

Decoding the Cosmos

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UNIVERSITY OF CAMBRIDGE







accuracy

precision

Observational frontier with galaxy surveys



<u>Spectroscopic</u> DESI (ground)



<u>Photometric</u> LSST (ground), Euclid (space), Roman (space)





CREDIT: RUBIN OBSERVATOR









Solving cosmological modelling challenges with machine learning





Emulation: ML-accelerated forwardmodelling of observations

high dimensional cosmological inference with ML-accelerated parts





Explainable AI: machine-assisted knowledge extraction







Simulation-based inference with ML-accelerated components



Justin Alsing (OKC/Stockholm)



Stephen Thorp (OKC/Stockholm)





Sinan Deger Boris Leistedt (IoA/KICC/Cambridge) (Imperial College London)

Arthur Loureiro (OKC/Stockholm)

With: Joel Leja, Daniel Mortlock



Photometric catalogues require redshift estimation



Blanton et al. (2003)

Loureiro et al. (2023)

Adapted from Justin Alsing



Key idea: learn joint distribution of galaxy properties over cosmic history

Machine learning models can accurately describe this complicated web of interdependencies

Figure: Hubble Ultra Deep Field





To decode the cosmos, we need to understand galaxies

Recipe for making galaxy spectra and colours

- mass
- star formation history
- dust
- gas
- metallicity
- active galactic nuclei
- redshift

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Model galaxy spectra using stellar population synthesis



add up light from all the stars (at their ages and metallicities)

CONROY, GUNN AND WHITE (2009), CONROY AND GUNN (2010)



Speeding things up with neural emulators



network weights and biases

 $\mathbf{w} = \{\mathbf{W}_1, \mathbf{b}_1, \mathbf{W}_2, \mathbf{b}_2, \dots, \mathbf{W}_n, \mathbf{b}_n\}$

Emulating spectra

I6-parameter SPS model | sub-percent accuracy | factors x I0^4 speed-up | differentiable

Alsing, Peiris, Leja, Hahn, Tojeiro, Mortlock, Leistedt, Johnson, Conroy (ApJS, 2020)

 $\mathbf{w} = \{\mathbf{W}_1, \mathbf{b}_1, \mathbf{W}_2, \mathbf{b}_2, \dots, \mathbf{W}_n, \mathbf{b}_n\}$

Emulating photometry



 $\mathbf{m}(\boldsymbol{\theta}; \mathbf{w})$

magnitudes,



Flexible neural models for distribution of galaxy properties



score-based diffusion model

learn transformation that degrades target distribution to Gaussian noise

~Brownian motion



~reverse Brownian motion

draw from random distribution: invert learned transformation to sample from target distribution

Gaussian distribution

Song and Ermon (2019), Song+ (2020A,B)



Pop-Cosmos: galaxy population model calibrated with COSMOS2020



~140,000 galaxies | 26 bands near-UV to mid-IR | deep z < 4 | simple selection r<25 Zero-point calibration | emission line corrections | Student-t uncertainty model

WEAVER ET AL (2021), ALSING ET AL (2022, APJS), LEISTEDT ET AL. (2022, APJS), ALSING ET AL (ARXIV: 2402.00935)



Learning the galaxy population model



TOY MODEL SIMULATION BY JUSTIN ALSING





Pop-Cosmos: a generative model for galaxy surveys



ALSING ET AL (2022, APJS), LEISTEDT ET AL. (2022, APJS), ALSING ET AL (2024, SUBMITTED), THORP ET AL (2024, SUBMITTED)





Forward-modelling other catalogues



ADAPTED FROM STEPHEN THORP



Pop-Cosmos as a prior for galaxy photo-z inference





THORP ET AL (2024, IN PREP)



Quality of individual redshifts









smaller errors

fewer outliers

THORP ET AL (2024, IN PREP)



Bonus: information on full galaxy population over cosmic time



star forming sequence

ALSING, THORP, DEGER, PEIRIS ET AL (SUBMITTED, ARXIV:2402.00935), THORP, PEIRIS, ALSING ET AL (SUBMITTED, ARXIV:2402.00930)











Knowledge extraction using deep learning



Luisa Lucie-Smith (MPA/Garching)



Andrew Pontzen (UCL)



Lillian Guo (UCL)



Davide Piras (Geneva)

With: Brian Nord, Jeyan Thiyagalingam



Frameworks for knowledge extraction using Al



LUCIE-SMITH, PEIRIS, PONTZEN, NORD + (PRD, 2022), LUCIE-SMITH, PEIRIS, PONTZEN, (PRL, 2024)

Model compression to enable "explainable" Al



Dark matter halo density profiles



Key ingredient in cosmological analyses and beyond (e.g. dark matter detection)



What physics is encoded in the dark matter halo profile?



Mutual information between latents and halo assembly history



Network not given assembly history of halos during training.

accretion history in specific, interpretable ways.

LUCIE-SMITH, PEIRIS, PONTZEN, NORD + (PRD, 2022), LUCIE-SMITH, PEIRIS, PONTZEN, (PRL, 2024)

Cosmological simulation

The learnt degrees of freedom nevertheless relate to halos' mass



Seeing halo properties in terms of their full evolution histories

IVE recovers known relation between inner profile and early assembly history



Halo concentration versus formation epoch (e.g. Wechsler et al 2002)

IVE discovers that **outer profile** depends on single dof related to most recent accretion rate



Relation to splashback radius (e.g. Diemer 2020)

LUCIE-SMITH, PEIRIS, PONTZEN (PHYS. REV. LETT, 2024)





Understanding non-universality in halo mass function



- Can we isolate and quantify universal and non-universal information in the HMF?

Previous work suggests linear growth related to non-universality (Ondaro-Mallea et al 2021; Euclid collaboration et al. 2023)

• How many independent degrees of freedom are needed to describe the HMF to percent-accuracy?

GUO, LUCIE-SMITH, PEIRIS, PONTZEN, PIRAS (ARXIV:2405.15850)

"More is different": emergent phenomena in cosmology

 Modern machine learning methods can help us embrace the complexity and assist us in gaining new insight.



What I found at KICP 2004—2007

It's all about the people



Successful environment for enabling interdisciplinary research

Laying down the tracks as the train of data-driven cosmology got going.

Inspiring peers

My current research snapshot







ERC Advanced Grant project focusing on Data-intensive (static and transient) cosmology with Rubin data (with a strong focus on explainable AI methods)

Fundamental physics from populations of compact object mergers

Leveraging ML methods and simulation-based inference to accelerate and enable population studies with next generation gravitational wave facilities.



Detecting Axion Dark Matter in the Sky and in the Lab

Exploring axion dark matter hypothesis through its cosmological signatures + searching for QCD axion in the lab via tunable plasmon haloscopes (founding member of ALPHA Collaboration)



Cosmology x Quantum

Emulating early universe physics within an analogue quantum simulator in the laboratory (collaboration with Cambridge Physics / Hazibabic BEC Lab)

COSMICEXPLORER: Exploring the Cosmos with the Vera Rubin Observatory



COSMOPARTICLE, WWW.PENELOPEROSECOWLE

Happy 20th birthday and congratulations KICP!



