- You want me to talk about calibrating photo-zs using the observable with the worst possible redshift resolution??
- Of course adding any two Fisher matrices generally reduces the errors on all parameters.
- One may hope that when cross-correlating a galaxy lensing survey with CMB lensing in the same area of sky there is some extra gain.
- I will do some ~pedagogical exploration of LSST $\times$ CMB lensing with emphasis on relevance of photo-z errors.

Pat McDonald (LBL), Future Surveys, 9/21/2016

## How to think about photo-zs

- Things like number density and bias (shear bias?) are generally functions of both true and photo-z.
- I think it is useful to think about them that way and then understand what priors you are imposing.

$$
\begin{aligned}
& b\left(z_{t}, z_{p}\right) \rightarrow b\left(z_{t}\right) \\
& n\left(z_{t}, z_{p}\right) \rightarrow \exp \left(-\frac{\left(z_{t}-z_{p}-\Delta_{z}\right)^{2}}{2 \sigma_{z}^{2}}\right)
\end{aligned}
$$

- $\Delta_{z}$ and $\sigma_{z}$ are free parameters in each of my $z$ bins representing systematic error in mean and rms of photo-z estimation. $b(z)$ is also free in each bin.
- Base of ~LSST galaxies in dz=0.2 (photo and true) bins with $z<2$.
- Include all correlations of density and lensing.
- $k<0.1 \mathrm{~h} / \mathrm{Mpc}$ for density, $\mathrm{l}<500$ for lensing.
- fiducial $\sigma_{z}=0.05(1+z)$
- My "CMB lensing" is an extra zero-noise source plane at CMB z (not "CMB-S4", but very clear what is added).
- Always include no-lensing Planck Fisher and DESI BAO.
- Who cares actually about nuisance parameters - tell me cosmological parameter improvements (for a relevant set of experiments and parameter space).

| base + | photo-z errors? | FoM |
| :--- | :---: | :---: |
| LSST | N | 336 |
| LSST + CMBL | N | 562 |
| LSST | Y | 231 |
| LSST + CMBL | Y | 337 |

TABLE I. FoMs for different scenarios. All include Planck ( $\sigma_{\tau}=0.009$ ) and DESI BAO. Standard DE FoM except marginalized
over neutrino mass.

- Mostly photo-z errors are just bad, although maybe surprisingly un-fatal. If anything CMBL adds less when you have them.
- Almost entirely from mean shift (width harmless).
- Maybe you're more interested in neutrino mass...

| base + | photo-z errors? | $\sigma_{m_{\nu}}[\mathrm{eV}]$ |
| :--- | :---: | :---: |
| LSST | N | 0.021 |
| LSST+CMBL | N | 0.020 |
| LSST | Y | 0.027 |
| LSST+CMBL | Y | 0.022 |

- Marginalized mean shift errors, FoM scenario.
- Neutrino mass case almost identical.



## Conclusions

- Photometric surveys and CMB lensing are complementary, sometimes more so with photo-z uncertainty.
- Not clear it is useful to think of this as CMBL calibrating photo-zs.
- This is a toy calculation, intended to stimulate thinking about more complete ones.
- Preliminary - I don't guarantee bug-free.

