

# Modeling Scintillation and Ionization Mechanisms in NEST for low energy recoils in Xenon



Kaixuan Ni  
University of California, San Diego

Low Energy Calibration Workshop  
University of Chicago, Sep.23-25, 2015



Thanks to the fast-growing field of dark matter searches with liquid xenon, data for **low energy** are growing.

<b>Xenon</b>	<b>Light</b>	<b>Charge</b>
<b>Nuclear Recoils</b>	Chepel (1999) Arneodo (2000) Akimov (2002) Aprile (2005) Aprile (2009) Manzur (2010) Plante (2011) Lin (2015) LUX DD (2015) PIXey (2015)	Columbia (2006) Case (2006) Sorensen (2009) 730 V/cm XENON10 (2010) 730 V/cm Manzur (2010) XENON100 (2013) 530 V/cm ZEPLIN-III Lin (2015) LUX DD (2015) PIXey (2015)
<b>Electronic Recoils</b>	Obodovskii (1994) Dahl (2009) Manalaysay (2010) Aprile (2012) Baudis (2013) Lin (2015) PIXey (2015) NeriX (2015)	Dahl (2009) Akimov (2014) Lin (2015) LUX Xe127 (2015) PIXey (2015) NeriX (2015)

# About low energy calibrations

Do we need a model?

**Yes!**

Do we need measurements to fit the model?

**Definitely!**

Do we need to calculate from microphysics to get the parameters?

**It will be interesting to do that.**

# NEST: The Noble Element Simulation Technique

NEST is a(n):

**Detector-independent, data-driven**  
simulation framework

**Comprehensive physical model** of low  
energy interactions in liquid xenon

External package compatible with **Geant4**,  
for easy integration into simulations

**Stand-alone code** for fast calculations of  
yields in simplified situations (Excel sheet  
available so far)



NEST is free and publicly available:

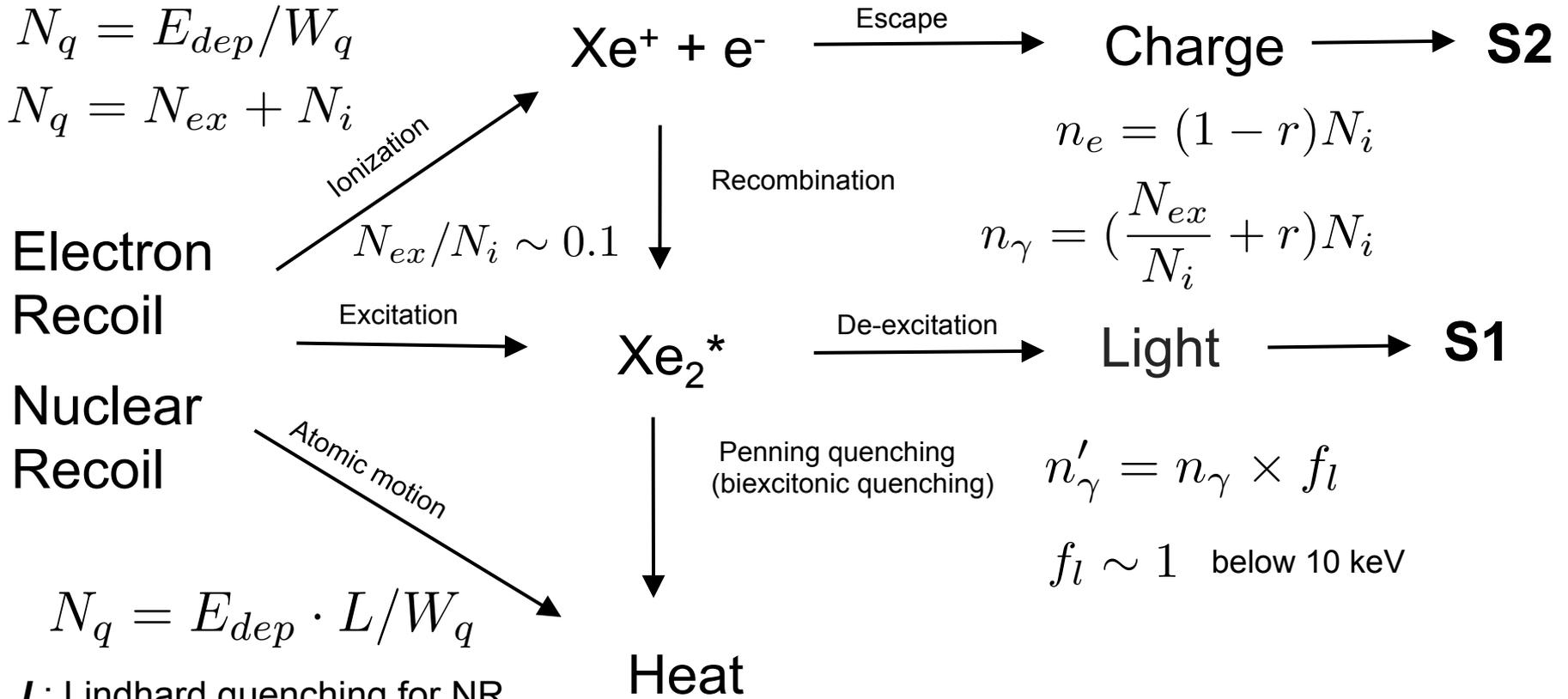
<http://www.albany.edu/physics/NEST.shtml>

<http://nest.physics.ucdavis.edu>

# This talk covers:

1. Low energy nuclear recoils
2. Low energy electron recoils
3. Recombination fluctuation and discrimination

# Signal production in liquid xenon



$L$ : Lindhard quenching for NR  
 Works well for total quanta,  
 should not be confused with  $L_{eff}$  which is only the light

**NEST models this process start-to-finish for both ER and NR**

NEST provides absolute number of electrons AND number of photons for a given **energy**, **field** and **type**.

$$n_e = L \times \frac{E_0}{W} \left( \frac{1}{1 + N_{ex}/N_i} \right) (1 - r)$$

$$n_\gamma = L \times f_l \times \frac{E_0}{W} \left[ 1 - \left( \frac{1}{1 + N_{ex}/N_i} \right) (1 - r) \right]$$

arXiv:1412.4417

$$L = \frac{k g(\epsilon)}{1 + k g(\epsilon)}$$

energy

$$N_{ex}/N_i = \alpha \overset{\text{field}}{F}^{-\zeta} (1 - e^{-\overset{\text{energy}}{\beta\epsilon}})$$

$$r = 1 - \frac{\ln(1 + N_i \zeta)}{N_i \zeta}$$

$$f_l = \frac{1}{1 + \eta \epsilon^\lambda}$$

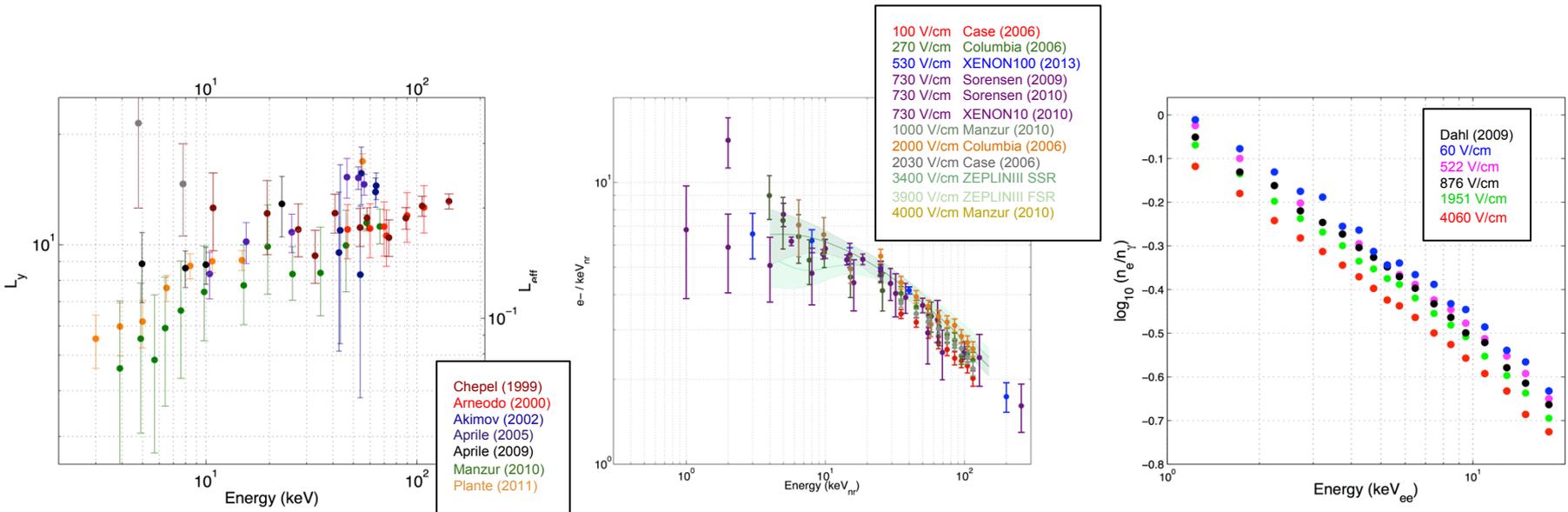
energy

$$\zeta = \gamma \overset{\text{field}}{F}^{-\delta}$$

Thomas-Imel  
recombination model

$$W = 13.7 \text{ eV}$$

# Global fit to the world's data (NR)



$L_{eff}$  - scintillation yield

$Q_y$  - ionization yield

Electron / photon ratio

To fit to all of these data, we construct a global likelihood function and optimize.

$$\mathcal{L} = \prod \frac{1}{\sqrt{2\pi}\sigma_{exp}} \exp\left(\frac{-(x_{exp} - \mu)^2}{2\sigma_{exp}^2}\right) \quad \mu \in \left\{ \mathcal{L}_{eff}, Q_y, \frac{N_e}{N_{ph}} \right\}$$

**arXiv:1412.4417: 2015 data not yet included, but pretty close to the predictions**

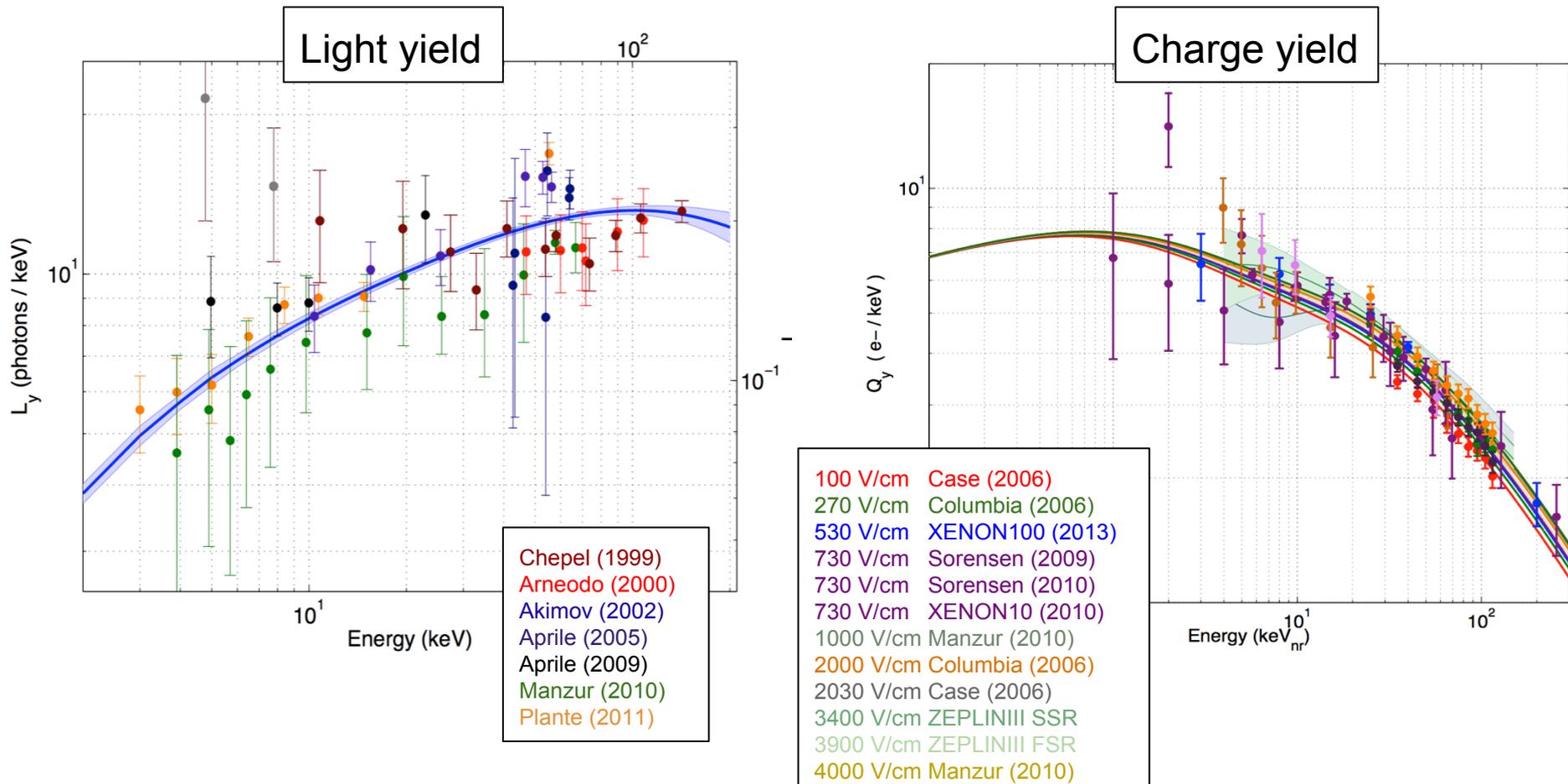
# Advantages of the global analysis

- Constrains light and charge simultaneously; one affects the other
  - Stronger constraints than looking at either individually
- Includes all measurements in the literature in an unbiased way
- Can easily incorporate new data as it becomes available

The resulting model allows us to interpret and simulate experimental data in a way that **reflects the cumulative body of research** on liquid xenon response.

# Global fit to the world's data (NR)

Nuclear recoil model constrained with a global analysis of available data, see arXiv:1412.4417



# NEST best fit for nuclear recoils

$$n_e = L \times \frac{E_0}{W} \left( \frac{1}{1 + N_{ex}/N_i} \right) (1 - r)$$

$$n_\gamma = L \times f_l \times \frac{E_0}{W} \left[ 1 - \left( \frac{1}{1 + N_{ex}/N_i} \right) (1 - r) \right]$$

$$L = \frac{k g(\epsilon)}{1 + k g(\epsilon)} \quad f_l = \frac{1}{1 + \eta \epsilon^\lambda}$$

$$r = 1 - \frac{\ln(1 + N_i \zeta)}{N_i \zeta}$$

$$N_{ex}/N_i = \alpha F^{-\zeta} (1 - e^{-\beta \epsilon})$$

$$\zeta = \gamma F^{-\delta}$$

$$\sigma_R^2 = n_e \times \mathcal{F} = (1 - r) N_i \times \mathcal{F} = (1 - r) \mathcal{C} N_i^2$$

TABLE I. Best fits and 68% confidence intervals of free parameters.

Parameter	Best Fit	68% conf.
$\alpha$	1.240	+0.079 -0.073
$\zeta$	0.0472	+0.0088 -0.0073
$\beta$	239	+28 -8.8
$\gamma$	0.01385	+0.00058 -0.00073
$\delta$	0.0620	+0.0056 -0.0064
$k$	0.1394	+0.0032 -0.0026
$\eta$	3.3	+5.3 -0.06
$\lambda$	1.14	+0.45 -0.09
$\mathcal{C}$	0.00555	+0.00025 -0.00025

**arXiv:1412.4417**

# Systematics in the literature

## Energy scale / uncertainties

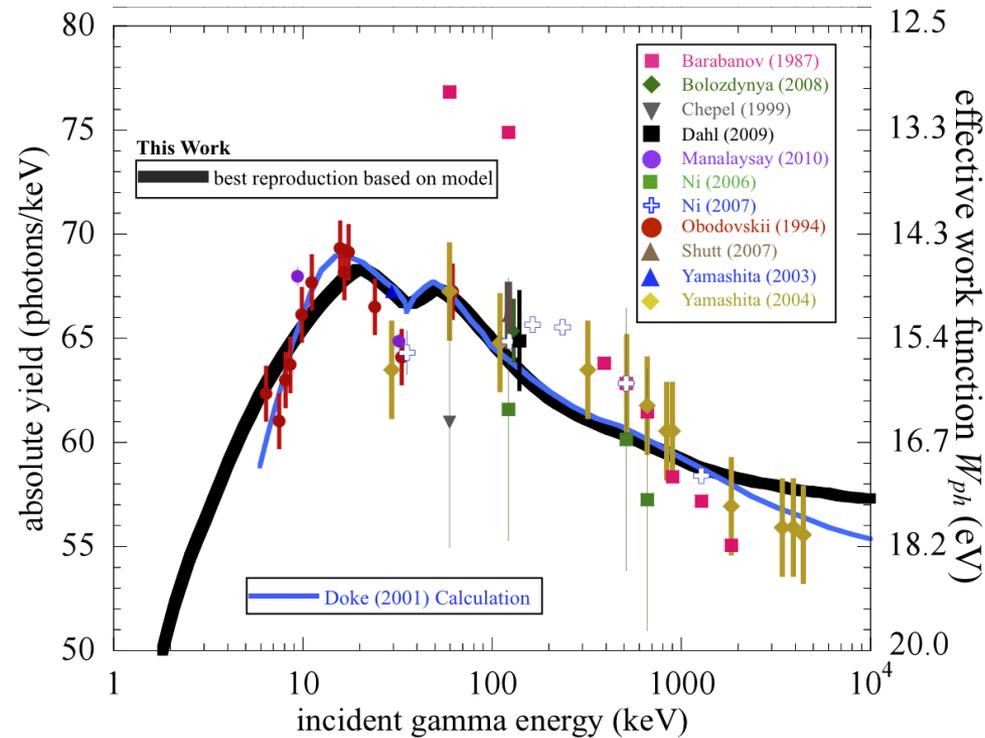
- Energy scales are determined differently in different papers
- Uncertainties in energy are inconsistently reported

## Electron extraction efficiency from liquid

- Results in systematic up/down shift in measured charge yield

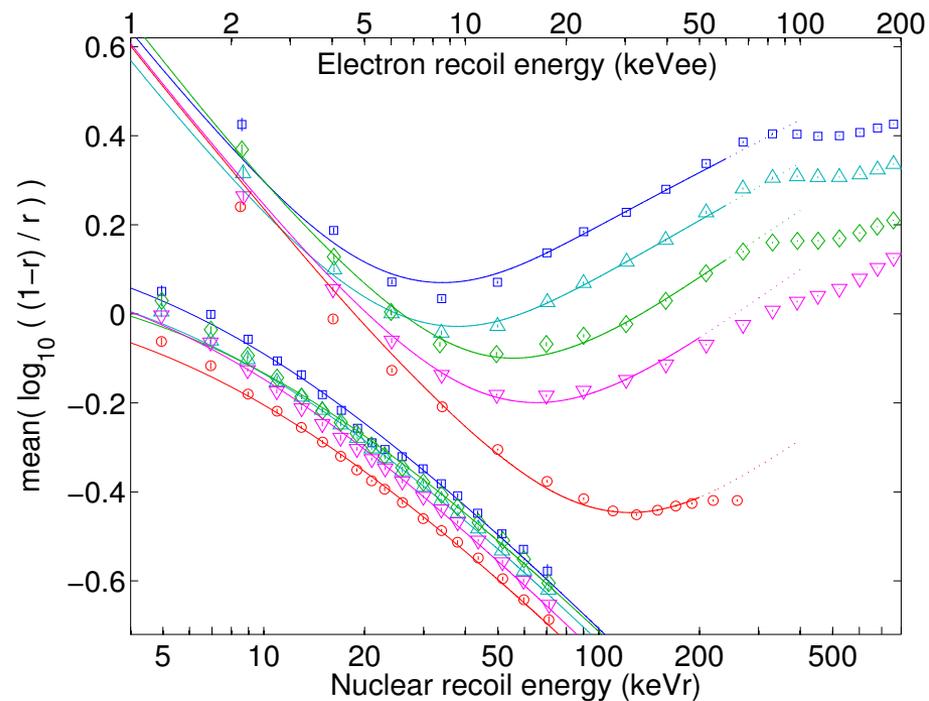
# Scintillation and Ionization for **Electron Recoils** (early work)

## Zero field scintillation (high energy)



Szydakis et al., JINST (2011), arXiv:1106.1613

## Ionization/Scintillation at different fields (low energy)

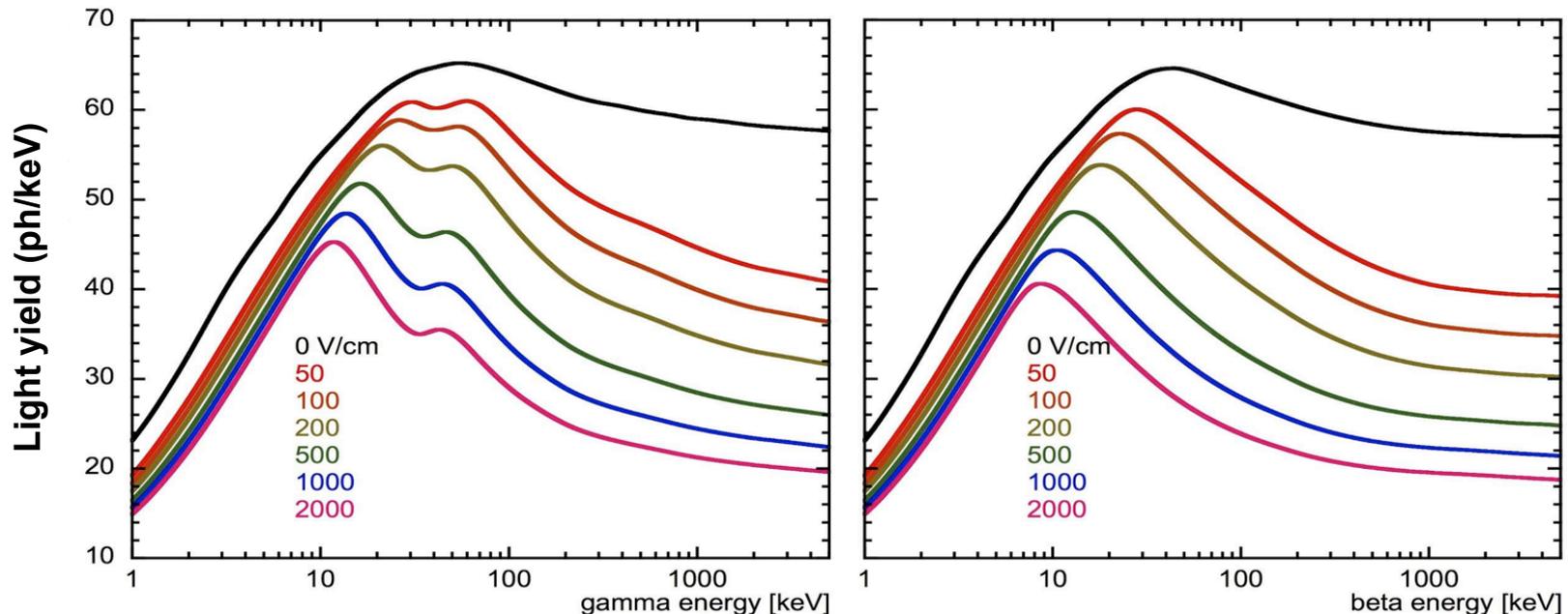


Dahl's PhD thesis (2009)

# The current NEST ER model (v0.98) based on data from early work

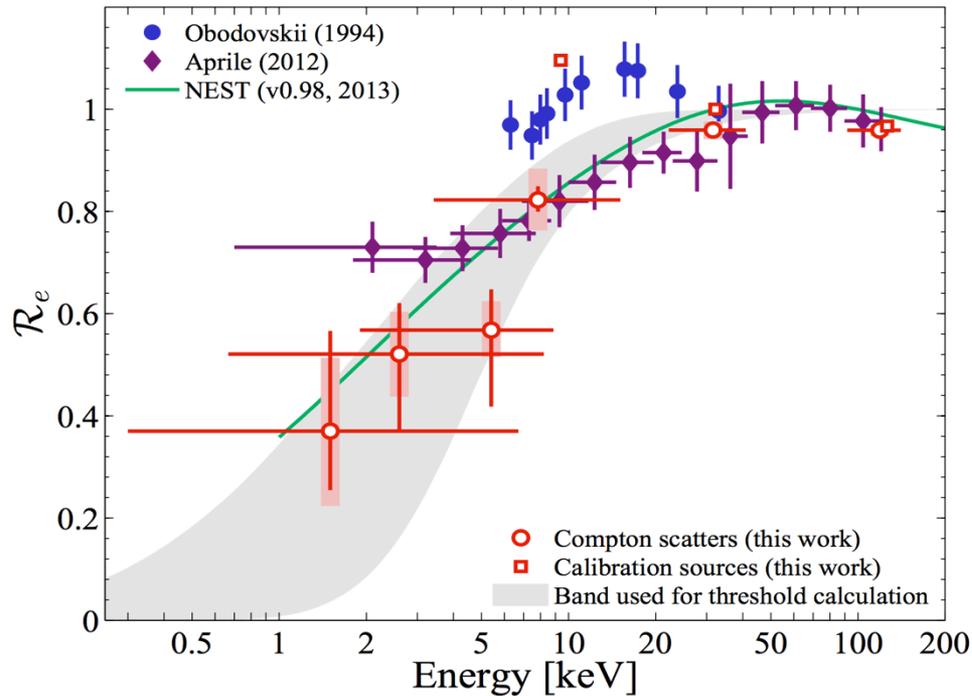
For long tracks (high energy), Birks-Doke law:  $r = \frac{A \frac{dE}{dx}}{1 + B \frac{dE}{dx}} + C, \quad C = 1 - A/B.$

For short tracks (low energy), Thomas-Imel Box model:  $r = 1 - \frac{\ln(1 + \xi)}{\xi}, \quad \xi \equiv \frac{N_i \alpha'}{4a^2 v}$

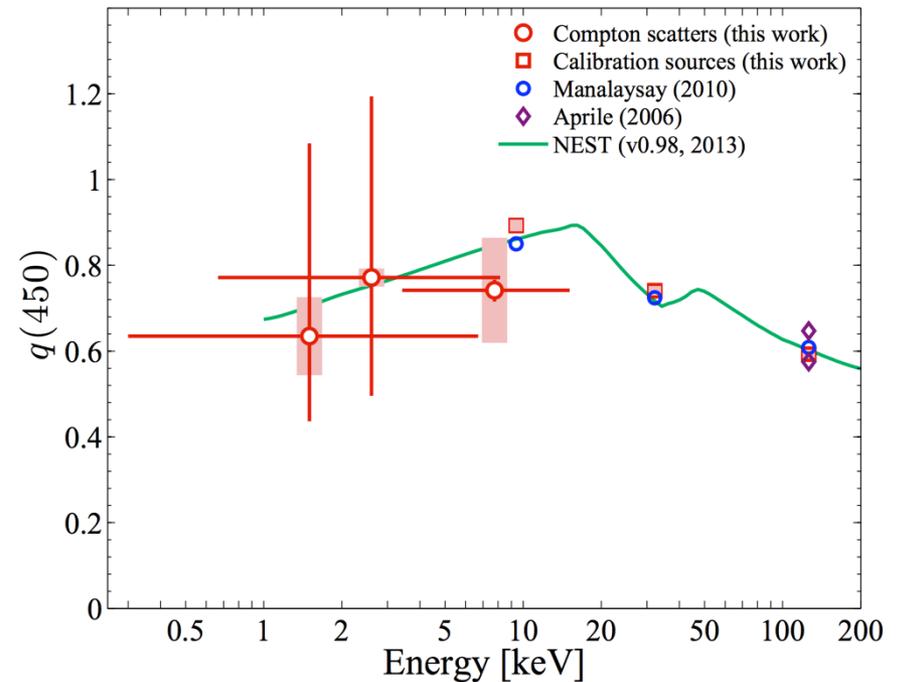


# ER scintillation from **direct** measurements (**Compton scattering, or mono-energetic lines**)

### Zero Field Scintillation

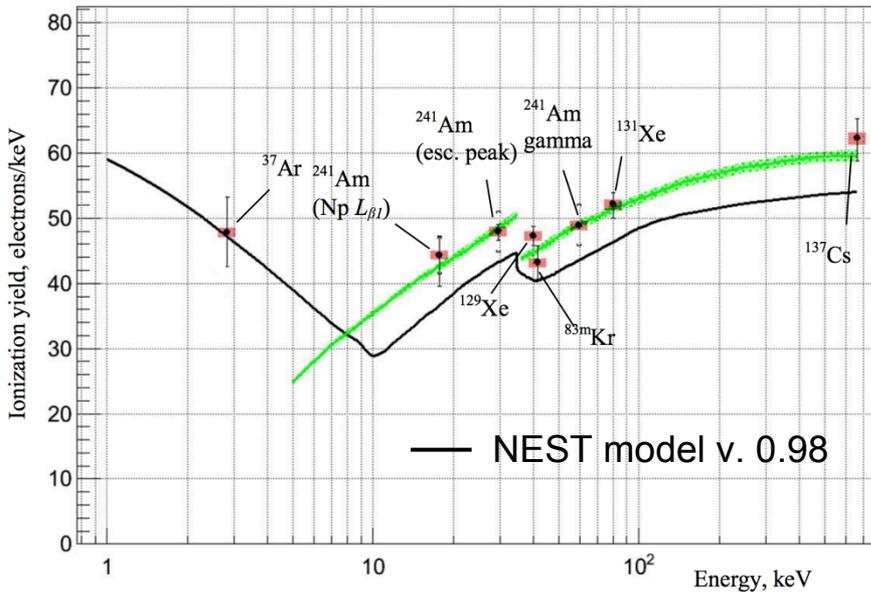


### Scintillation Quenching at 450 V/cm

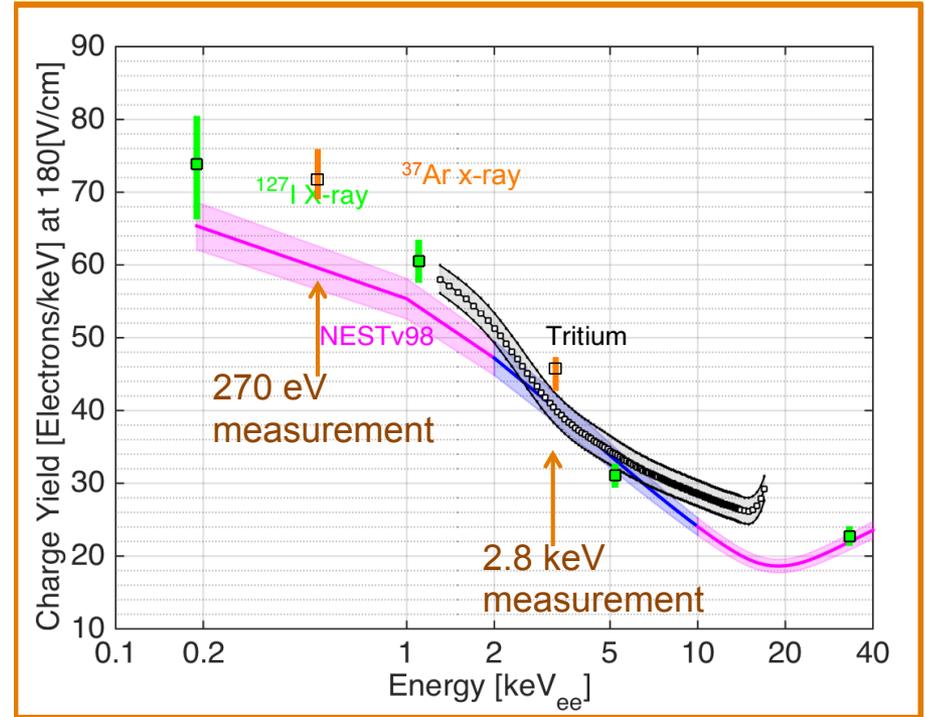


Aprile et al., Phys. Rev. D 86, 112004 (2012)  
Baudis et al., Phys. Rev. D 87, 115015 (2013)

# ER ionization from **direct** measurement – very few data



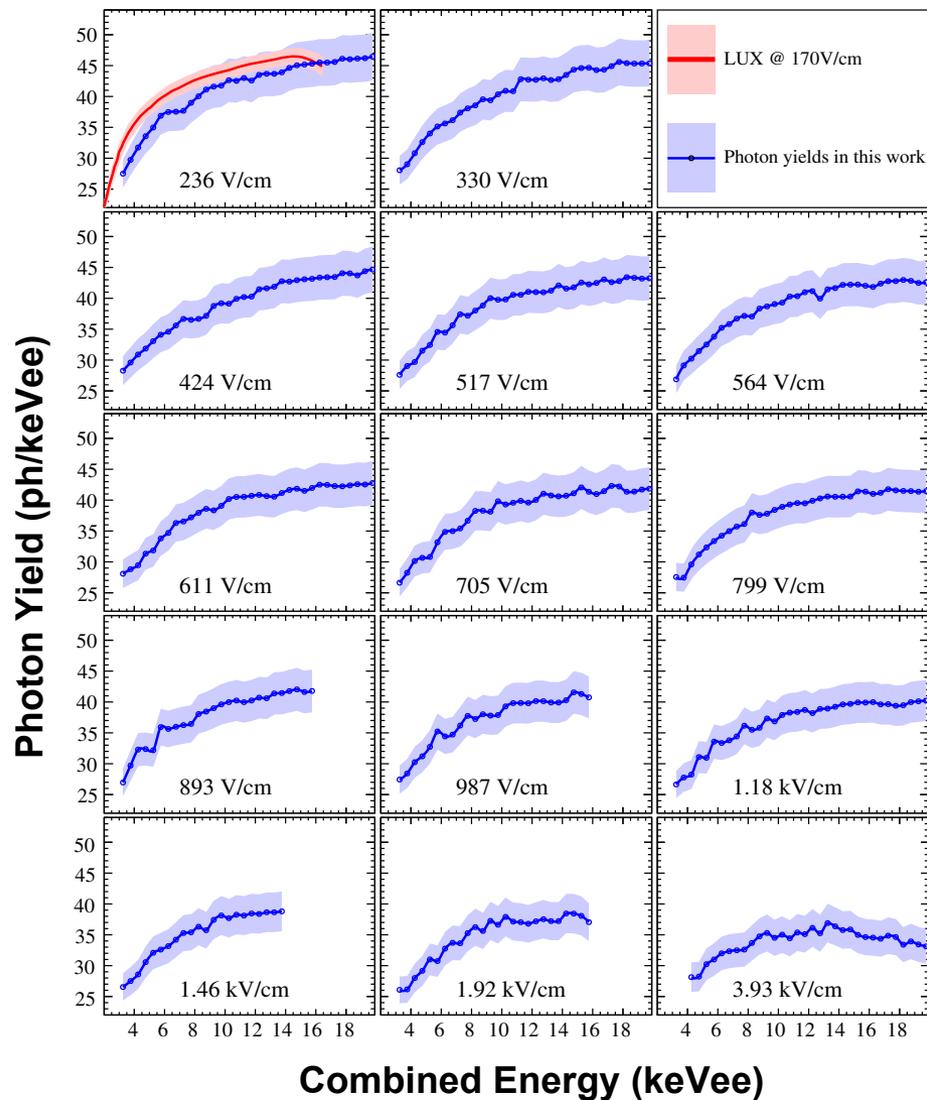
Akimov et al., JINST 9 (2014) P11014



D. Huang, LIDINE2015 (Xe127/LUX)  
E. Boulton, LIDINE2015 (Ar37/PIXeY)

New results from L. Goetzke (this workshop)!

# More recent **indirect** measurement results



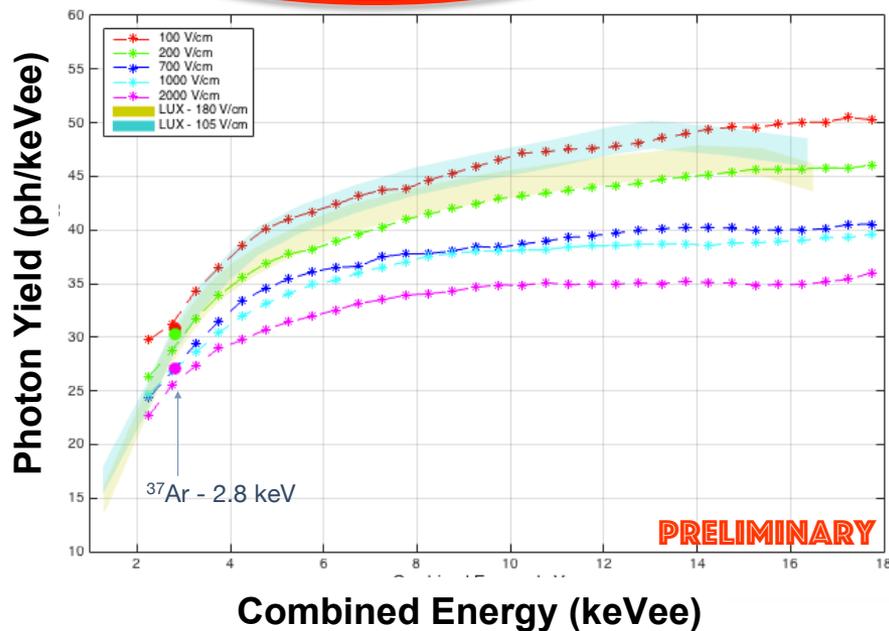
Lin et al., Phys. Rev. D 92, 032005 (2015)

Assumptions:

$$E = (n_\gamma + n_e)W_q$$

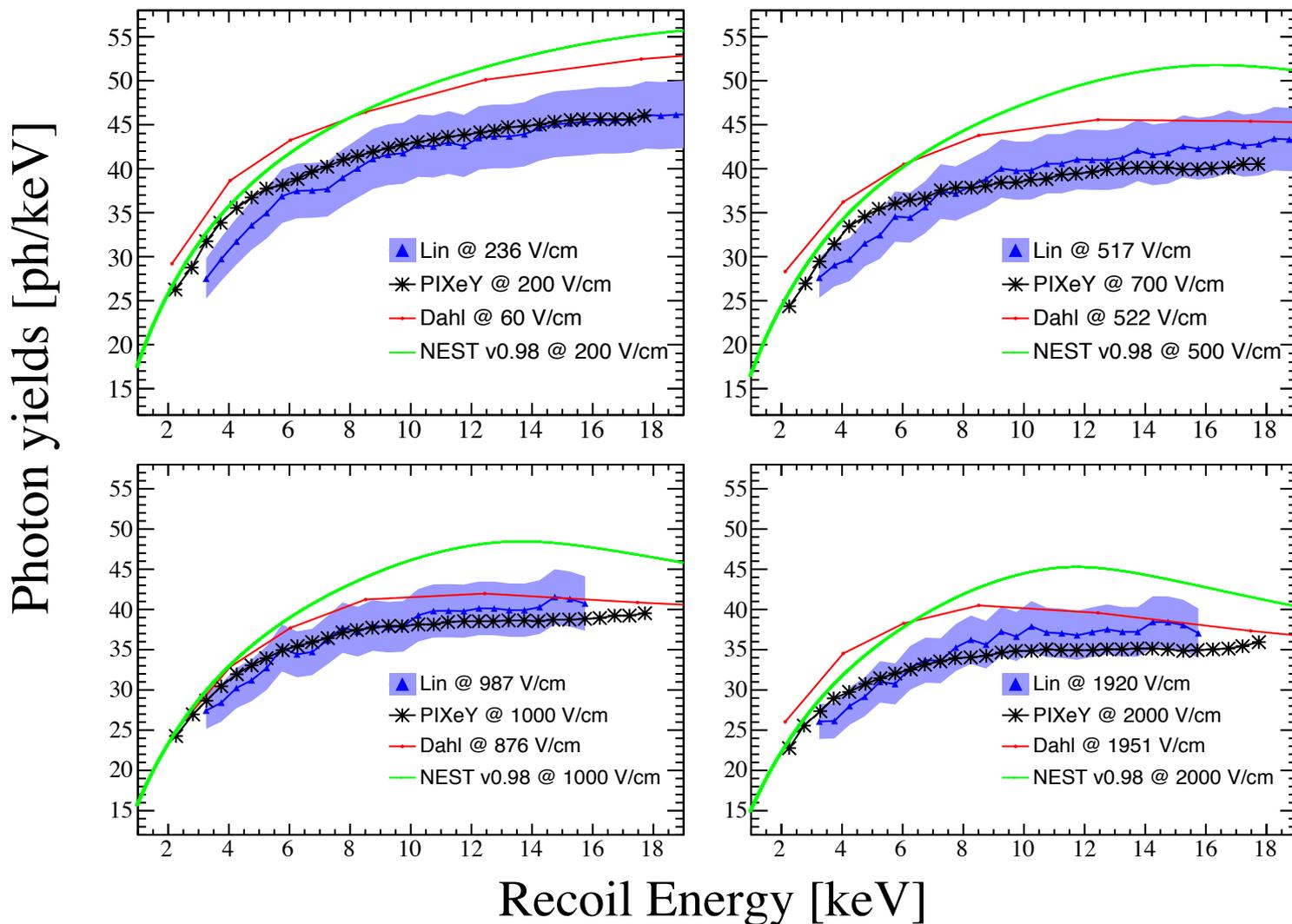
$$= \left( \frac{S1}{g_1} + \frac{S2}{g_2} \right) W_q$$

$$W_q = 13.7 \text{ eV}$$

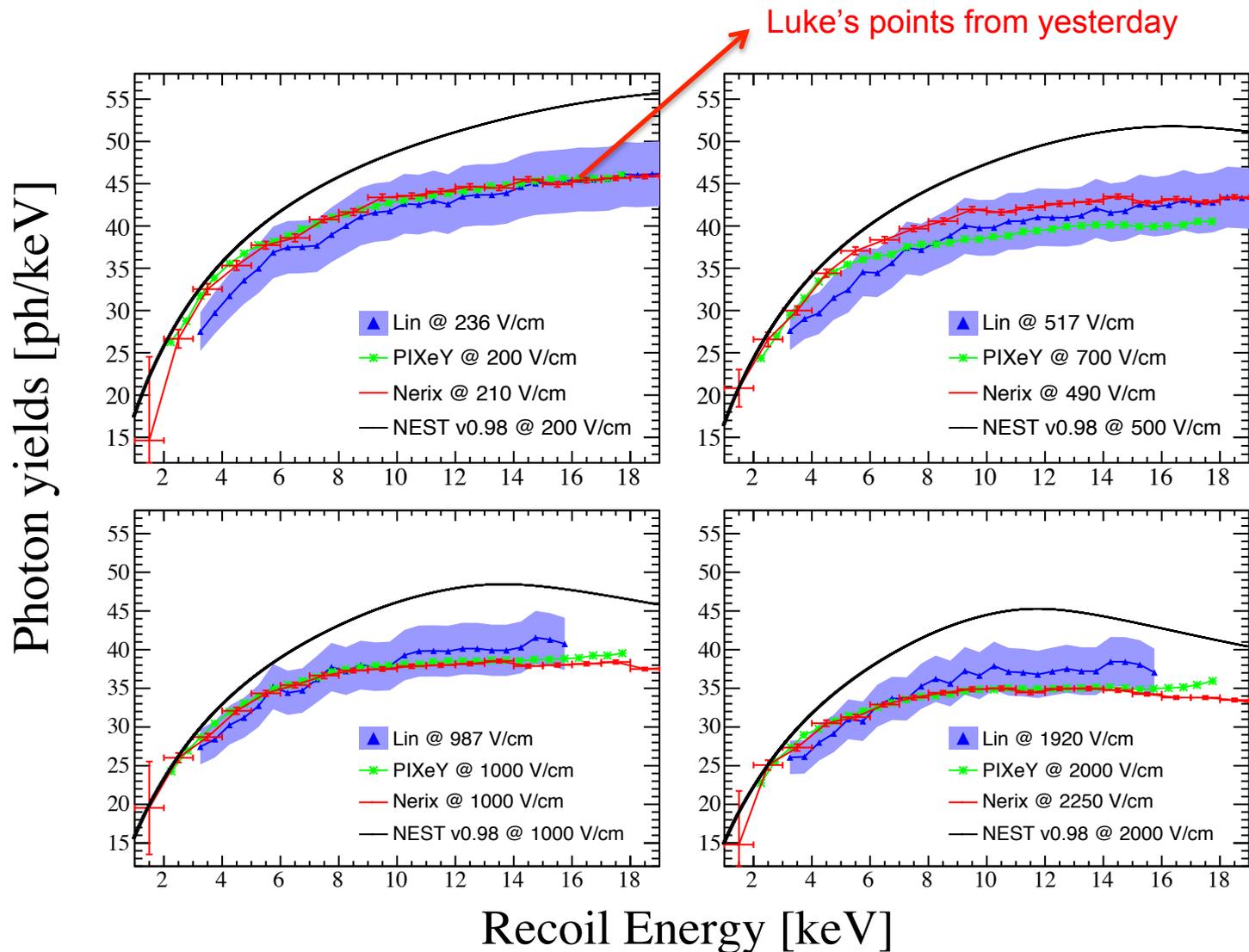


D. McKinsey & B. Edwards, LIDINE 2015

# Comparing the indirect measurements with NEST v0.98 ER model



# Comparing the **recent** measurements with NEST v0.98 ER model



# Assumptions of $W = 13.7$ eV at low energy

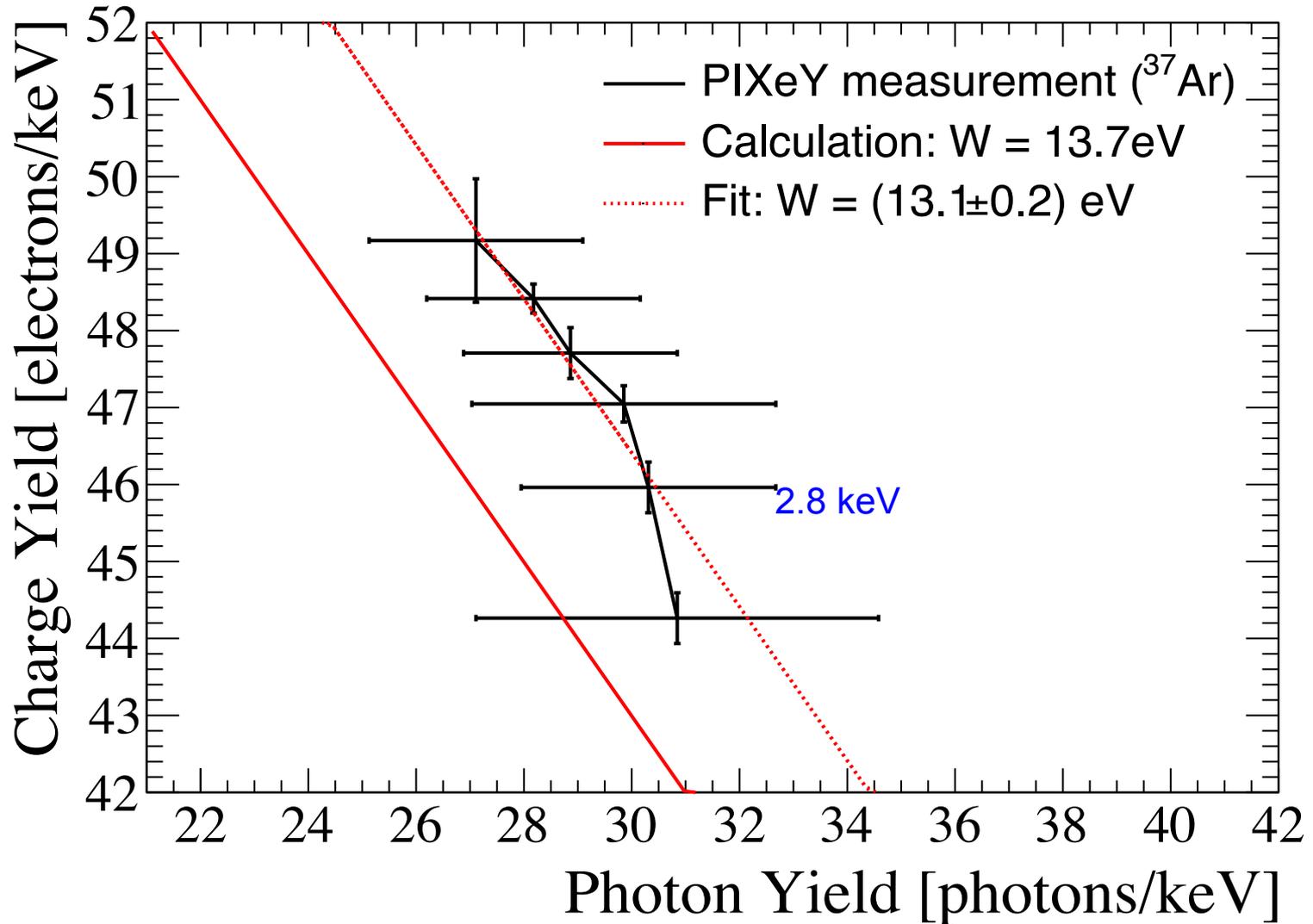


Figure made by Fei Gao from data points from [E. Boulton, LIDINE2015]

# Need a new model for Electron Recoils

- Current NEST ER model (v0.98) **describes new data within 10%, but needs to be improved**
- Current NEST ER model (v0.98) is **not analytical**, require “full” Geant4 simulation of energy deposition
  - with small-scale energy-loss steps recorded
  - with tiny track length cut-off (threshold)
- It needs to merge Thomas-Imel & Doke-Birk laws at medium energy (starting at 5~10 keV and field dependent)
  - require knowledge of  $dE/dx$  and exact track structure
- E. Dahl has a modified Thomas-Imel model but require track structure simulation with PENELOPE

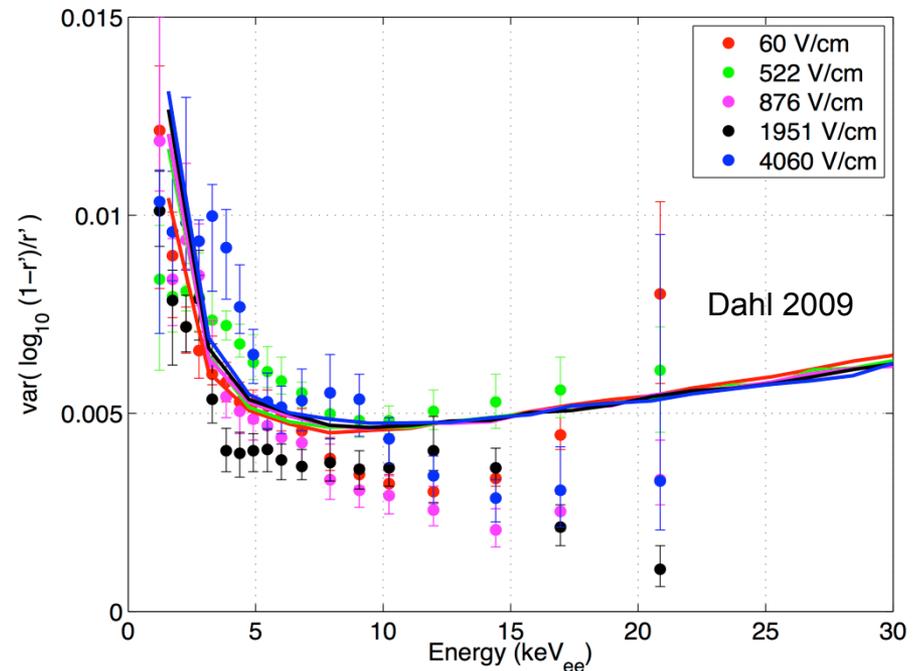
# Modeling recombination fluctuations

In LXe, recombination fluctuations are very different than the expected due to a binomial process (Dobi 2014):

$$\sigma_{bin}^2 = r(1-r)N_i \quad \sigma_{obs}^2 = \frac{1}{185} \times N_i^2$$

In **NEST v1.0 NR model**, a Fano-like factor proportional to  $N_i$  to fit to Dahl's data

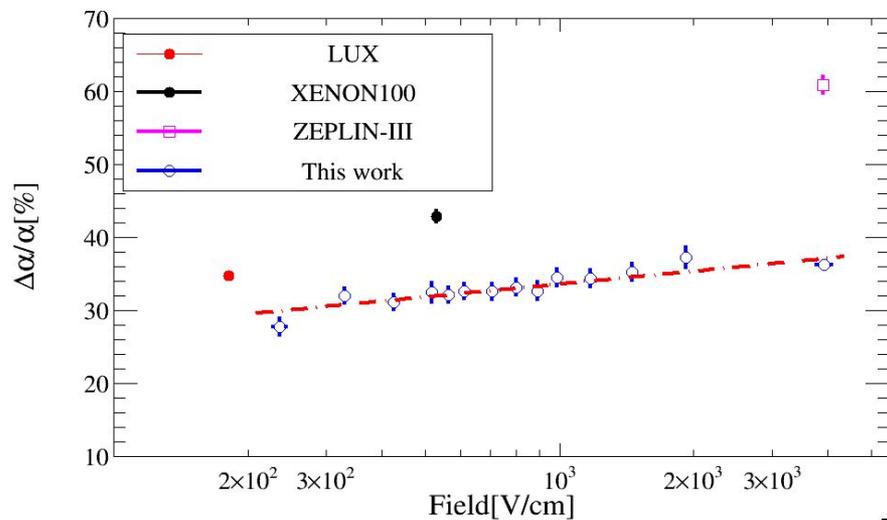
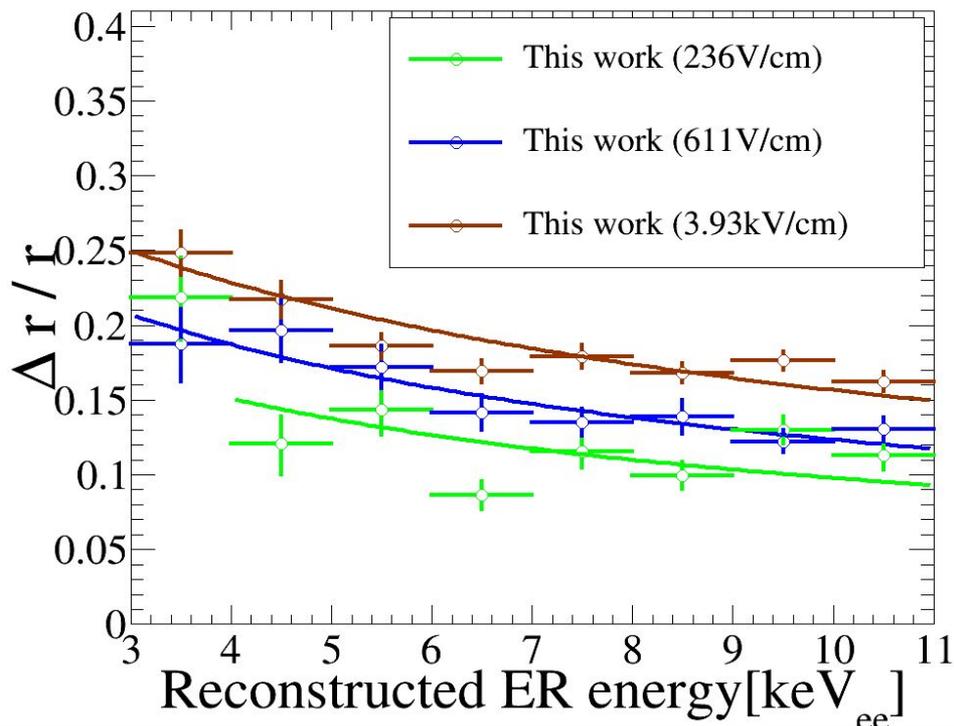
$$\sigma_R^2 = n_e \times \mathcal{F} = (1-r)N_i \times \mathcal{F} = (1-r)\mathcal{C}N_i^2$$



# Probing the intrinsic recombination fluctuations (for **Electron Recoils**, Lin's talk yesterday)

$$r(\xi) = 1 - \frac{1}{\xi} \ln(1 + \xi), \xi = \frac{N_i \alpha}{4a^2 \mu F}$$

$$\Delta r = \left[ \frac{1}{\xi} \ln(1 + \xi) - \frac{1}{1 + \xi} \right] \frac{\Delta \alpha}{\alpha}$$



# The NEST collaboration

## University at Albany, SUNY

Prof. Matthew Szydakis

Dr. Jeremy Mock



## University of California, Davis

Prof. Mani Tripathi

Prof. Emilija Pantic

Dr. Aaron Manalaysay

Brian Lenardo



## Lawrence Livermore National Laboratory

Dr. Kareem Kazkaz

## University of California, San Diego

Prof. Kaixuan Ni

Fei Gao



## University of Tennessee

Prof. Sergei Ovchinnikov

## Stanford Linear Accelerator Center (SLAC)

Prof. Tom Shutt

Prof. Dan Akerib



# Conclusions

- NEST is a tool **by and for the liquid xenon** community for modeling detector response
- Global analysis allows the inclusion of all available data in an unbiased way
  - For low energy NR data, v1.0 is now public ( [arXiv: 1412.4417](#) )
  - For low energy ER data, the current model (v0.98) is outdated. A modified TIB model is being developed (v1.0) incorporating all recent data
- Approaches to understand the intrinsic **recombination fluctuations** in order to predict ER/NR discrimination are underway