

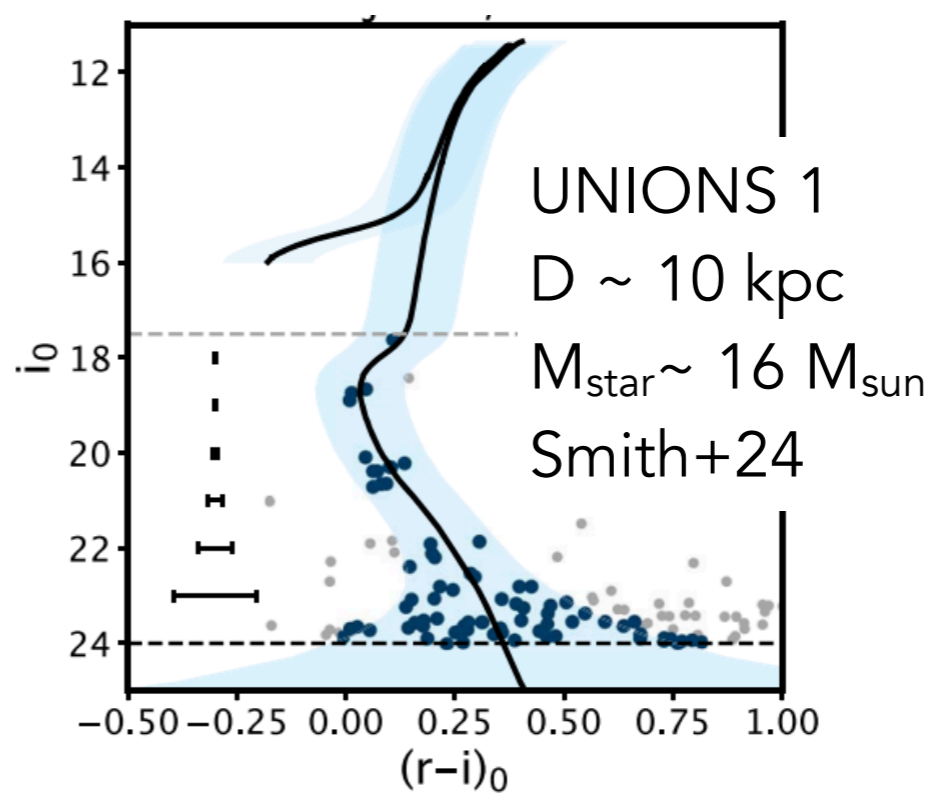
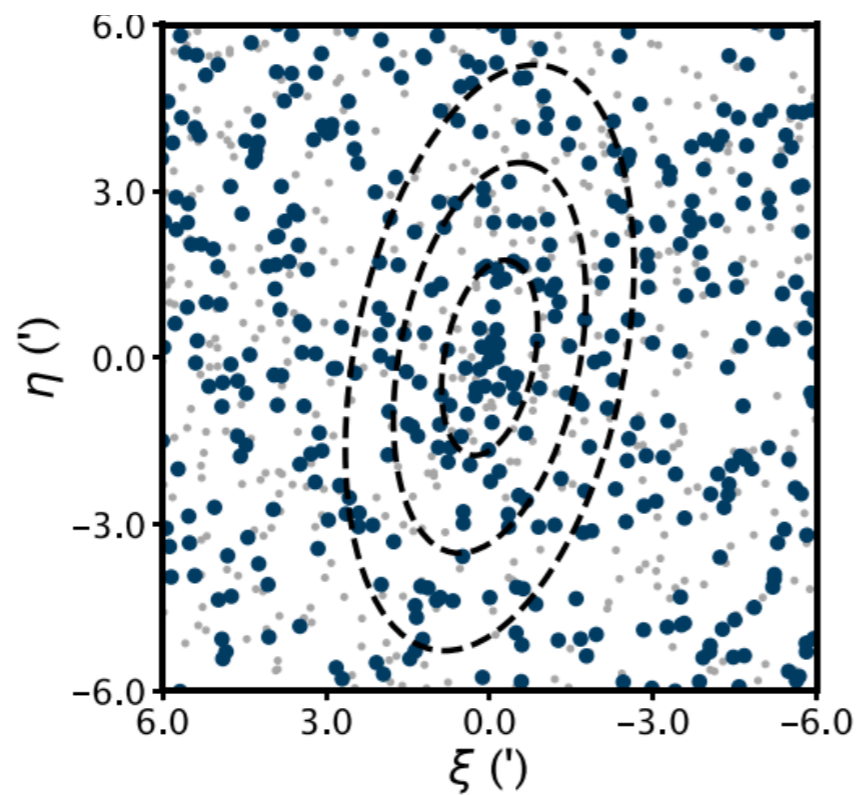
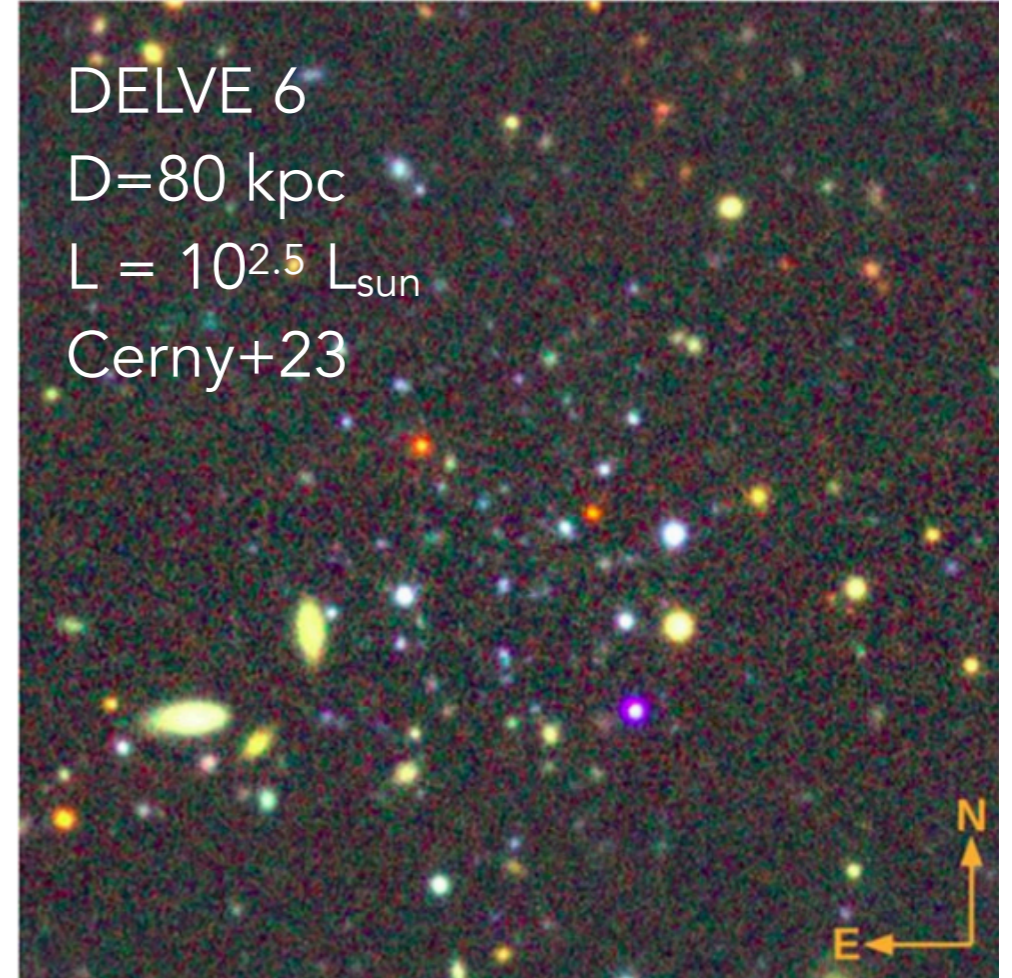
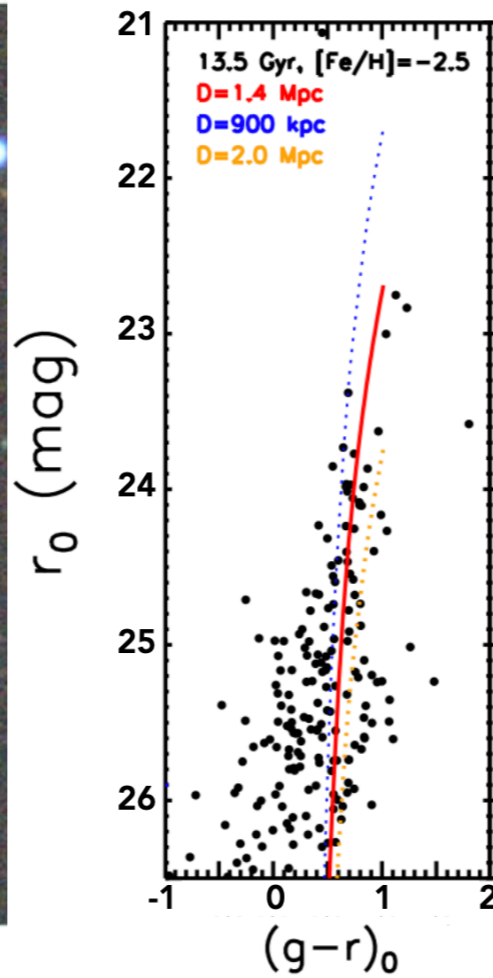
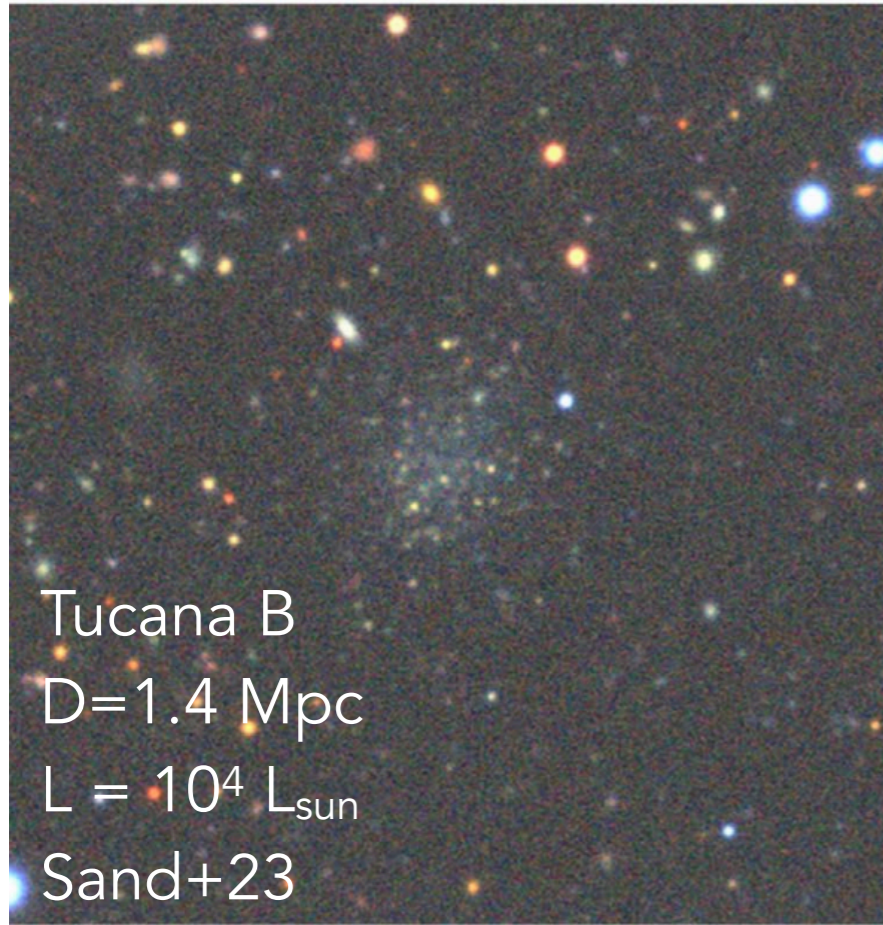


Detailed Views Into the Baryon
Cycle of Dwarf Galaxies from HST
Narrowband Imaging

Sal Wanying Fu
UC Berkeley

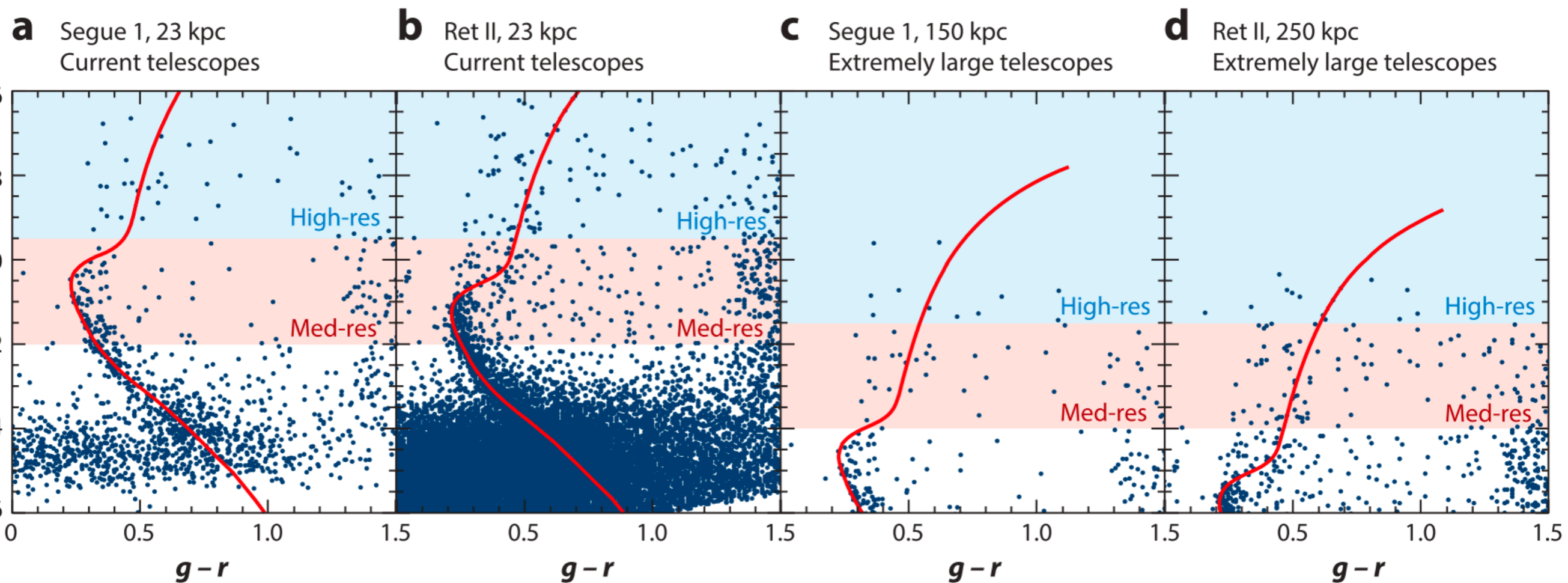
credit: NOIR Lab

Dwarf Discoveries Enabled by Photometric Surveys



Difficulty of Measuring Stellar Metallicities in Distant and/or Faint Galaxies

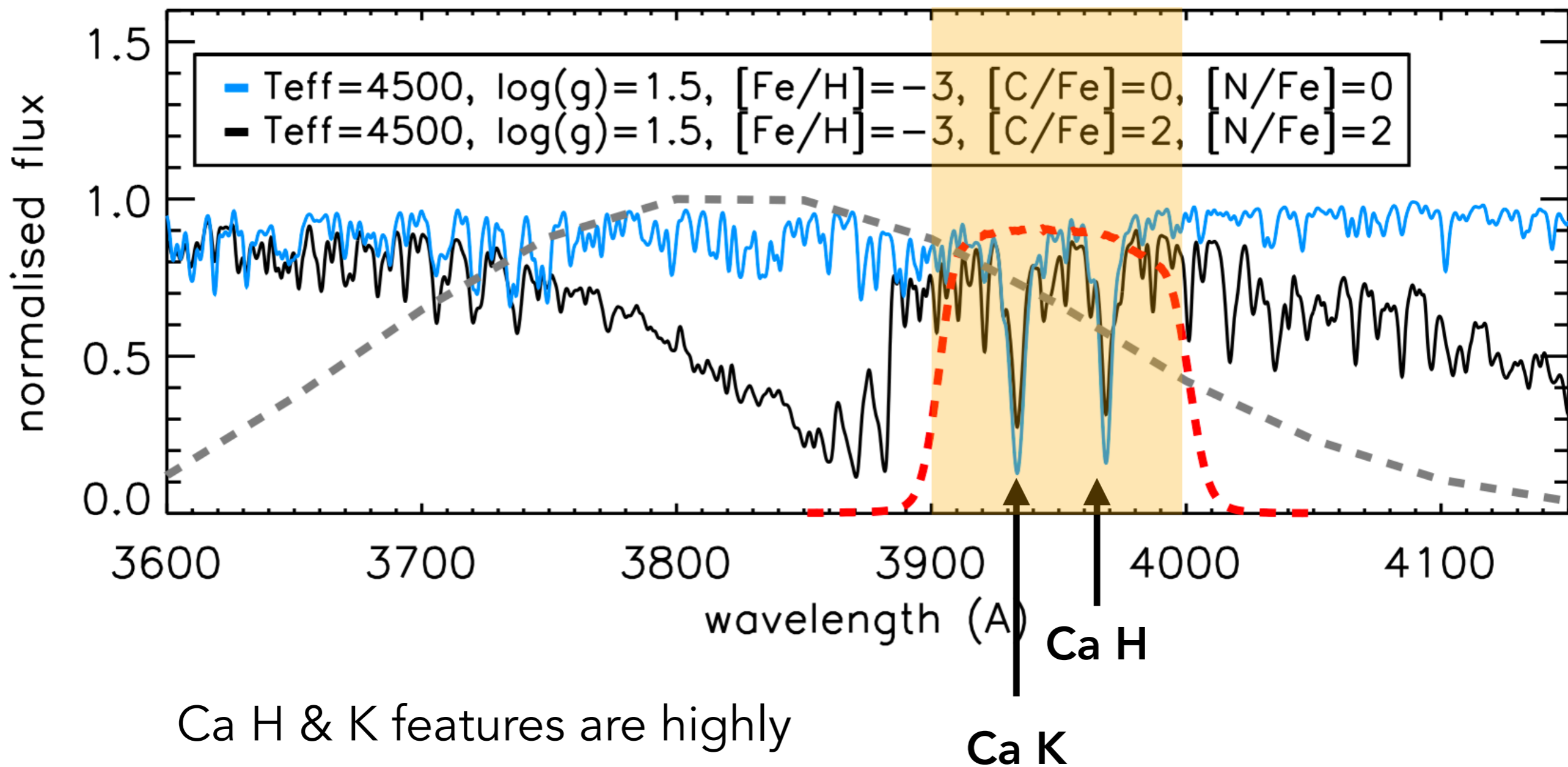
- Only handful of spectroscopically accessible stars for radial velocity, metallicity, chemical abundance measurements
- Problem in era of accelerated dwarf galaxy discovery: Methods for efficiently sampling MDFs?



observing capacity for current spectrographs
of currently-known UFDs

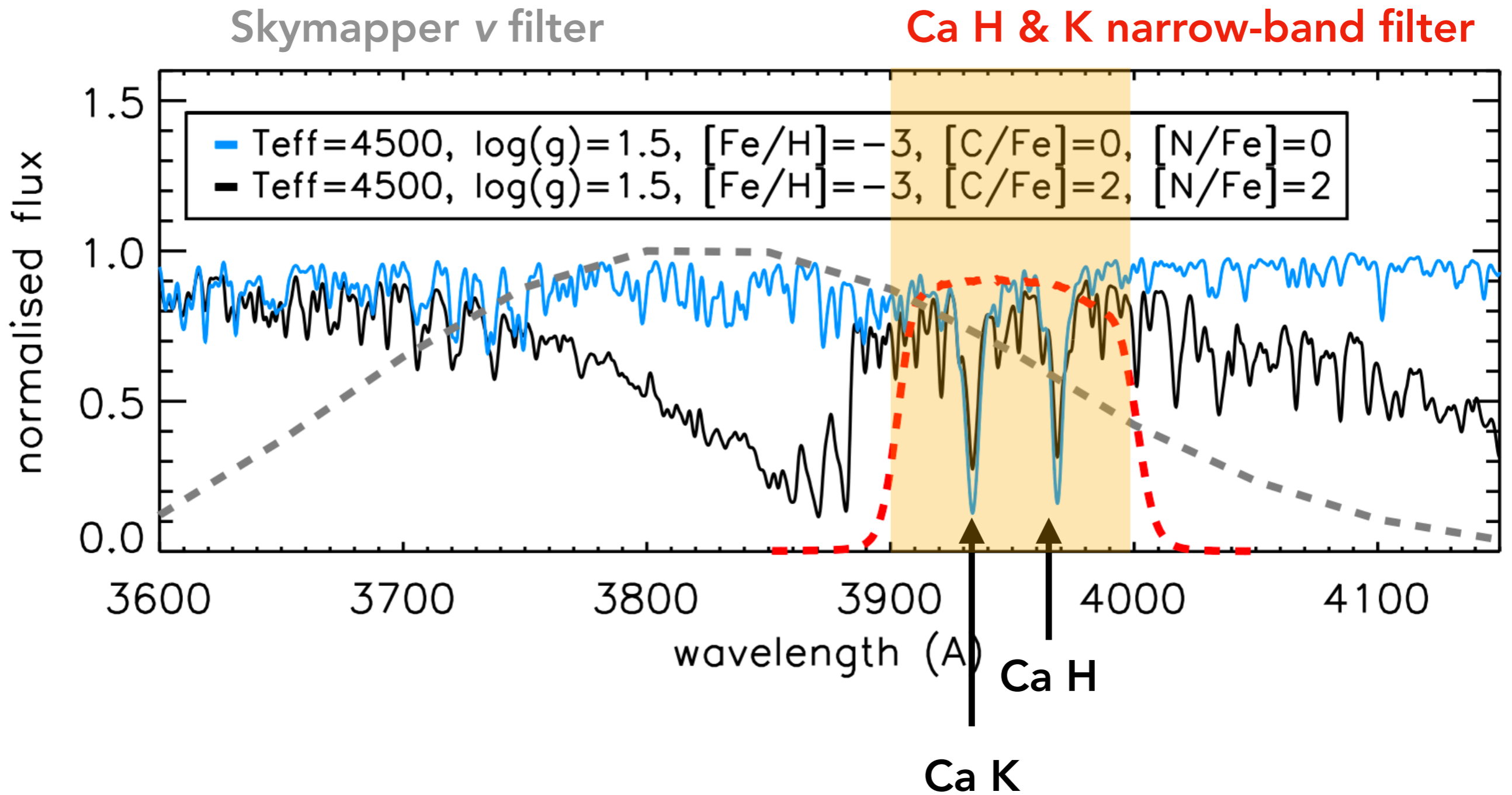
projected observing capacity for ELT
spectrographs for anticipated UFD discoveries

Photometric Metallicities Using Ca H & K



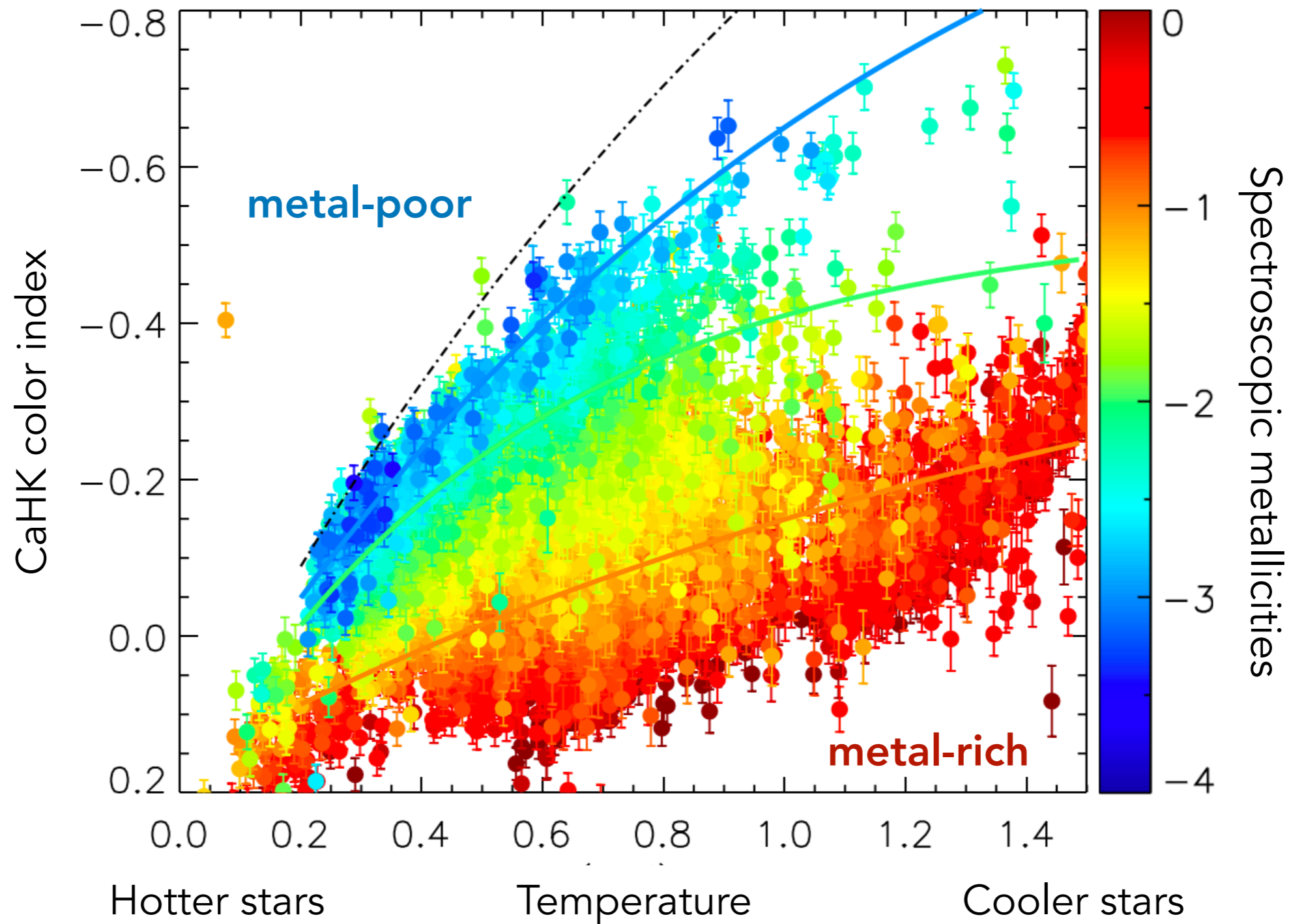
Ca H & K features are highly sensitive to $[\text{Fe}/\text{H}]$ (e.g., Strömberg 1966)

Photometric Metallicities Using Ca H & K



Narrow-band photometry around Ca H & K can trace stellar metallicity

Applying photometric metallicities to Milky Way surveys



Stars in Ca H & K color space separate in metallicity, especially cooler stars

Narrow-band photometric metallicities of faint stars in distant galaxies using HST

GO 15901

Wed May 17 20:30:27 GMT 2023

Principal Investigator: Daniel Weisz

PI Institution: University of California - Berkeley

Investigators ([xml](#))

Title: The Metallicity Distribution Functions of Ultra-Faint Dwarf Galaxies

Cycle: 27

Allocation: 43 orbits

GO 16226

Fri Jun 24 00:13:38 GMT 2022

Principal Investigator: Sal Fu

PI Institution: University of California - Berkeley

Investigators ([xml](#))

Title: Metallicity Distribution Functions of Quenched Field Dwarf Galaxies

Cycle: 28

Allocation: 23 orbits

GO 16686

Fri Jun 24 00:14:40 GMT 2022

Principal Investigator: Daniel Weisz

PI Institution: University of California - Berkeley

Investigators ([xml](#))

Title: The Metallicity Distribution Functions of Faint M31 Satellites

Cycle: 29

Allocation: 30 orbits

> 1 dex Discrepancy Between Observed and Simulated Metallicities for Galaxies Below $10^4 L_{\text{sun}}$

$$\text{Log}M(M_{\odot}), (M/L)_{\odot} = 2$$

3

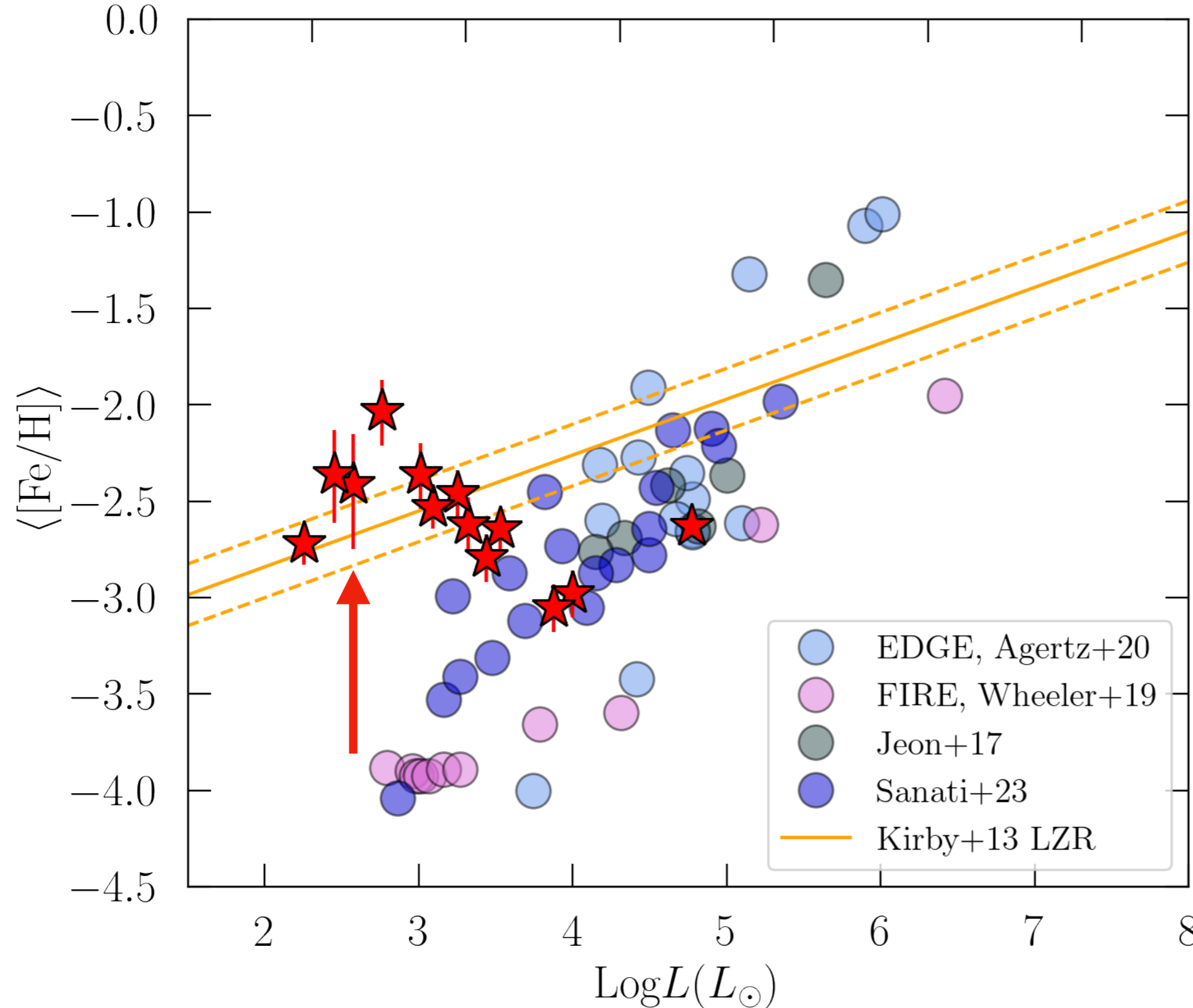
4

5

6

7

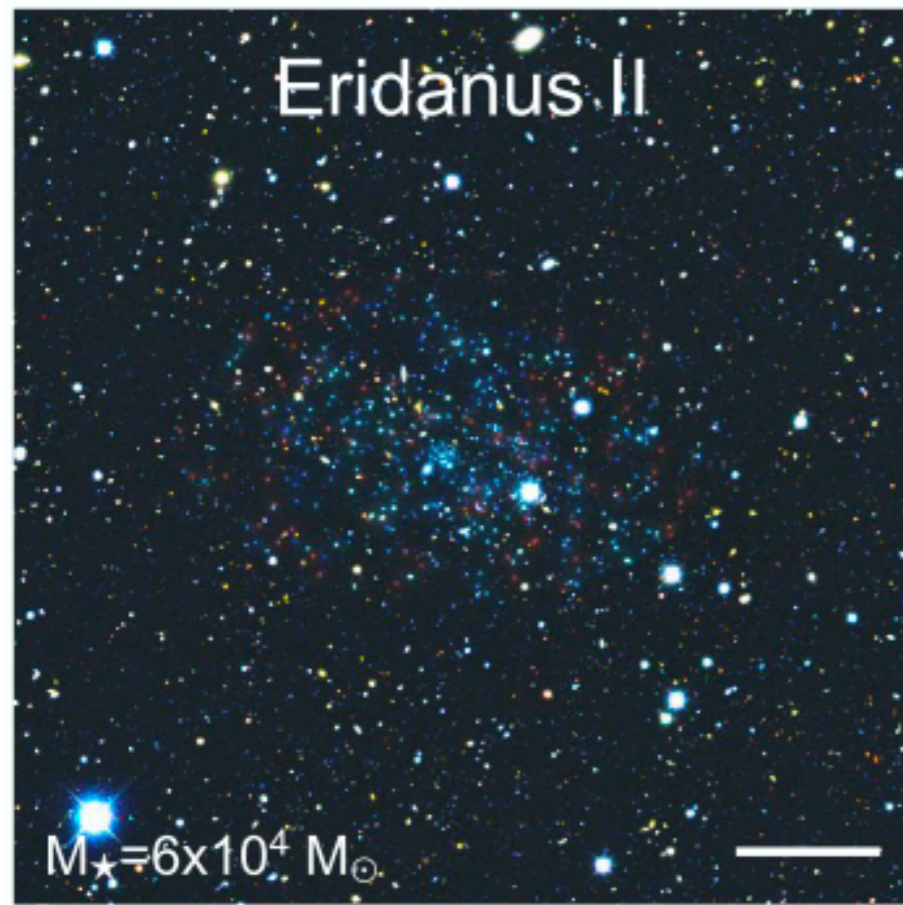
8



Implementation of simulation physics:

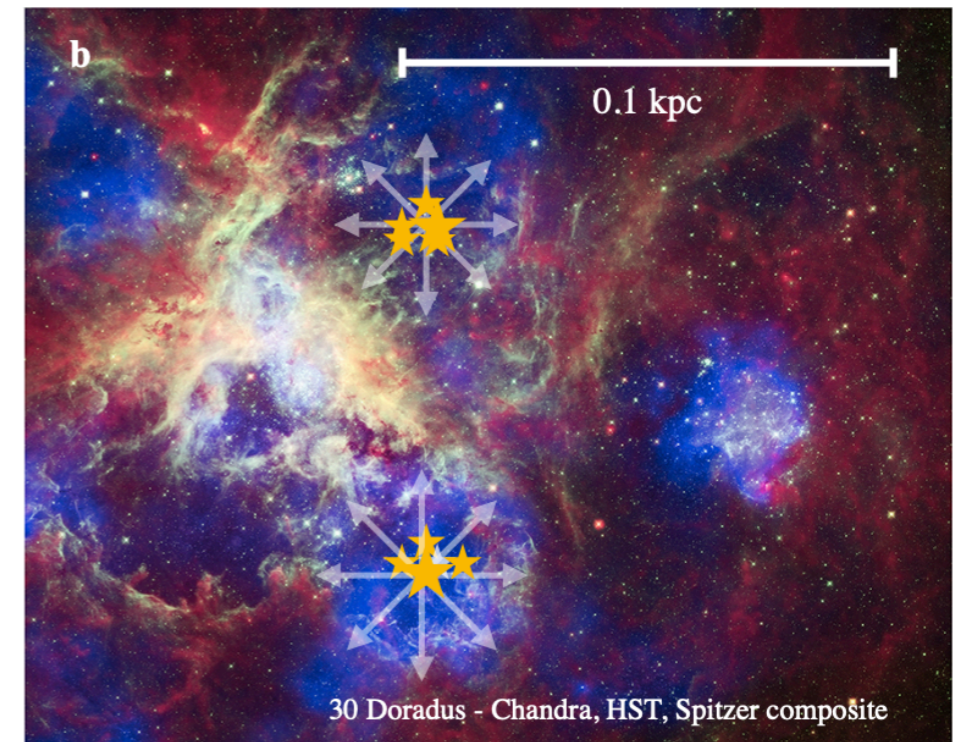
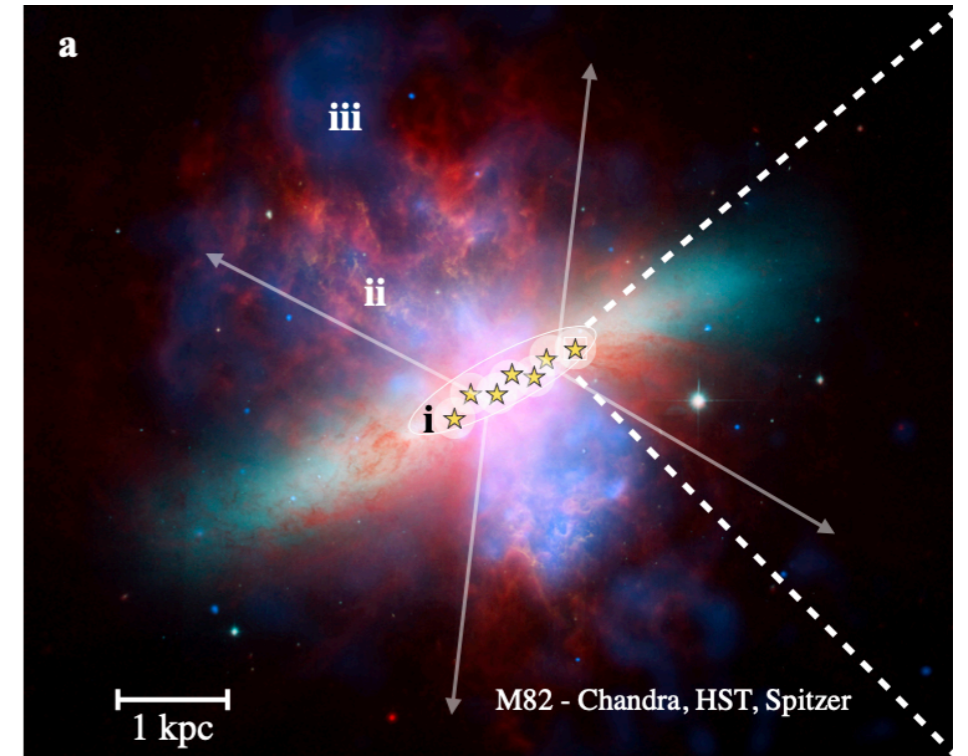
- SNe feedback energy (Agertz+2020)
- Metal diffusion (Escala+2018)
- IMF sampling (Prgomet+2022)
- [Fe/H]-dependent Type Ia SNe (Gandhi+2022)
- Pop III stars (Sanati+2023)
- SNe implementation (Azartash-Namin+2024, Zhang+2024)
- Simulation resolution, pre-enrichment of natal gas, ...

Chemical Evolution as an Archaeological Lens into Star-Forming Astrophysics



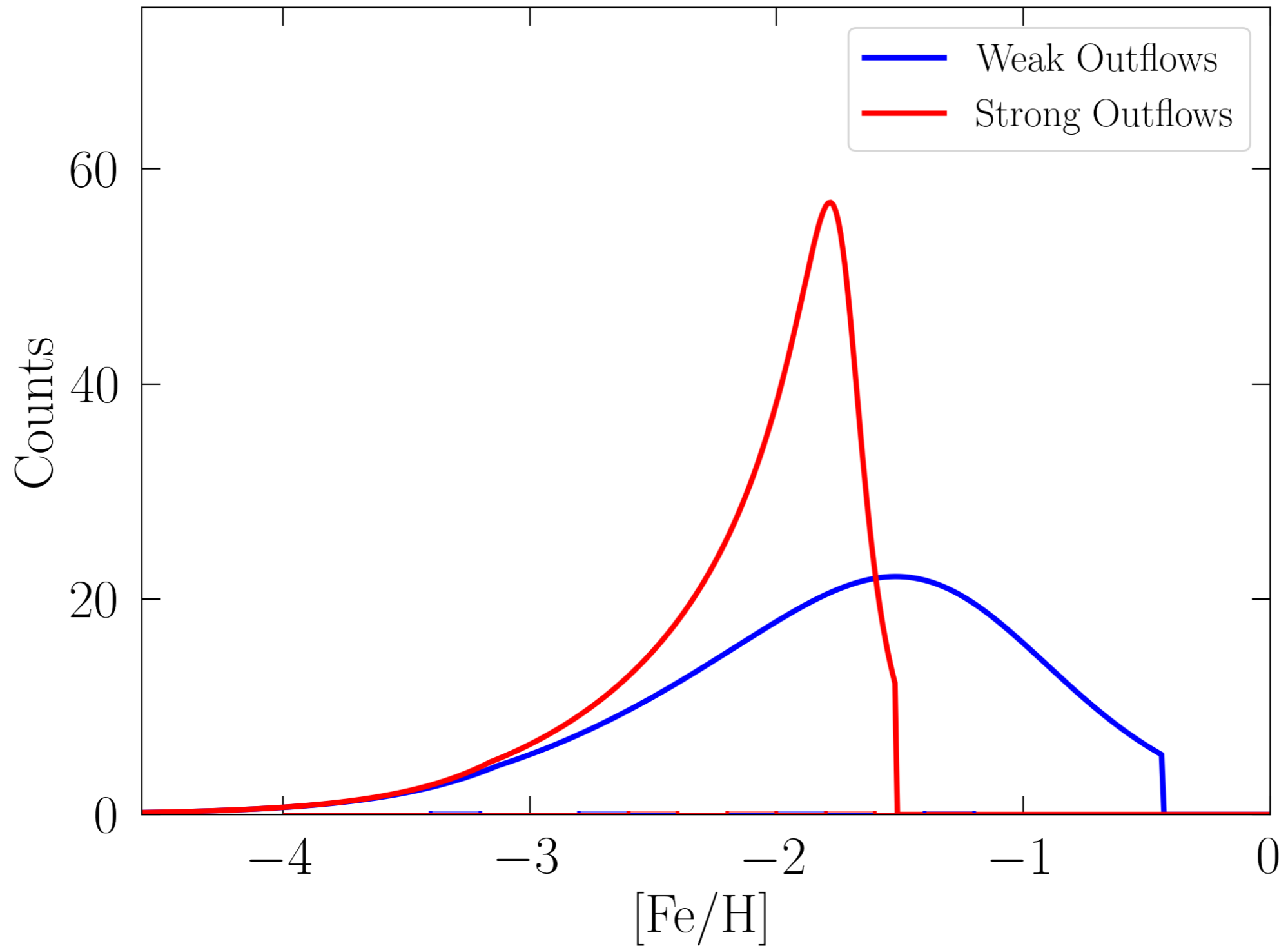
$z = 0$

HST CaHK MDF +
chemical evolution
models
(Sandford+2024)



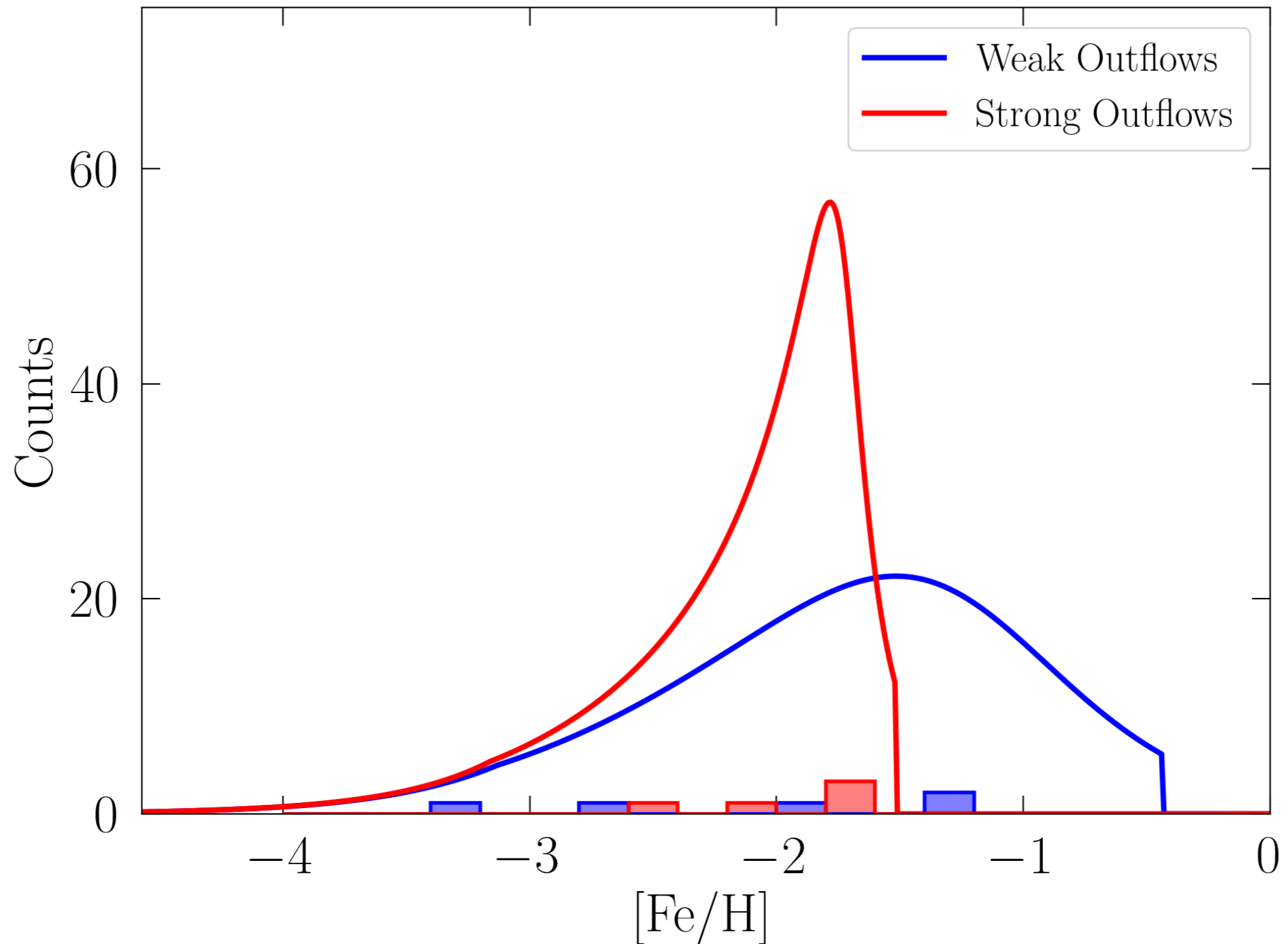
Eri II, $z \sim 7$?

Outflow Strength on MDF Shape



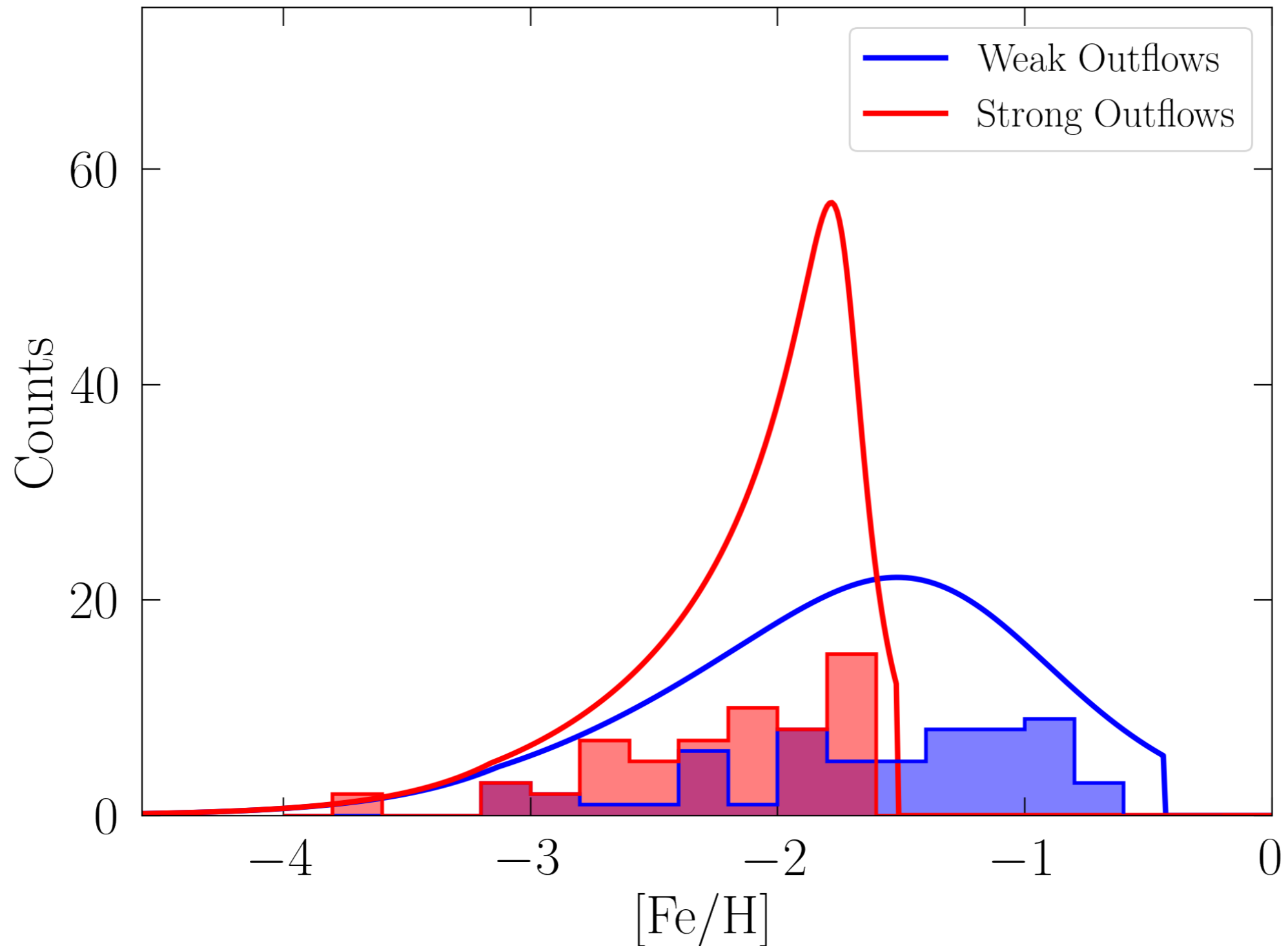
Outflow Strength on MDF Shape

$N = 5$



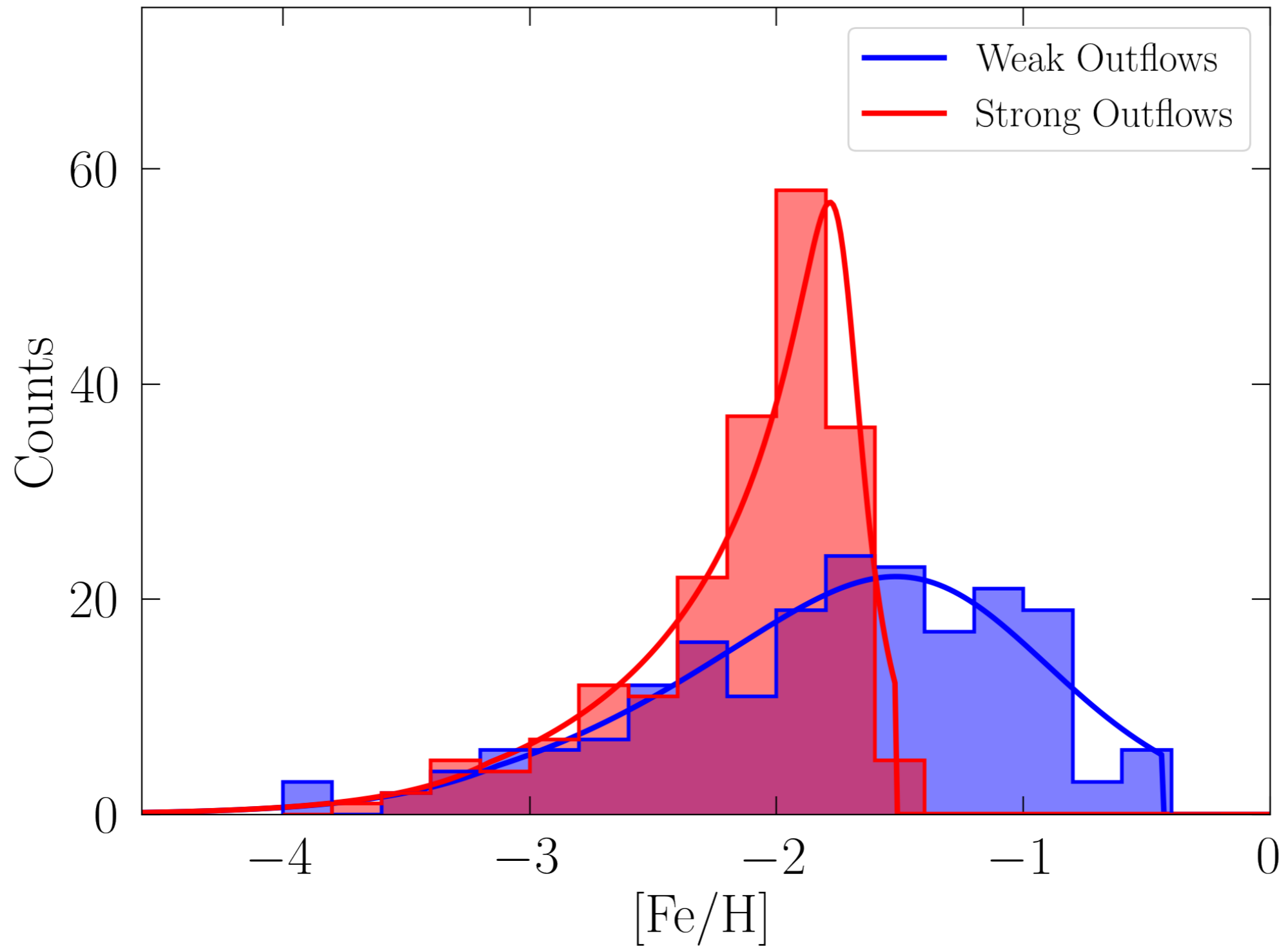
Outflow Strength on MDF Shape

$N = 60$

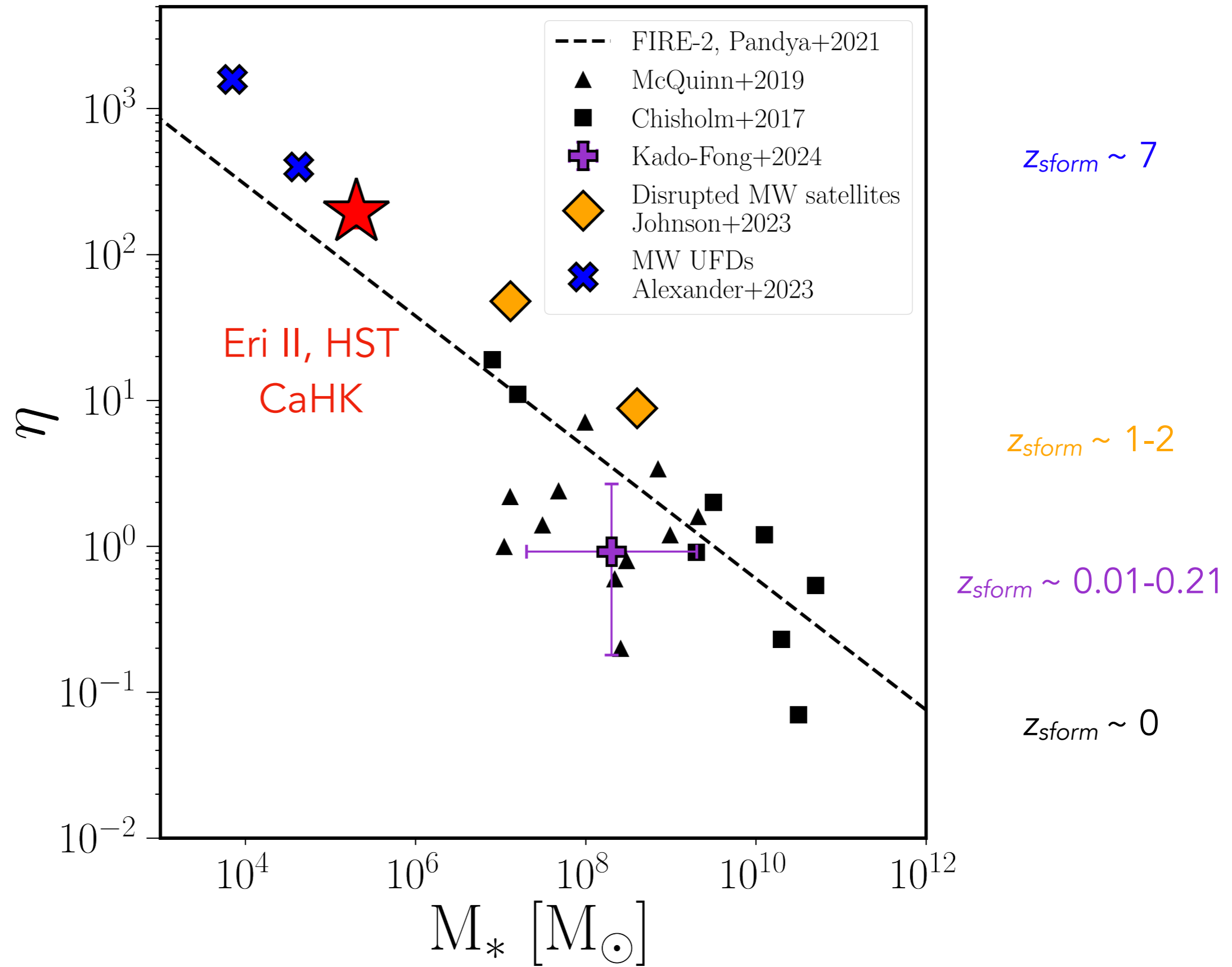


Outflow Strength on MDF Shape

$N = 200$

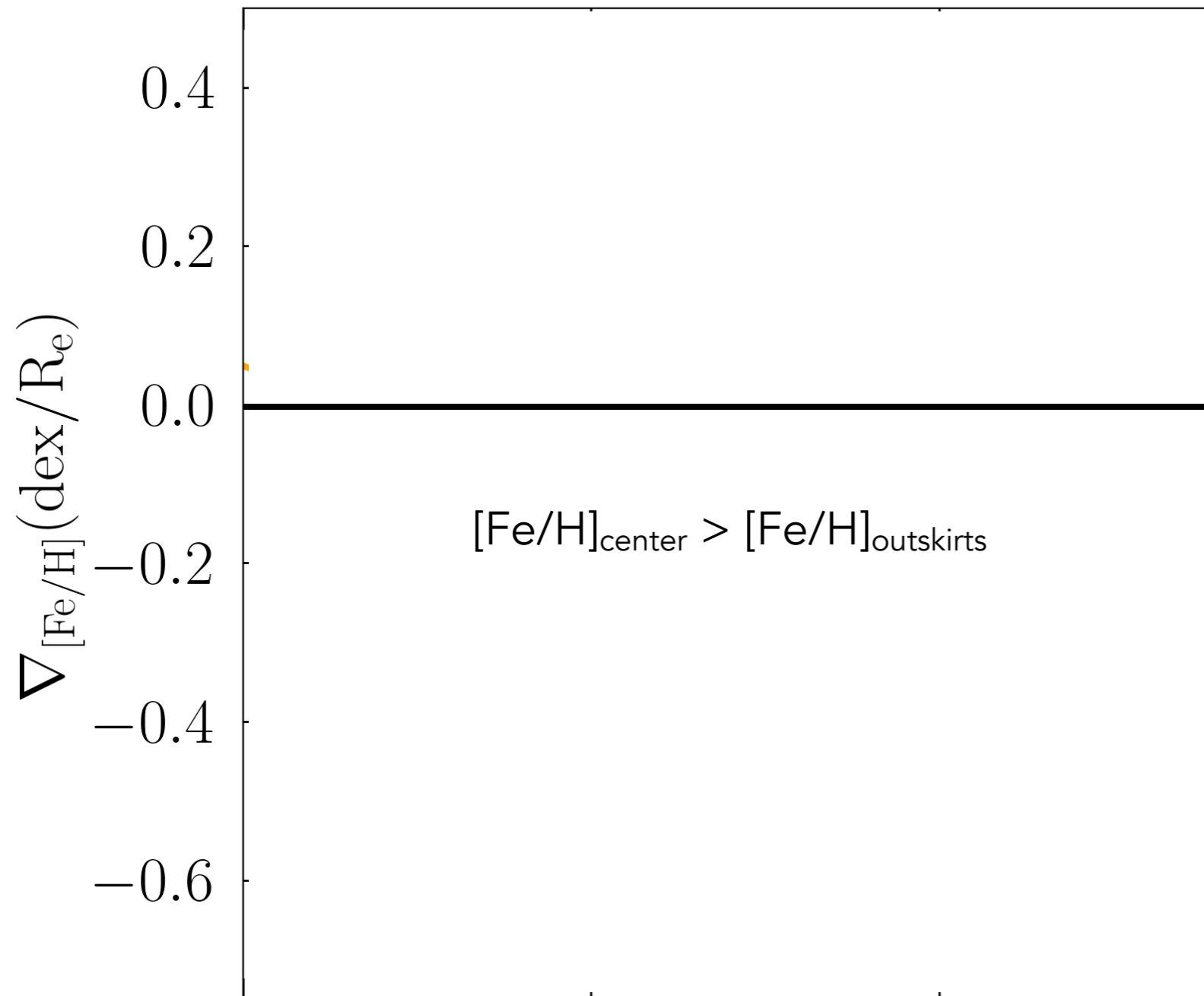


Outflow Strength in Dwarf Galaxies



Adapted from Sandford+2024

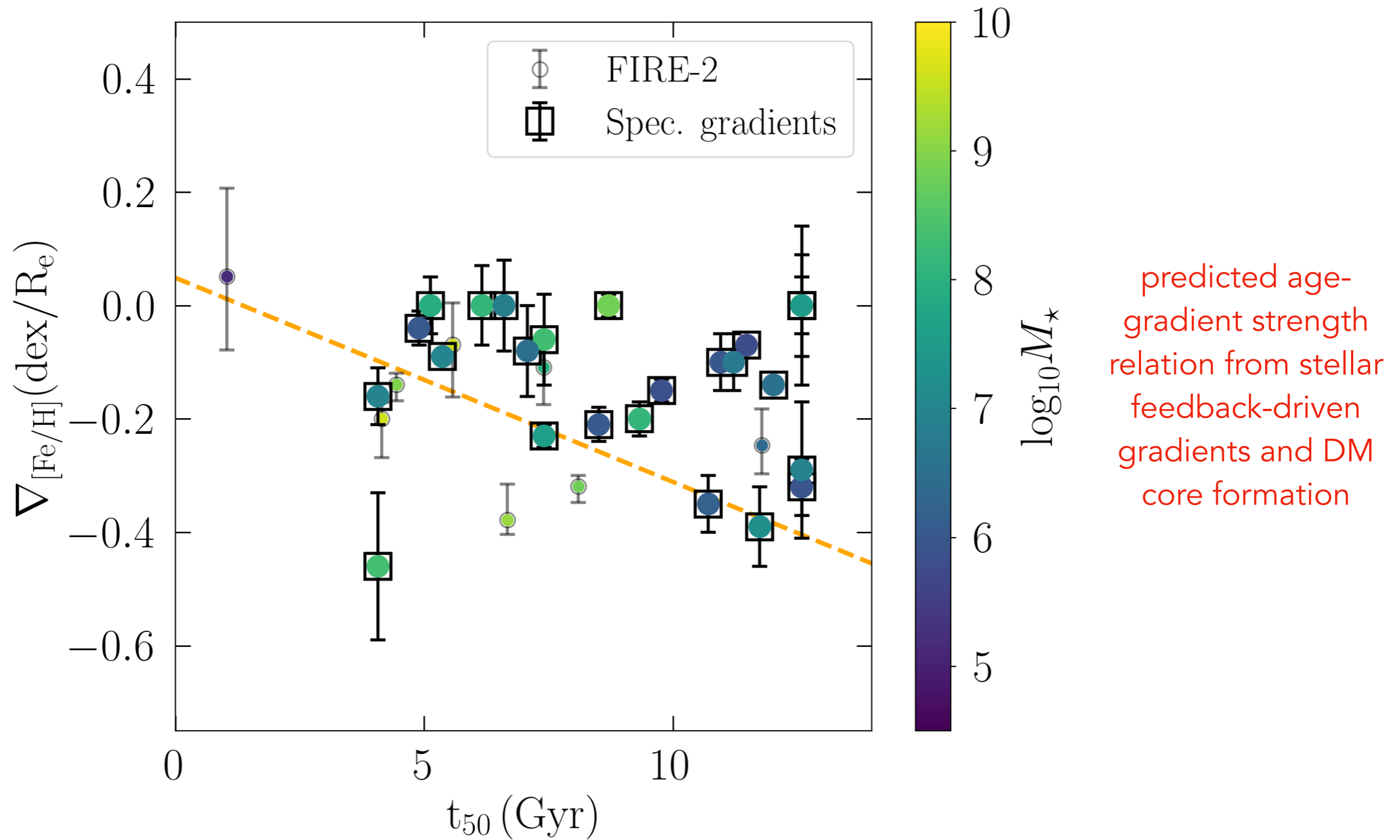
Metallicity Gradients as Tracers of the Baryon Cycle



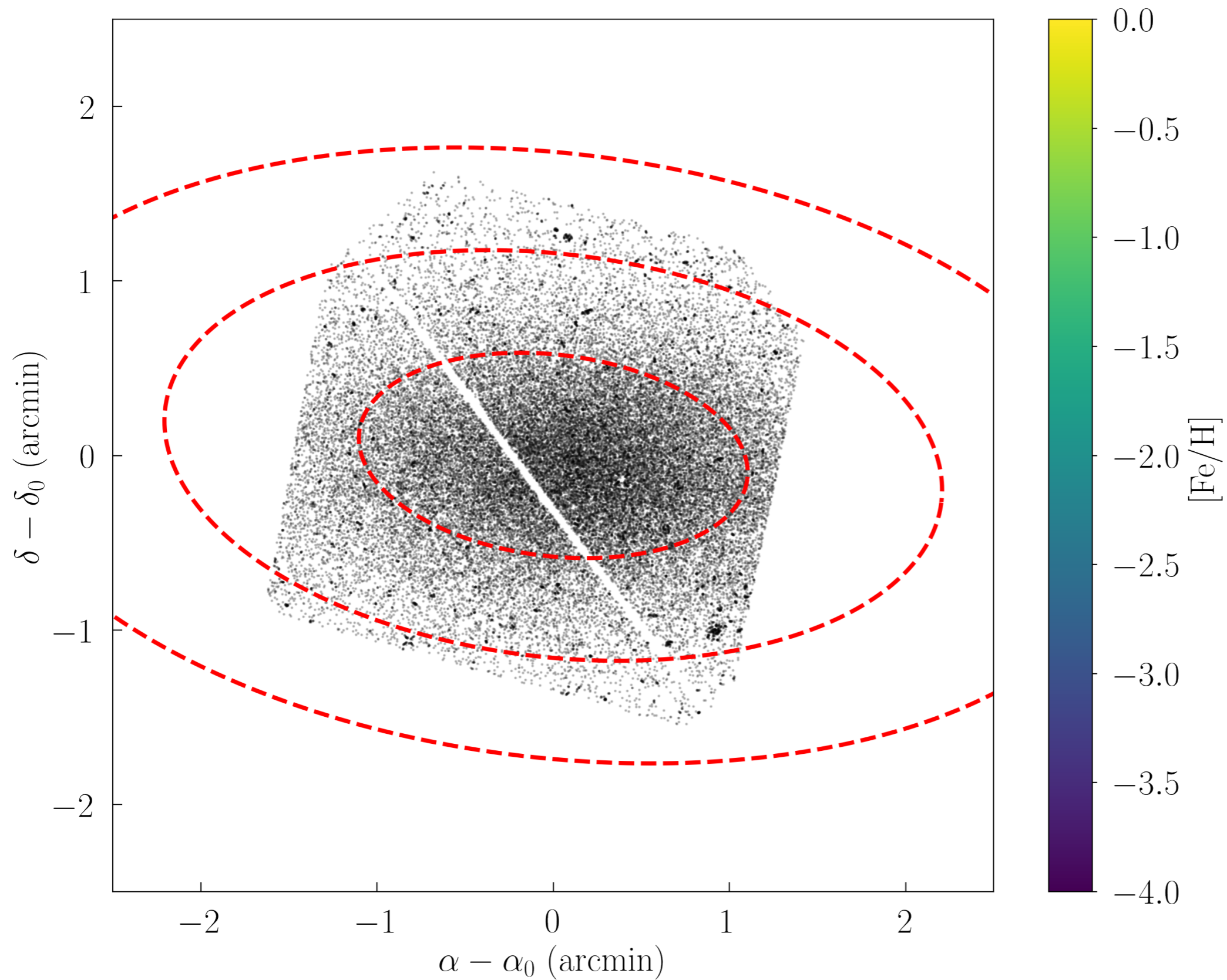
Gradient strength causes?

- M_{star}
- Environment (Sales+2010)
- Angular momentum (Schroyen+2013)
- Merger events (Cardona-Berrero+2021)
- Median stellar age (El-Badry+2016, Mercado+2021)

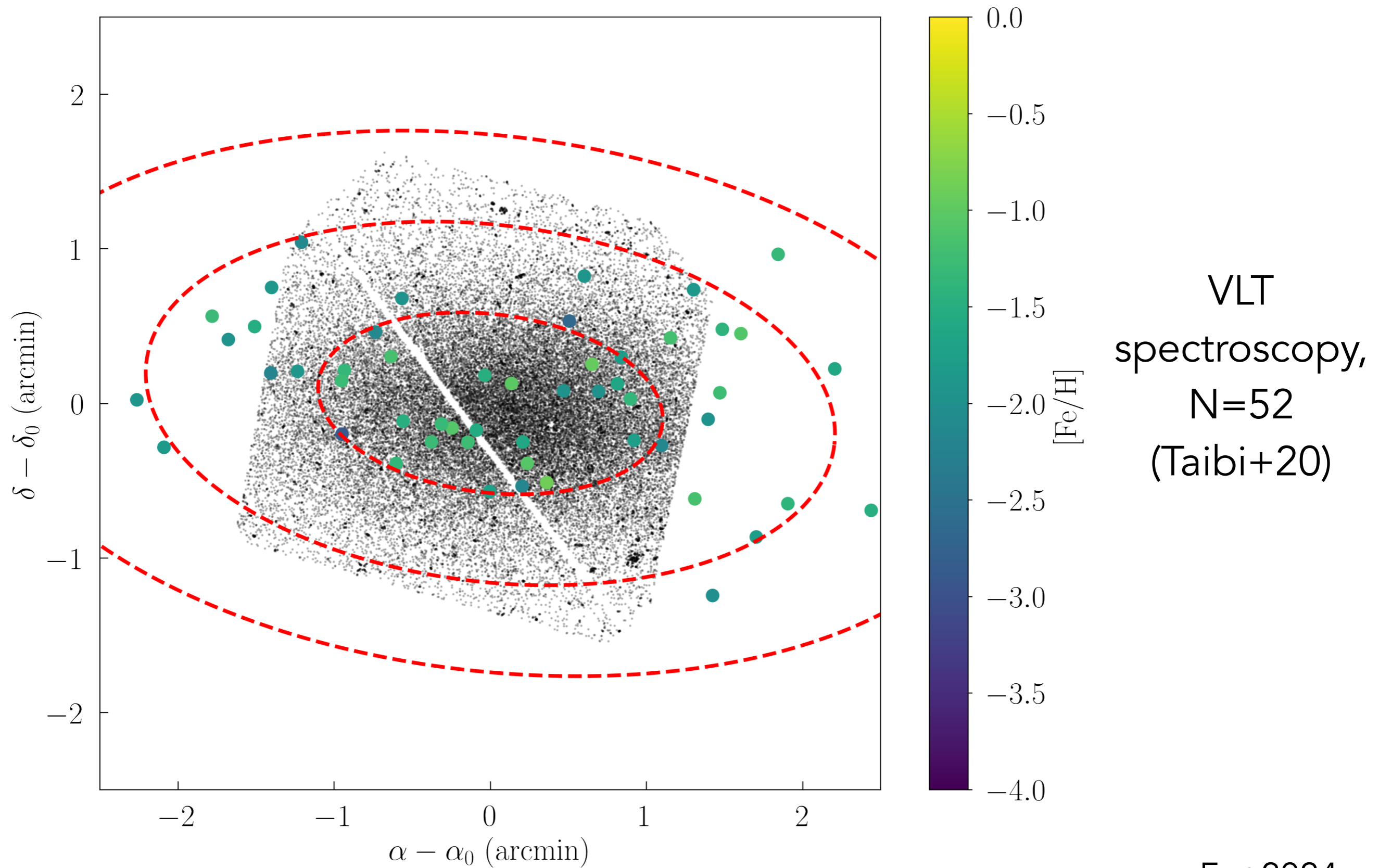
Metallicity Gradients as Tracers of the Baryon Cycle



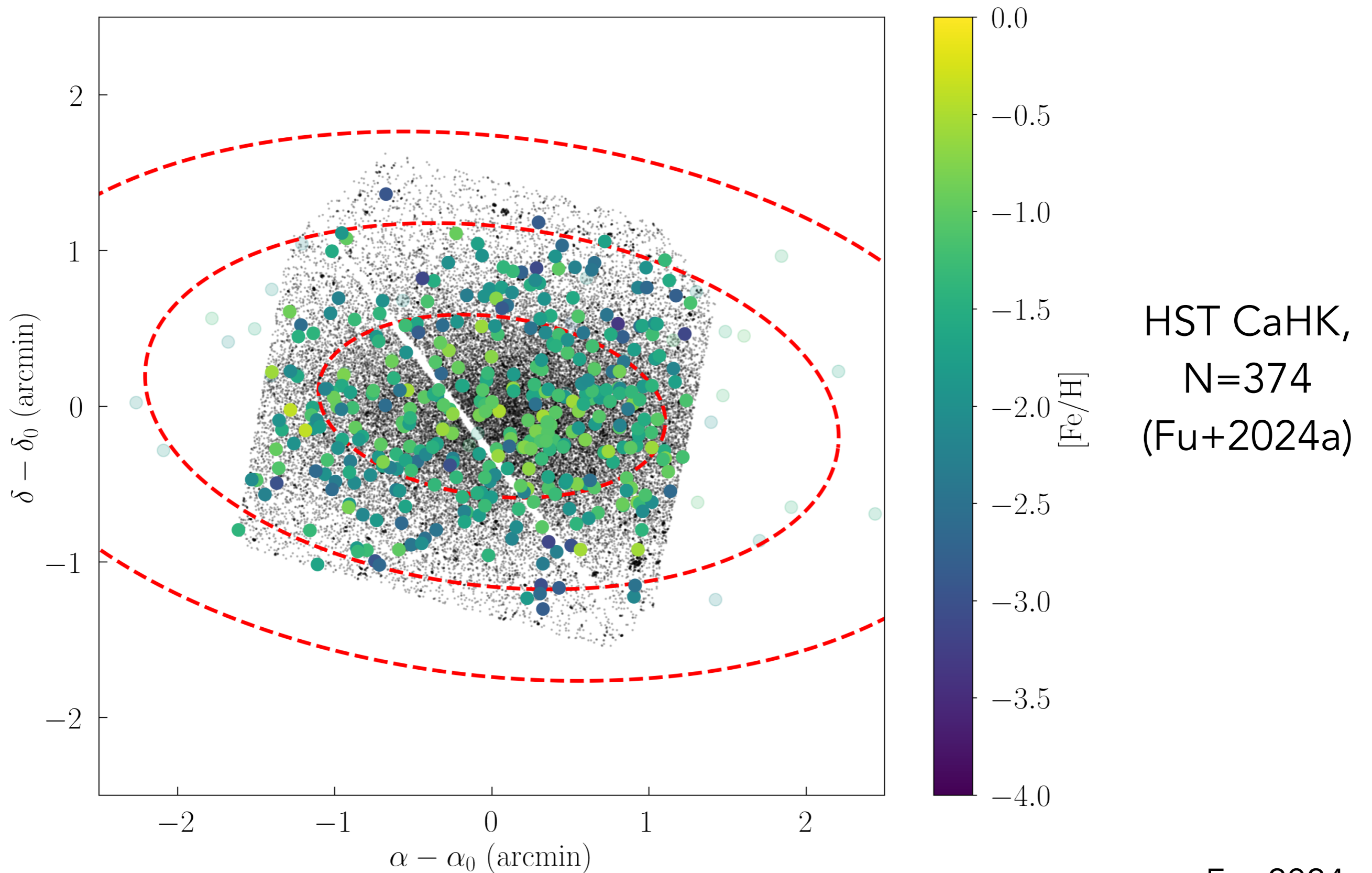
Spatially Resolved Stellar Metallicity Maps of Distant Galaxies: Tucana dSph (~ 1 Mpc)



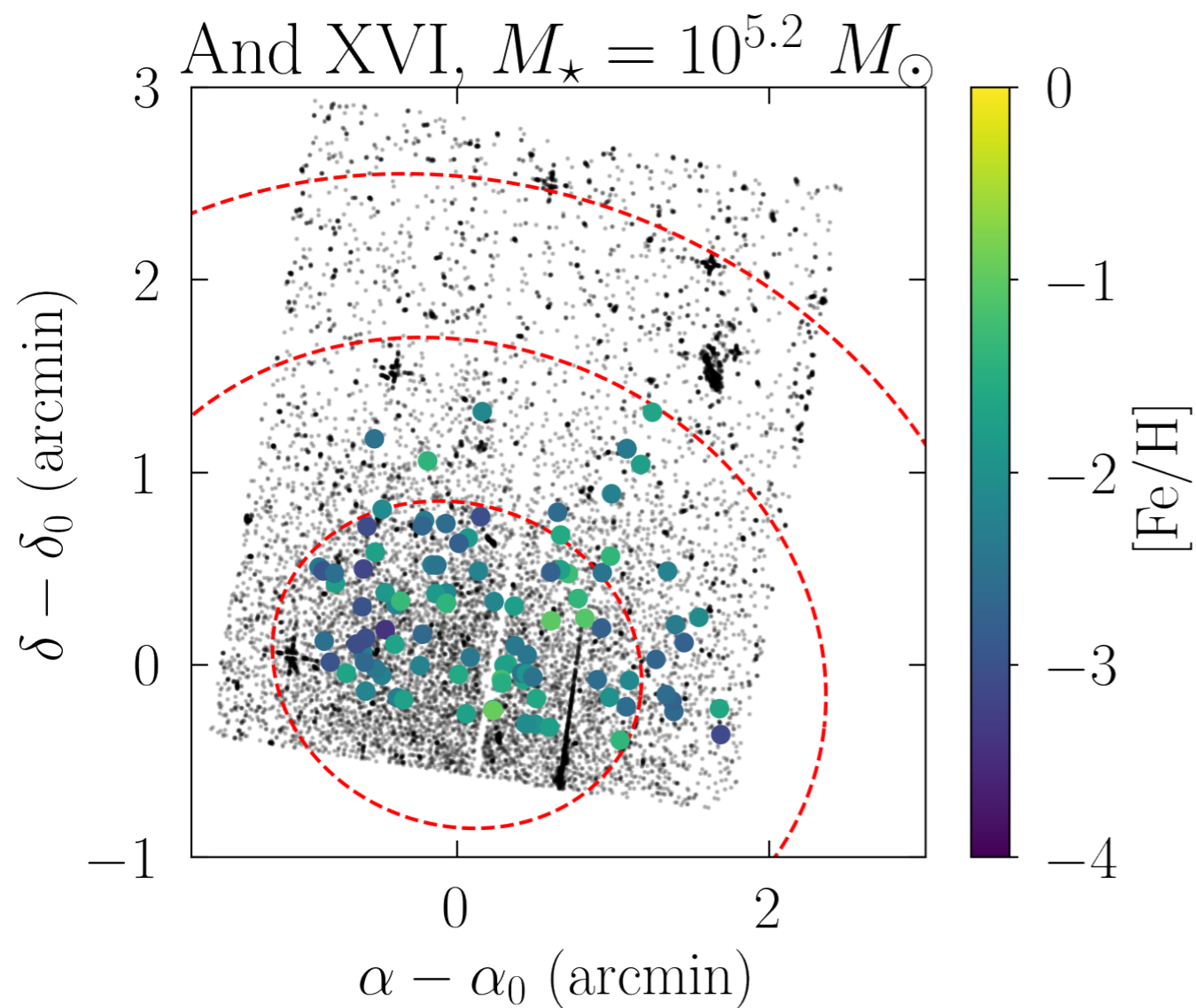
Spatially Resolved Stellar Metallicity Maps of Distant Galaxies: Tucana dSph (~ 1 Mpc)



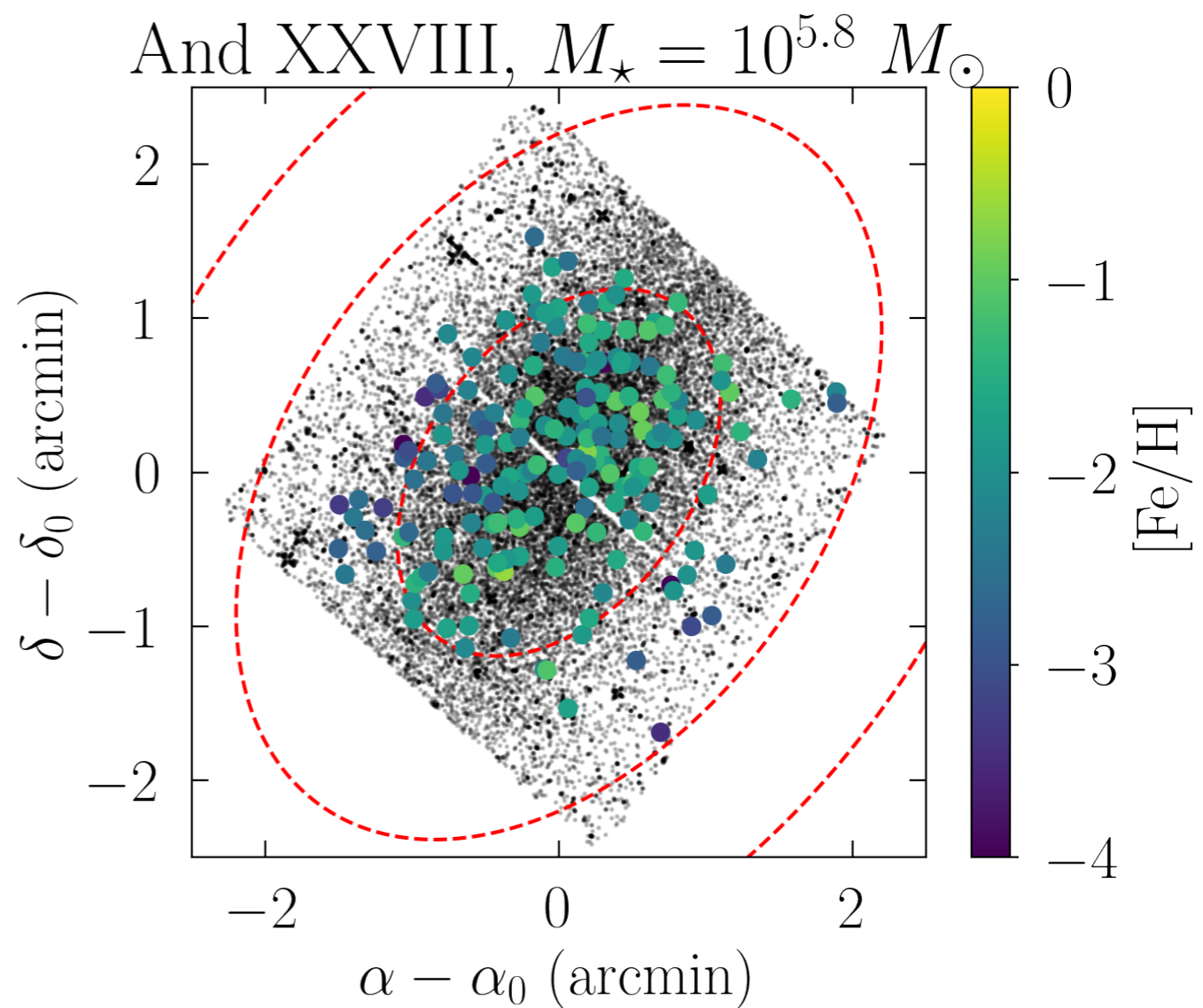
Spatially Resolved Stellar Metallicity Maps of Distant Galaxies: Tucana dSph (~ 1 Mpc)



New on arXiv: Spatially Resolved Stellar Metallicity Maps of M31 Dwarf Galaxies

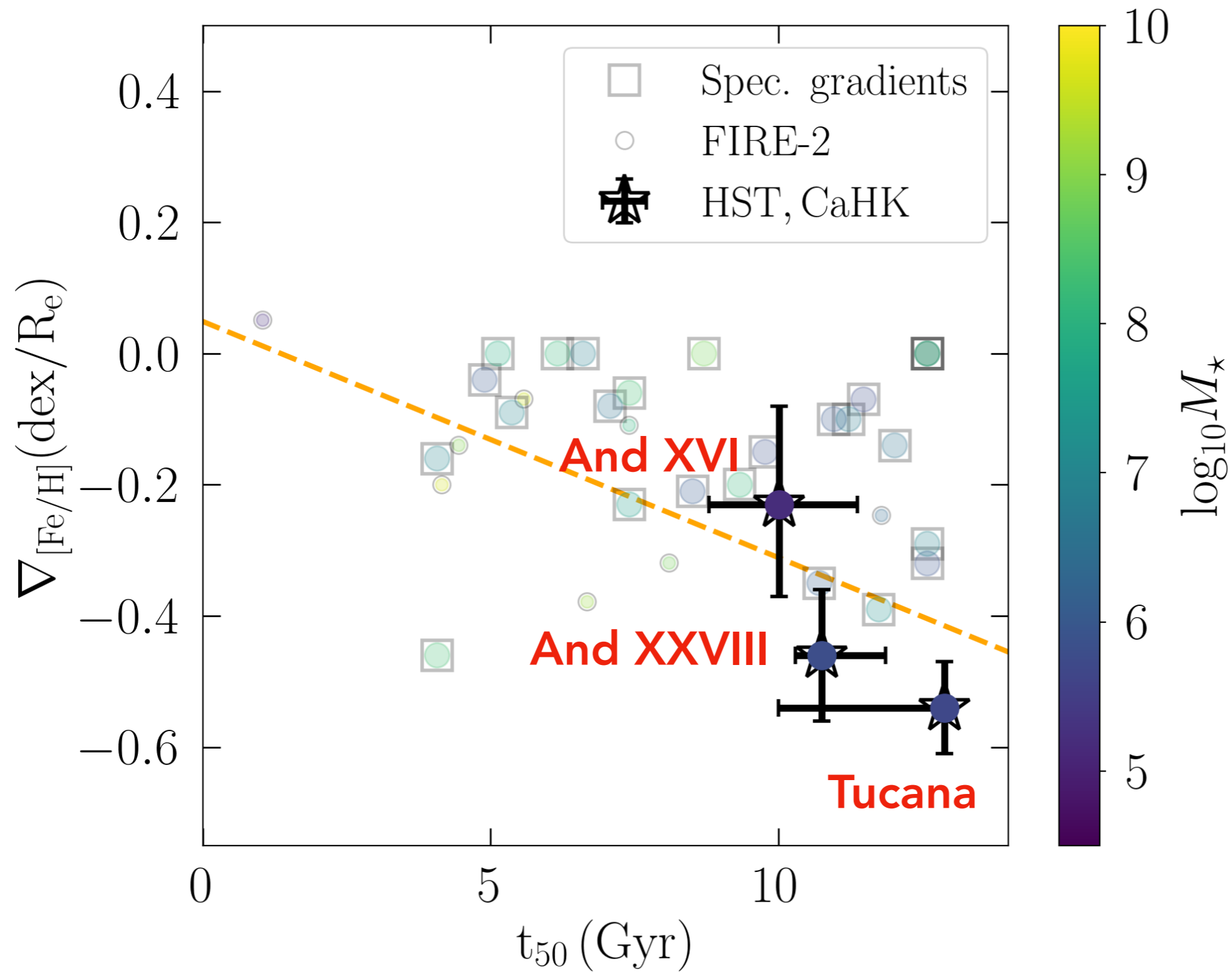


95 stars
($N_{lit} = 0$)

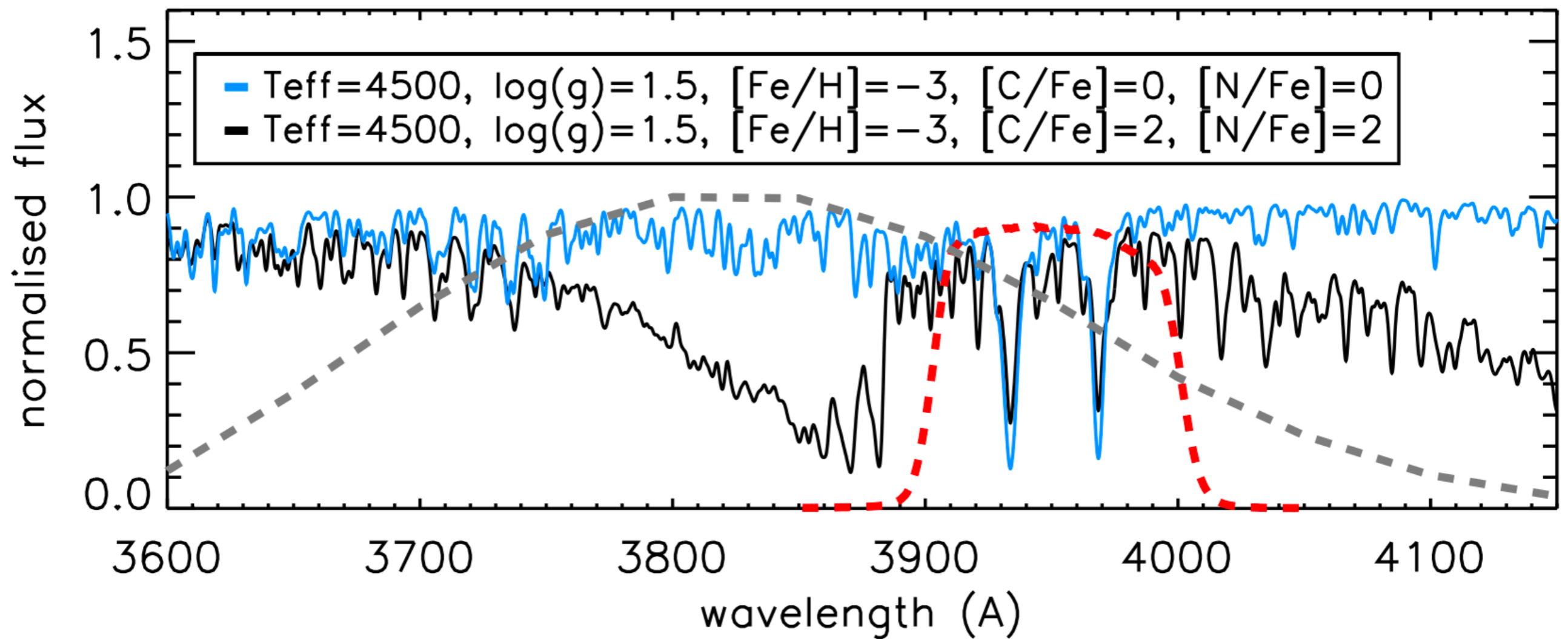


191 stars
($N_{lit} = 13$)

Metallicity Gradients as Tracers of the Baryon Cycle

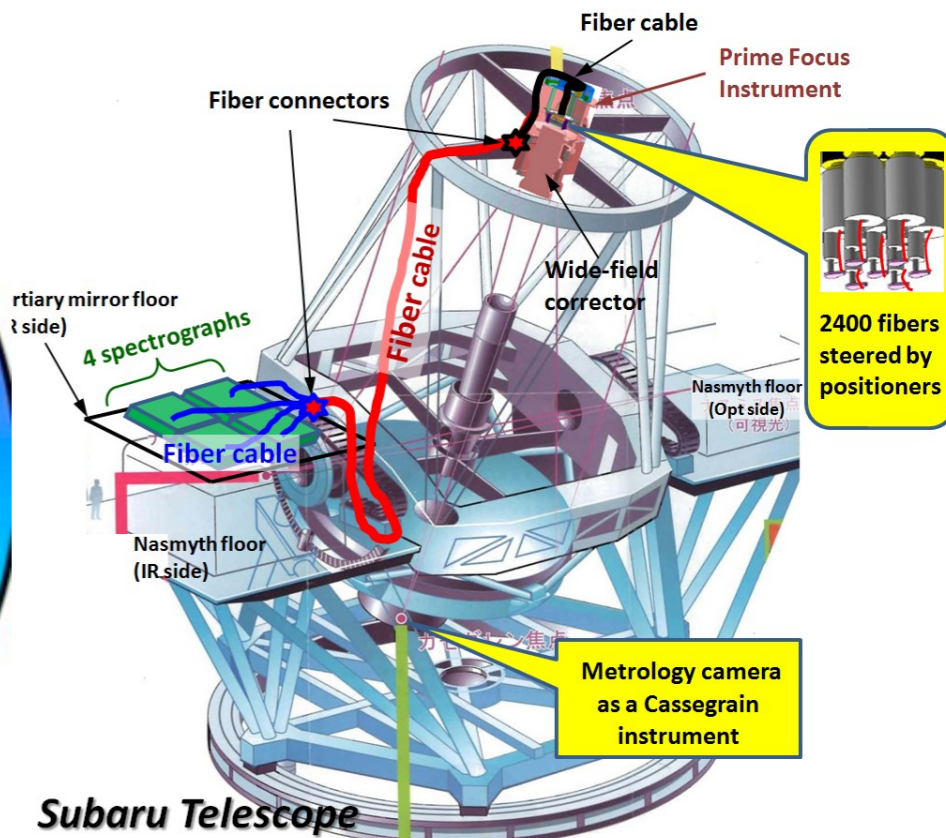
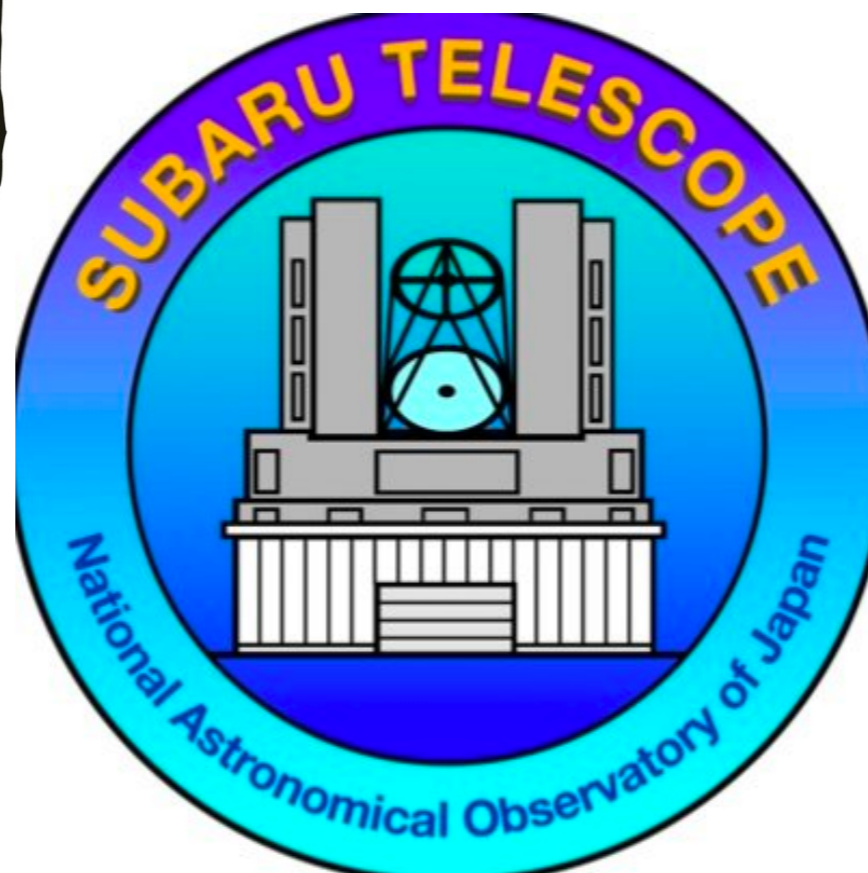
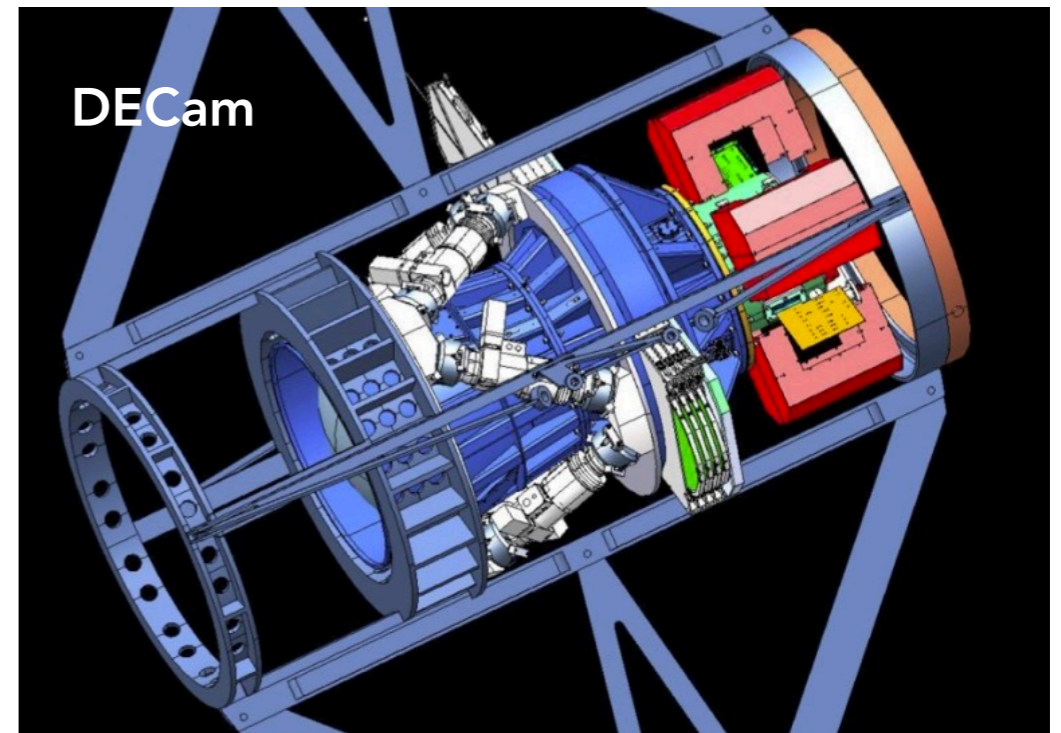
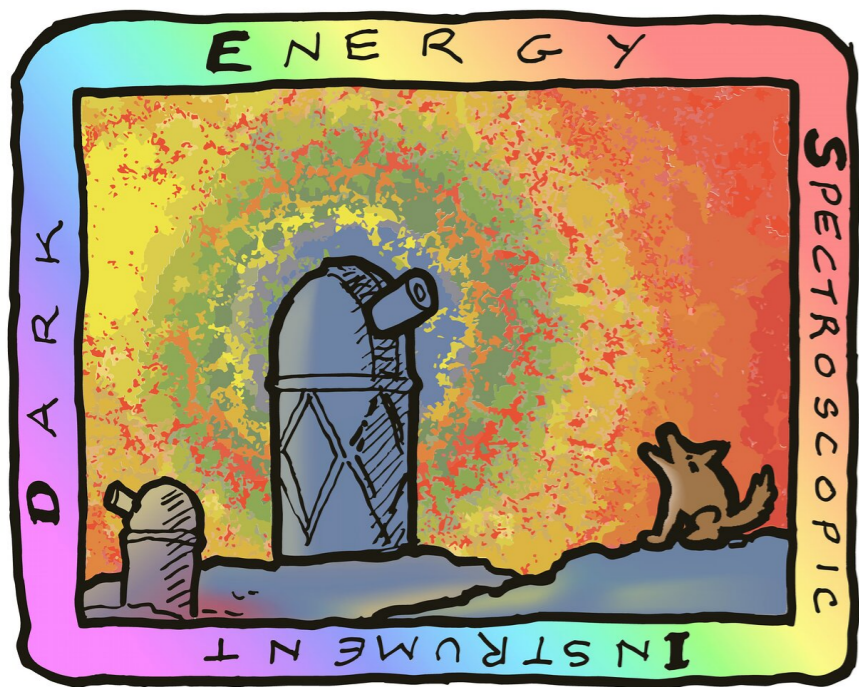


Photometric metallicities with Rubin Observatory?



- *u*-band metallicity sensitivity?
- Kahn+19 on narrowband filter legacy surveys: minimum filter width of ~100 angstroms

Contemporary Photometric Metallicity Surveys & Synergies



Conclusions

- Photometric metallicities using HST CaHK narrow-band imaging enables stellar MDF measurements in distant and/or faint galaxies
- Data enable following perspectives into the baryon cycle of dwarf galaxies:
 - Measuring the mass-metallicity relation in ultra-faint dwarf galaxies for benchmarking against simulations
 - Quantifying outflows faint dwarf galaxies through chemical evolution modeling
 - Novel metallicity gradient measurements in faint dwarf galaxies to constrain metallicity gradient formation physics
- Abundant potential for photometric metallicities with upcoming surveys, including LSST