

# Odd Nuclei, “ $g_A$ Renormalization”

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# Neutrino-Nucleus Differential Cross Section

From J.D. Walecka, *Semileptonic Weak Interactions*

$$\begin{aligned} \left(\frac{d\sigma}{d\Omega}\right)_{\nu}^{\text{ERL}} &= \frac{G^2 \epsilon^2}{2\pi^2} \frac{4\pi}{2J_i + 1} \left\{ \cos^2 \frac{\theta}{2} \left[ \sum_{J=0}^{\infty} \left| \langle J_f || \hat{\mathfrak{M}}_J - \frac{q_0}{|\mathbf{q}|} \hat{\mathfrak{L}}_J || J_i \rangle \right|^2 \right] \right. \\ &\quad + \left. \left[ \frac{q^2}{2q^2} \cos^2 \frac{\theta}{2} + \sin^2 \frac{\theta}{2} \right] \right. \\ &\quad \times \left. \left[ \sum_{J \geq 1}^{\infty} \left( \left| \langle J_f || \hat{\mathfrak{J}}_J^{\text{mag}} || J_i \rangle \right|^2 + \left| \langle J_f || \hat{\mathfrak{J}}_J^{\text{el}} || J_i \rangle \right|^2 \right) \right] \right. \\ &\quad \mp \sin \frac{\theta}{2} \frac{1}{|\mathbf{q}|} \left( q^2 \cos^2 \frac{\theta}{2} + q^2 \sin^2 \frac{\theta}{2} \right)^{1/2} \\ &\quad \times \left. \left[ \sum_{J \geq 1}^{\infty} 2 \operatorname{Re} \langle J_f || \hat{\mathfrak{J}}_J^{\text{mag}} || J_i \rangle \langle J_f || \hat{\mathfrak{J}}_J^{\text{el}} || J_i \rangle^* \right] \right\} \end{aligned}$$

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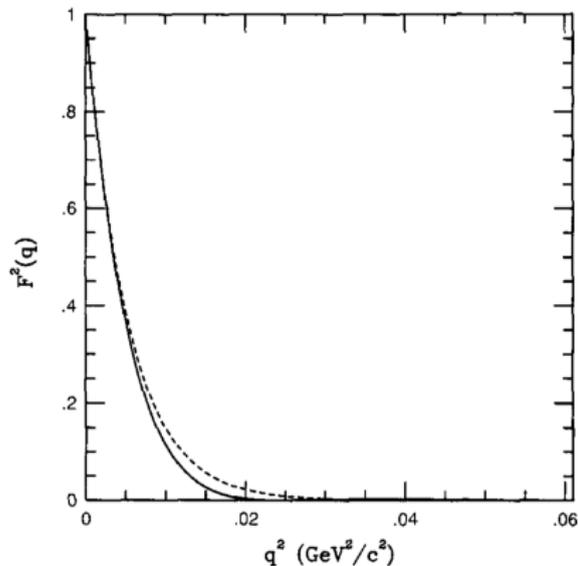
Doesn't contribute when  $J_i = J_f = 0$ .

But calculable in reasonable nuclear models.

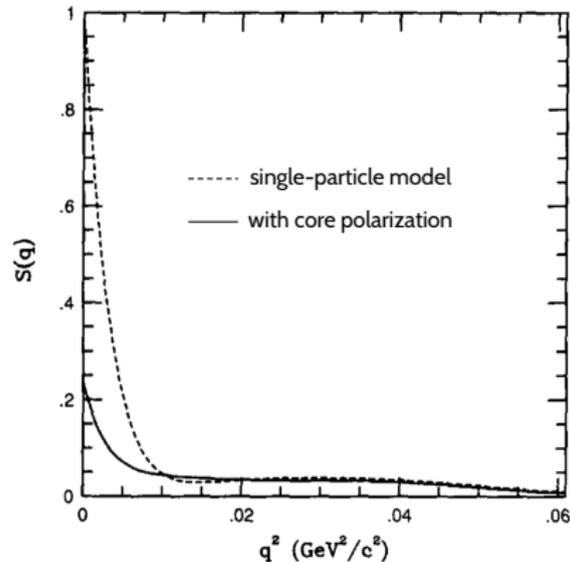
$O(1/A)$  in elastic scattering.

# WIMP-Nucleus Elastic Scattering

$\tilde{B}$  on  $^{131}\text{Xe}$ : a similar process



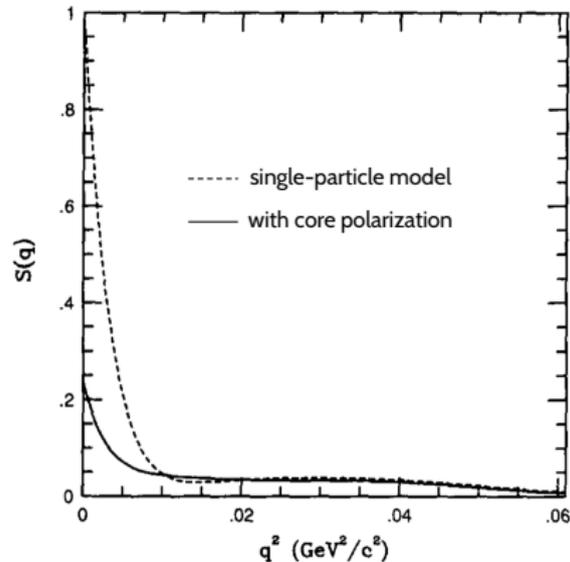
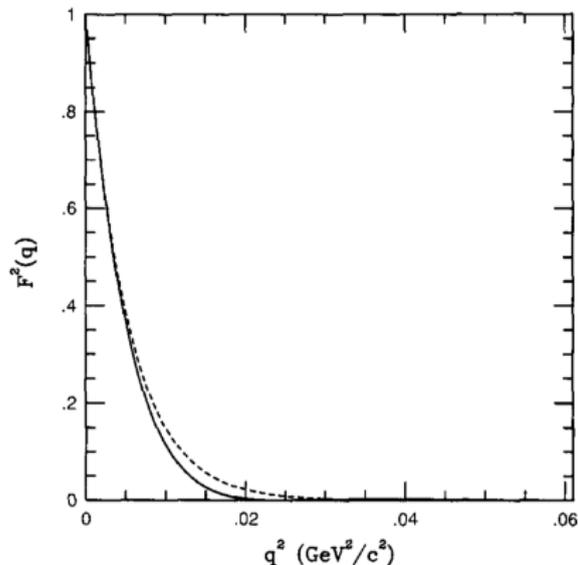
Spin-Independent (Coherent) Form Factor



Spin-Dependent Form Factor

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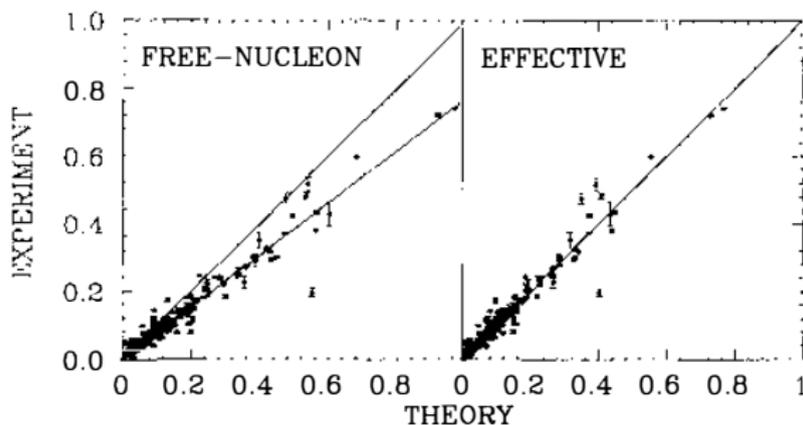
Spin-Dependent Form Factor

Spin-dependent scattering down by  $1/A$  but has long tail.

# $g_A$ “Renormalization:” Gamow-Teller $\beta$ Decay

Leading order decay operator is  $\vec{\sigma}\tau_+$ .

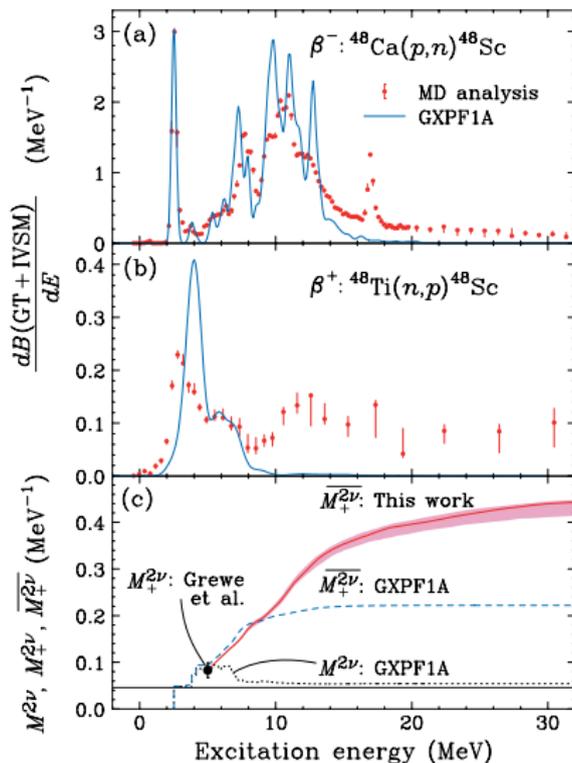
**40-Year-Old Problem:** Effective  $g_A$  needed in all calculations of shell-model type.



Lots of suggestion about the cause but no consensus until quite recently.

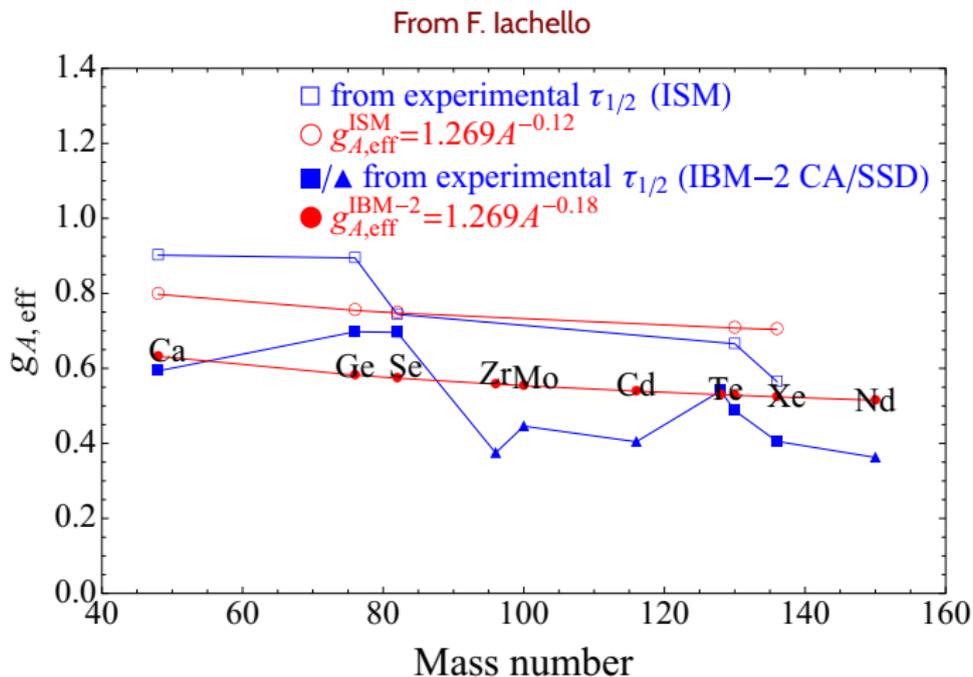
# Other Tests of $\vec{\sigma}\tau$ Strength Also Show Suppression

From Yako et al., PRL 103, 012503 (2009)



Only about 2/3 of theoretically expected strength observed.

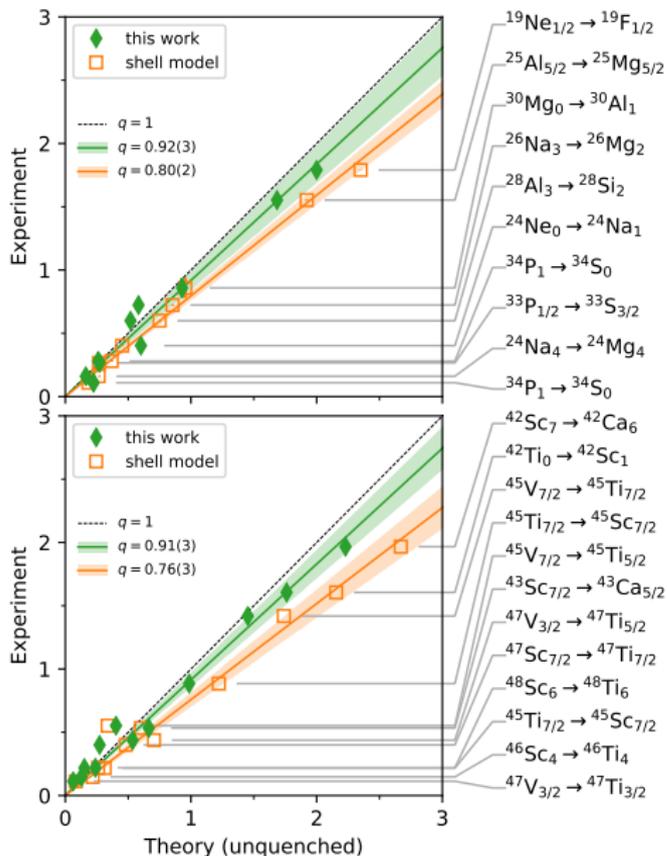
# And $2\nu\beta\beta$ Decay...



If quenching is this severe in  $0\nu$  decay, experimentalists will not be happy.

# Ab-Initio Calculations in the $sd$ -Shell

## Explains Most of Quenching



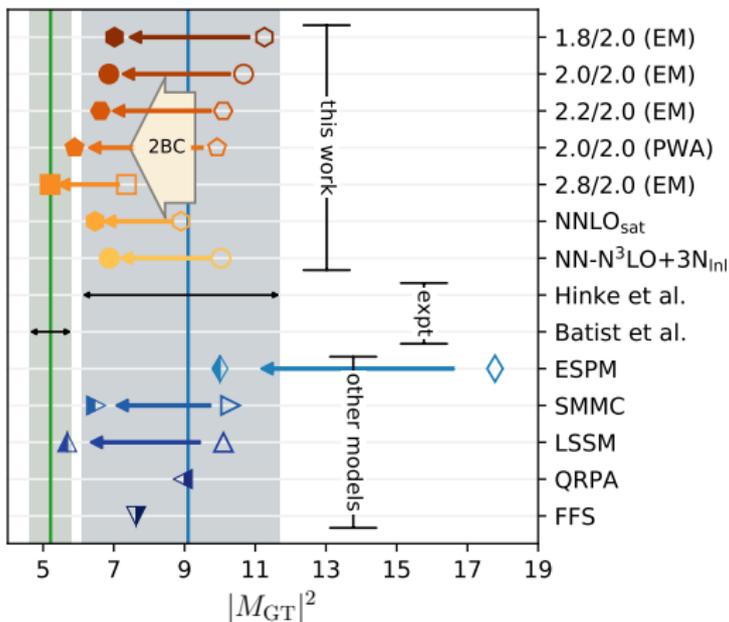
Recent ab-initio work, Holt et al,  
preliminary

Shell model seems to include some correlations. Quenching comes from additional correlations and “meson-exchange currents” (roughly speaking).

# ...And in $^{100}\text{Sn}$

Another Ab-Initio Calculation

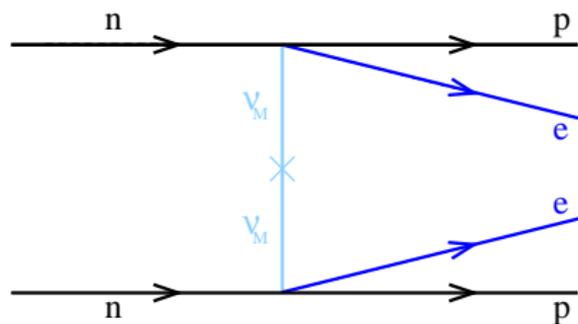
Hagen et al, unpublished



Again, most of the quenching from correlations, meson-exchange currents.

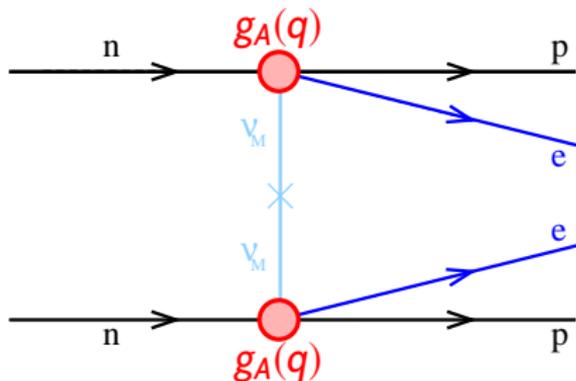
Quenching increases with mass.

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Intermediate neutrino has 100 or so MeV of momentum on average, so would like to know  $g_A(q)$  at  $q \approx 100$  MeV.

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The End. Thanks.