Oguri et al. (2009)
  - But for the few clusters studied to date, this is not undisputed: eg., Limousin et al. (2008); Morandi et al. (2010)
- Broadhurst & Barkana (2008) recast this from concentration measurement to comparison between virial masses and core masses. They showed that observed Einstein radii for 4 well-studied clusters lie well beyond predicted distribution in the standard  $\Lambda$ -CDM cosmology

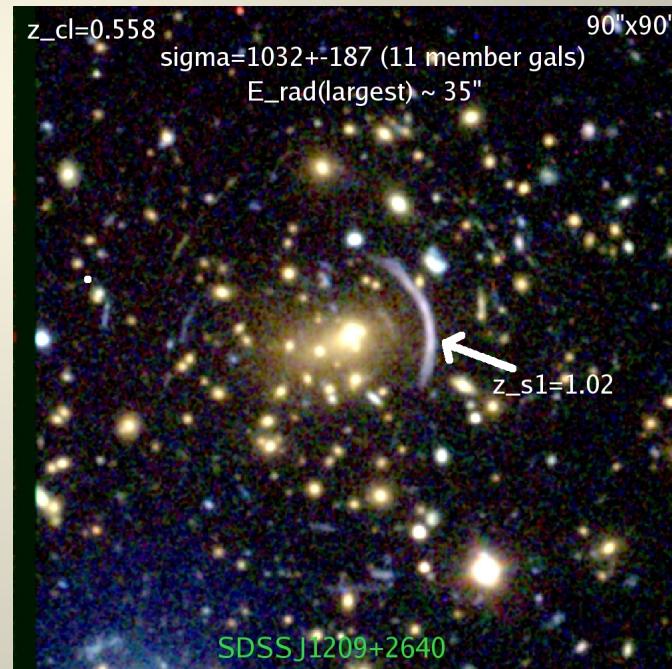
# Strong lensing clusters in our sample



# SZA observations of SDSSJ1209

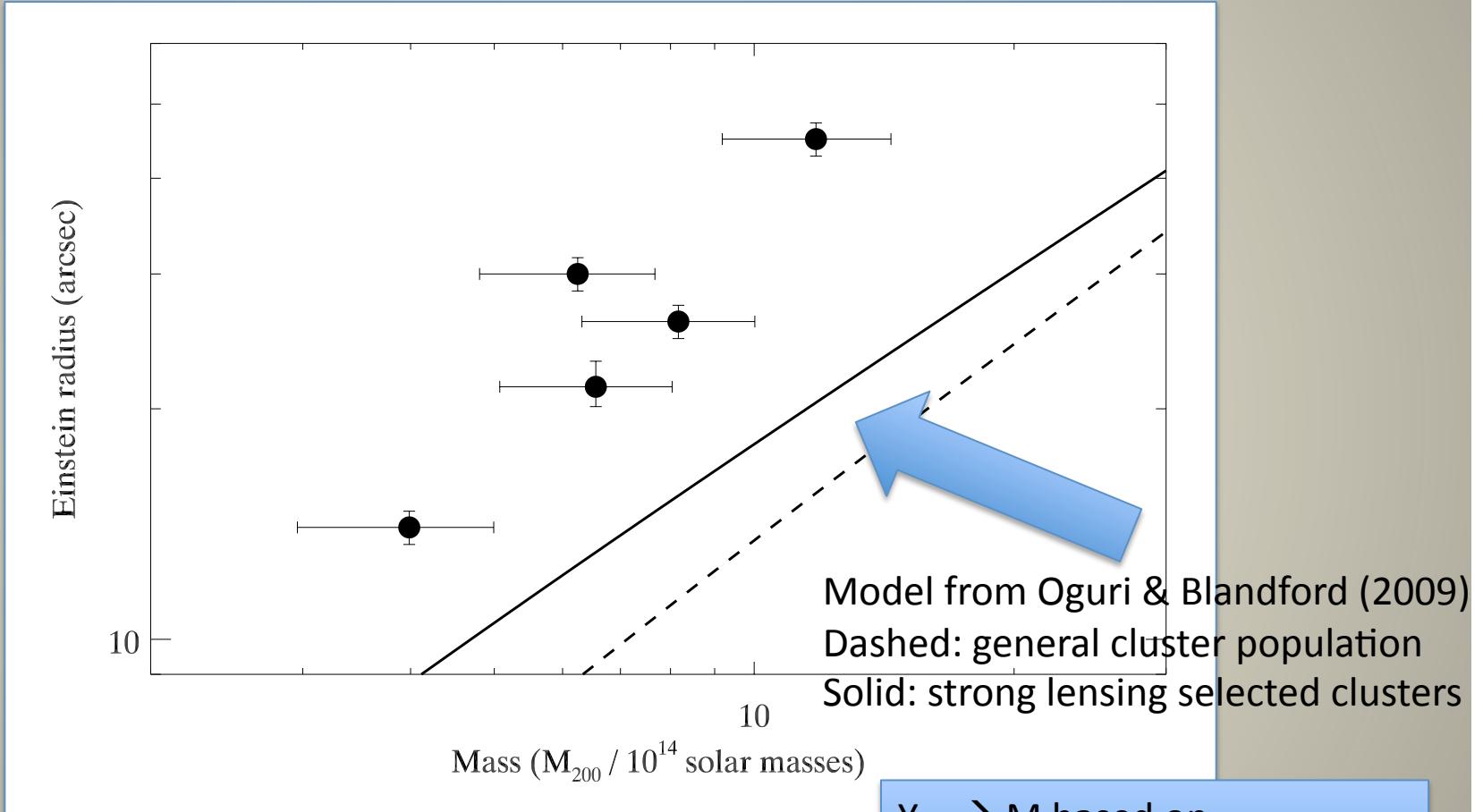


30 GHz SZ observations. 2 off-center point sources removed from map. 11.3 hours of on-source, un-flagged observing time



Discovery paper: Ofek, Seitz & Klein (2008)

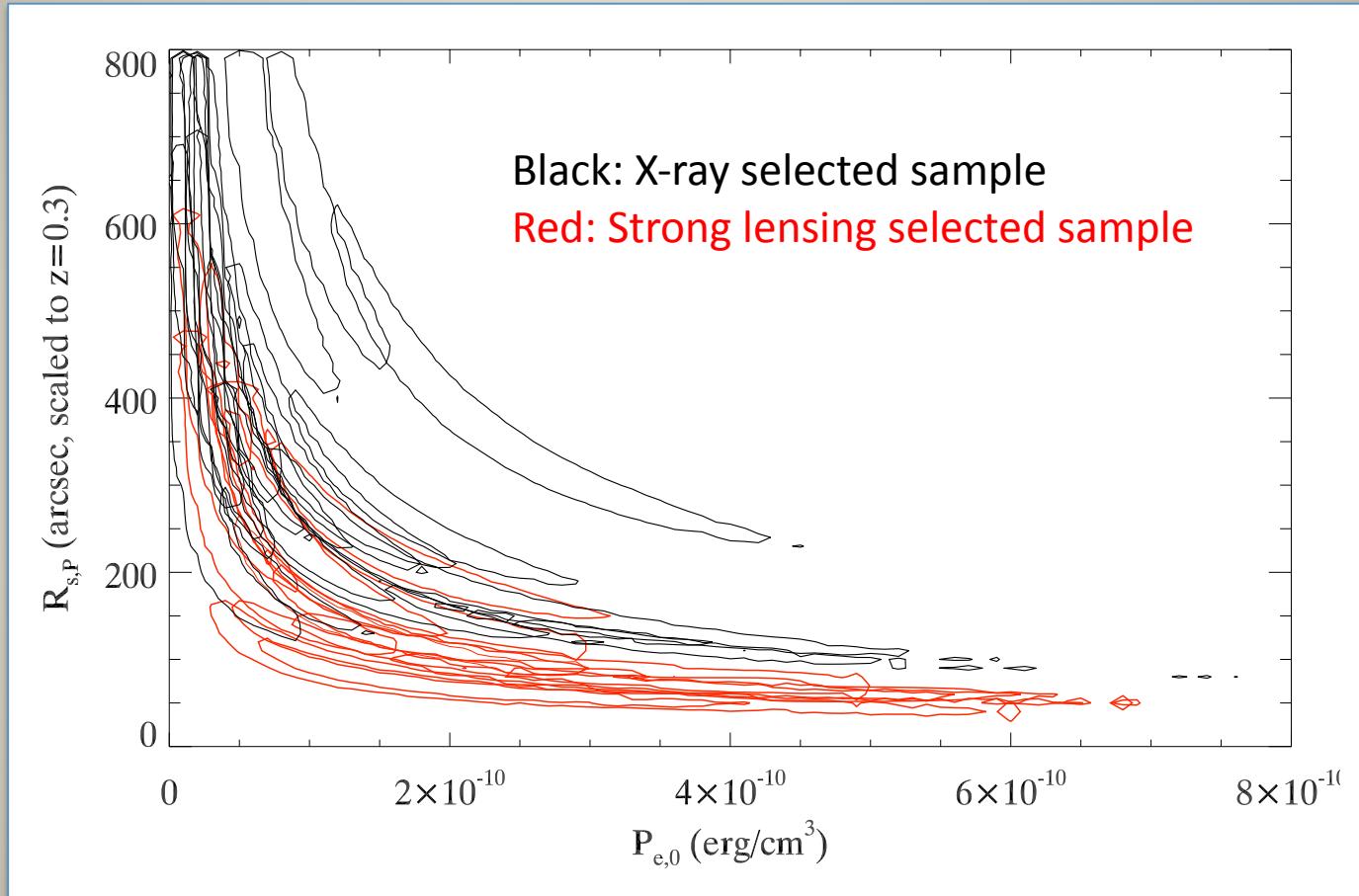
# Masses of lenses



Theory and observations do not agree!

$Y_{\text{SZ}} \rightarrow M$  based on  
scaling relation using LOCUSS  
Weak lensing: Okabe et al.  
SZ: Marrone et al.

# Strong Lensing Selection in the SZ Data



The SZ data indicate that for a given pressure ( $\rightarrow$ mass), strong lensing selected clusters have smaller scale radii (higher 2D concentration) than X-ray selected clusters

# Conclusions

- SZ observations indicate that strong lensing clusters have smaller scale radii
- We find that strong lensing selected clusters of a given mass are more efficient strong lenses than theory predicts
- Future work:
  - Observing a subsample at 90 GHz to better constrain shapes of pressure profiles at higher angular resolution
  - Combining with archival X-ray data, particularly to test for elongation along the line of sight
    - Have analyzed data for SDSSJ1115 and find line of sight elongation 1.6

