

The Experimental Challenge of 21 cm Observations

Miguel F. Morales

Future Cosmic Surveys, September 22nd 2016

EoR: PAPER, LOFAR, MWA, HERA

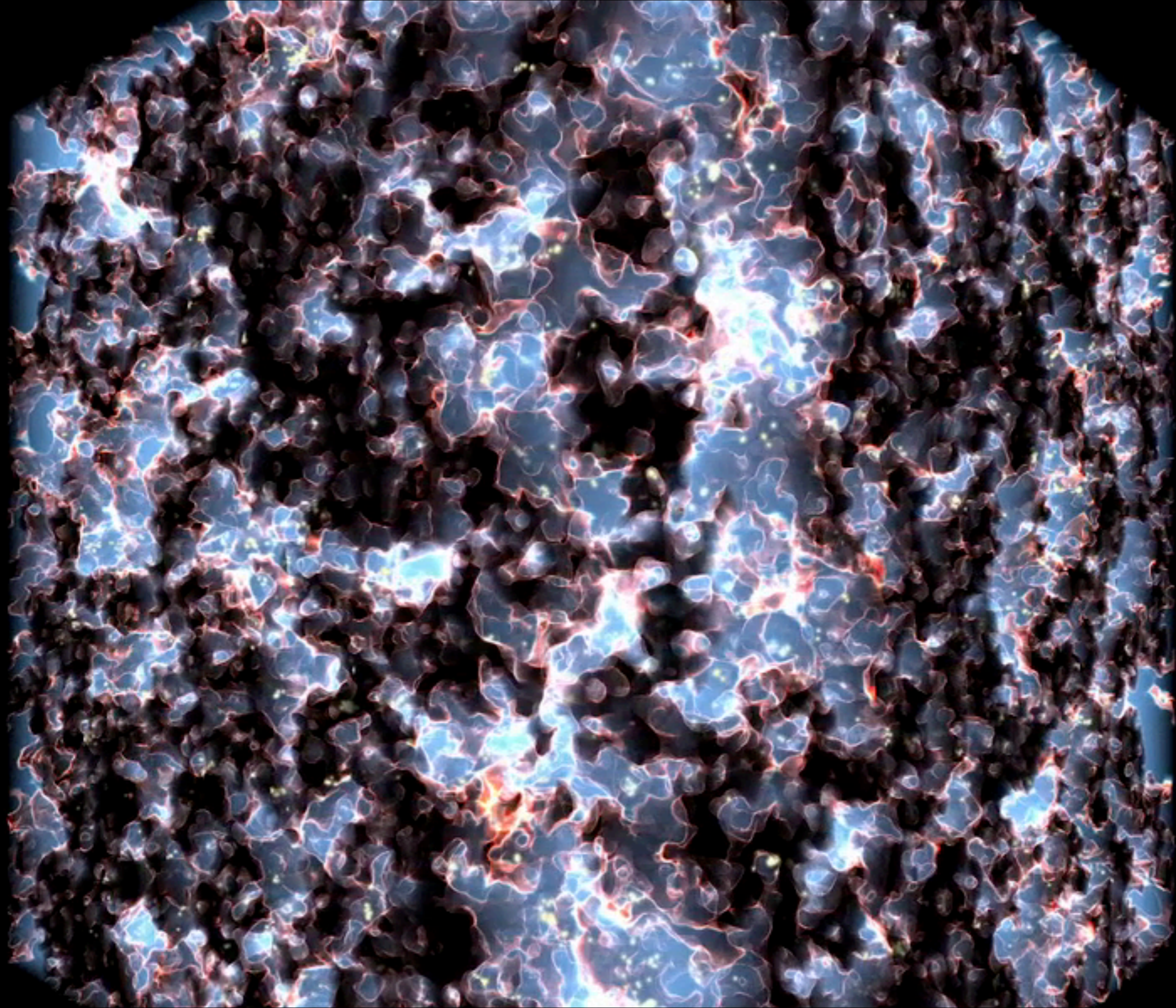


Dark Energy: CHIME, HIRAX

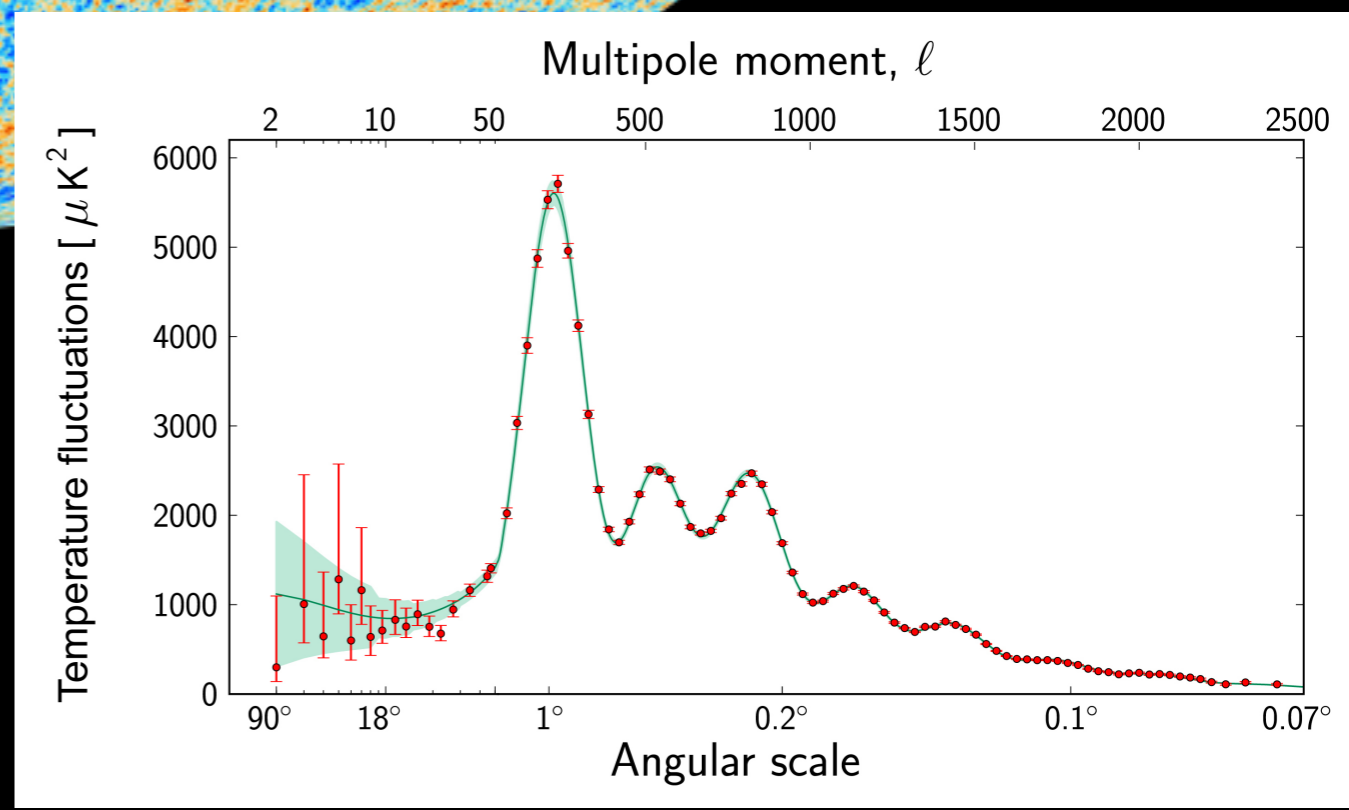
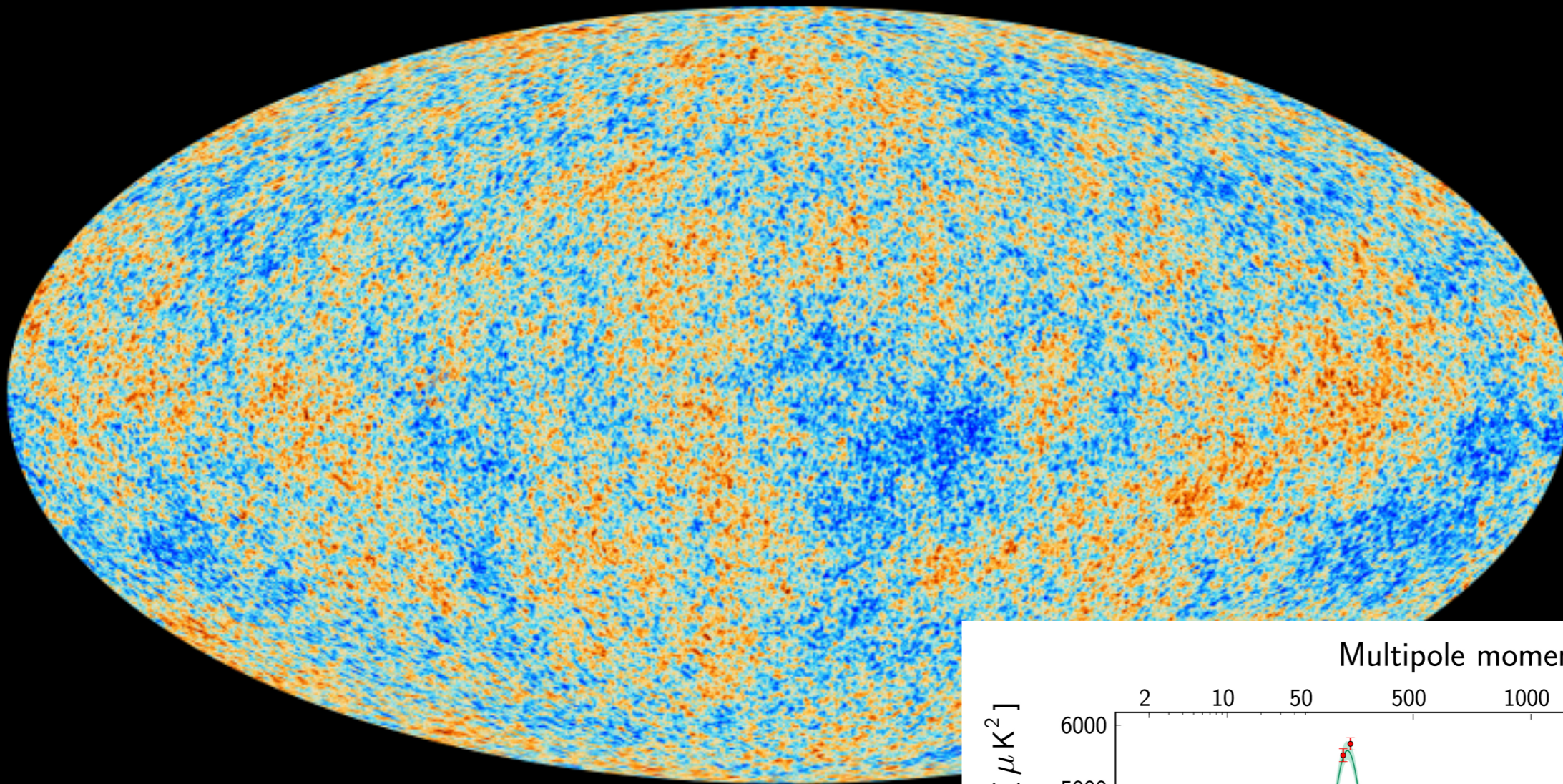


COMAP

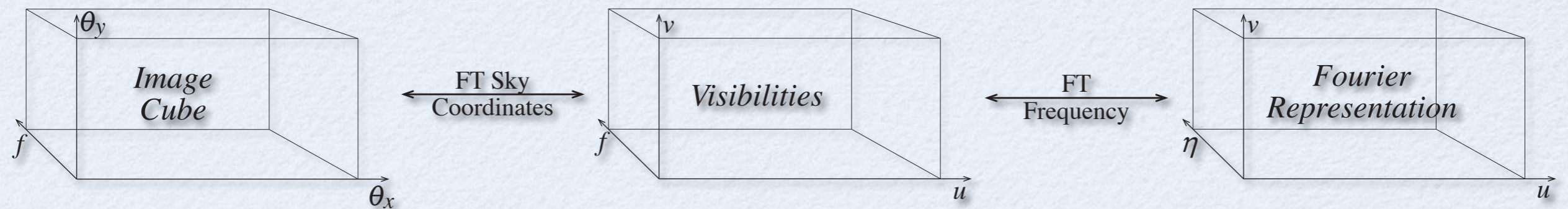
Carbon Monoxide intensity mapping to
study star formation at $z=2.4-2.8$ and
reionization



CMB

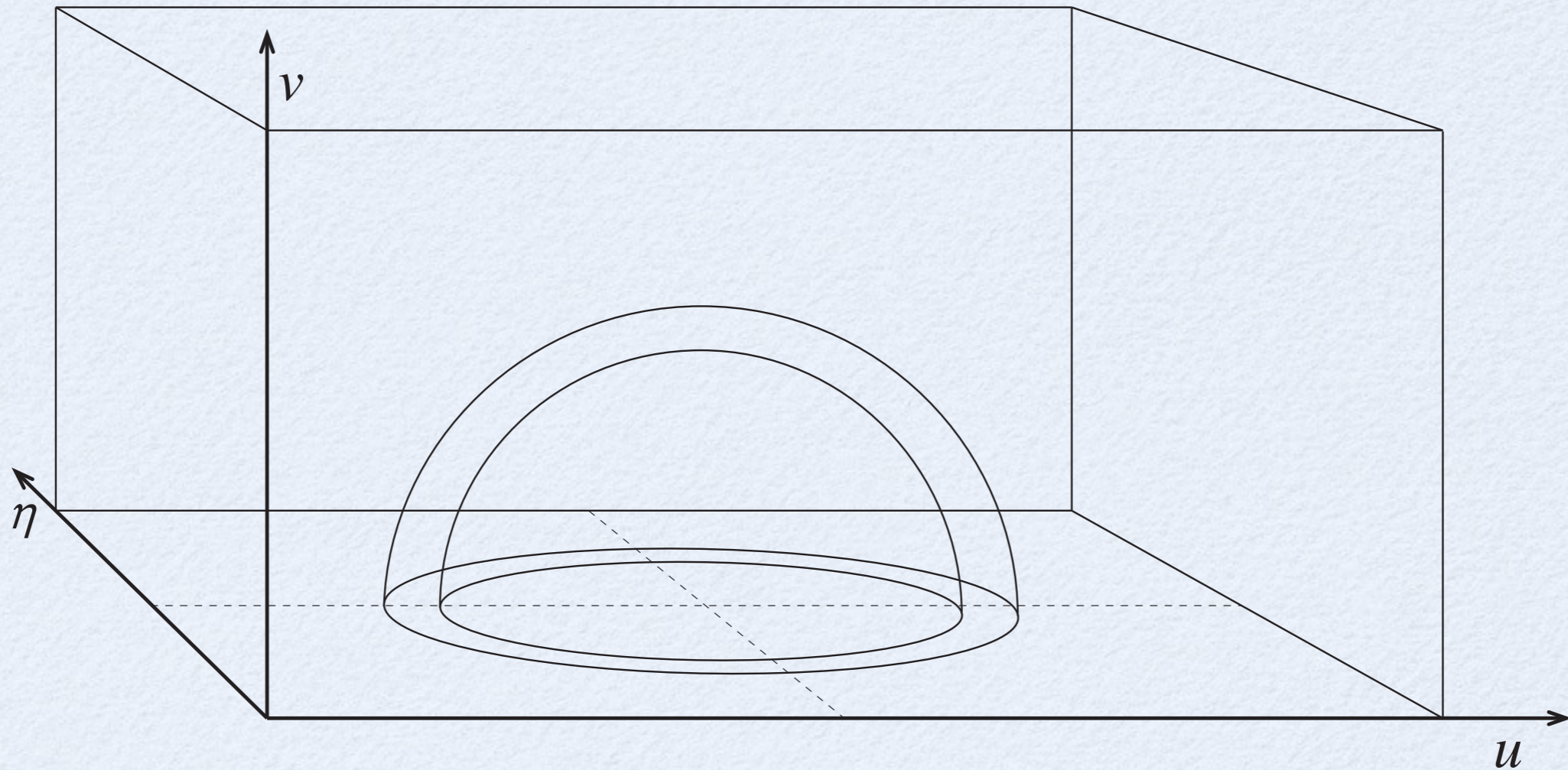


Statistical EoR detection



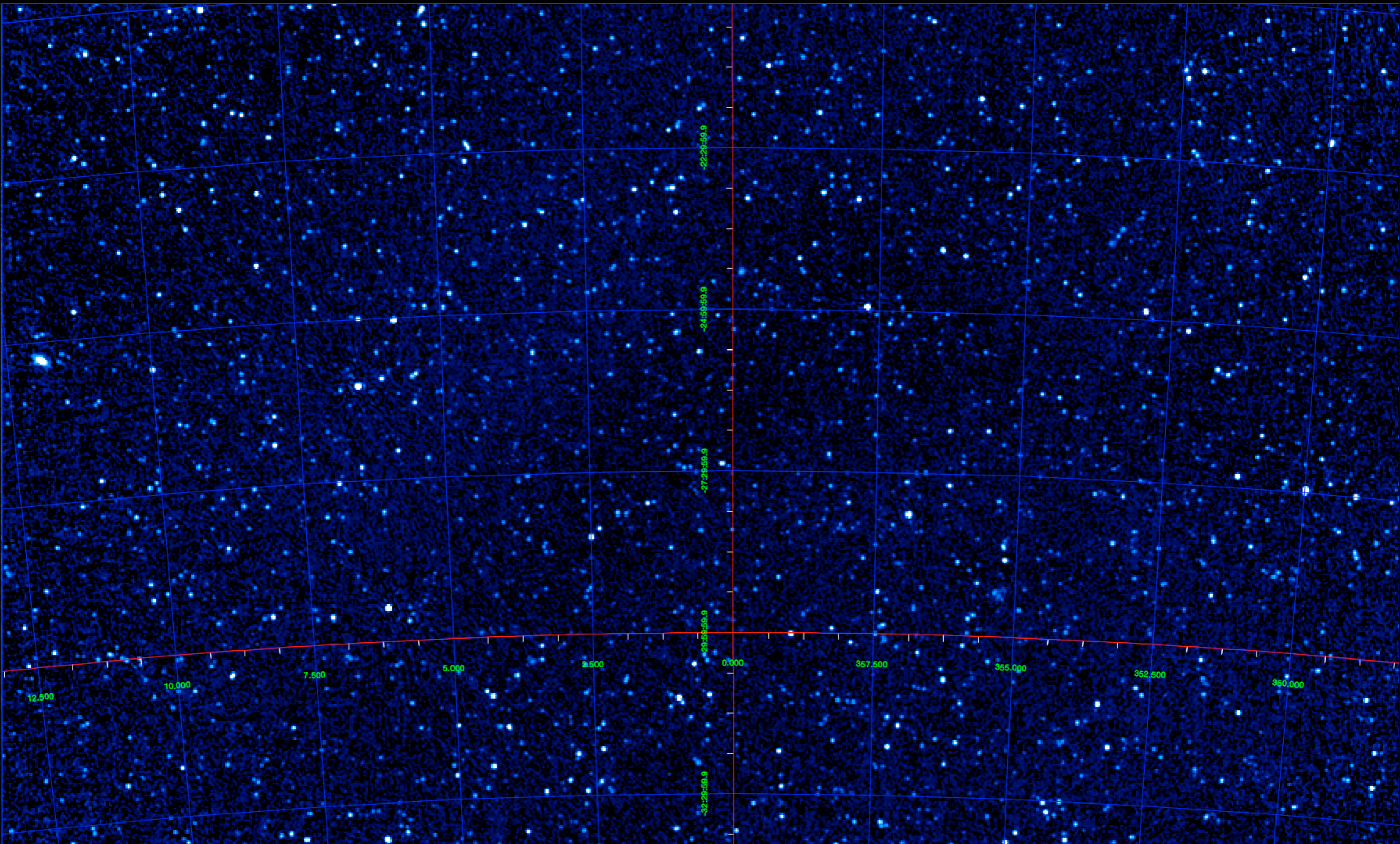
Morales & Hewitt (2004)

Spherical symmetry



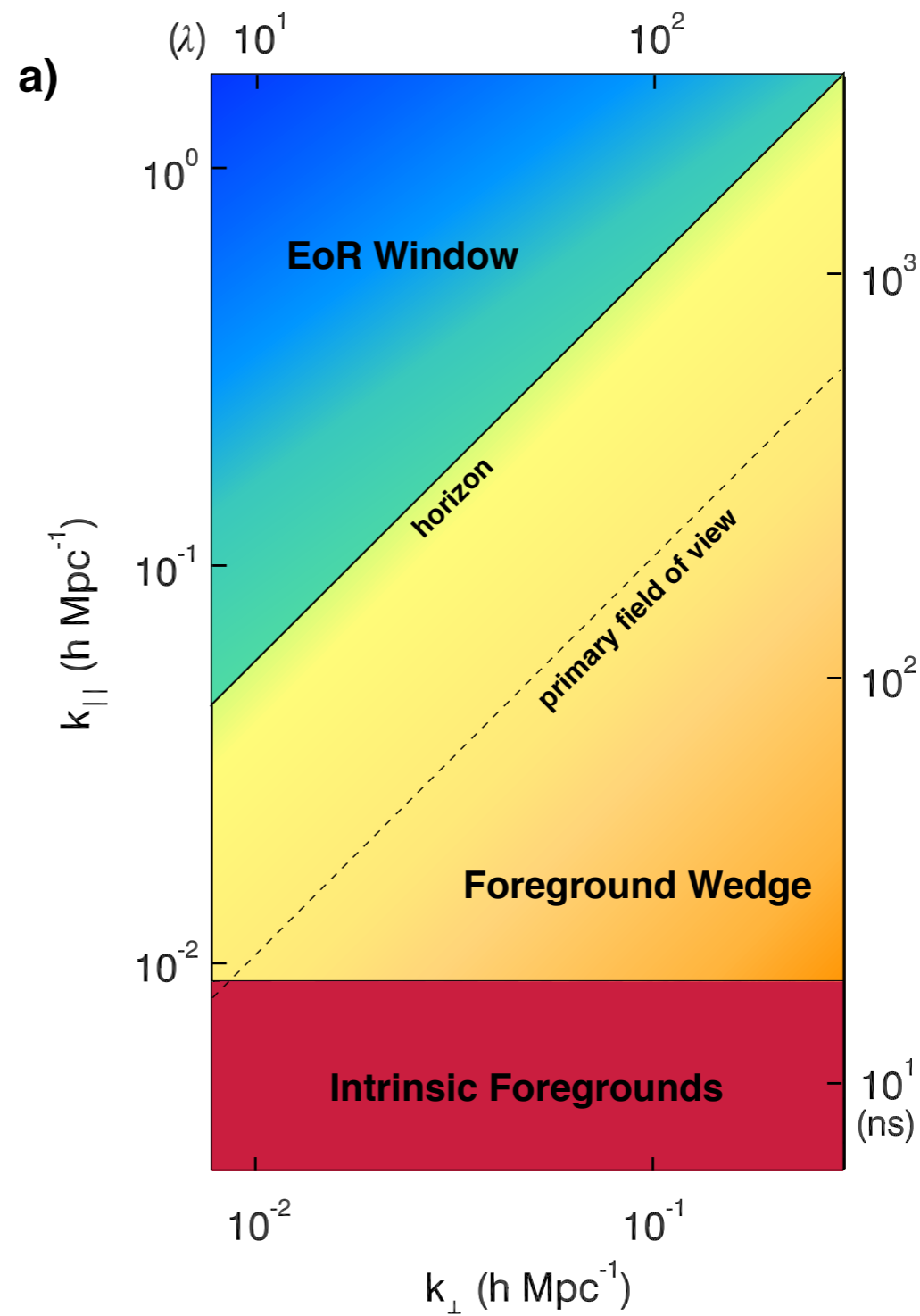
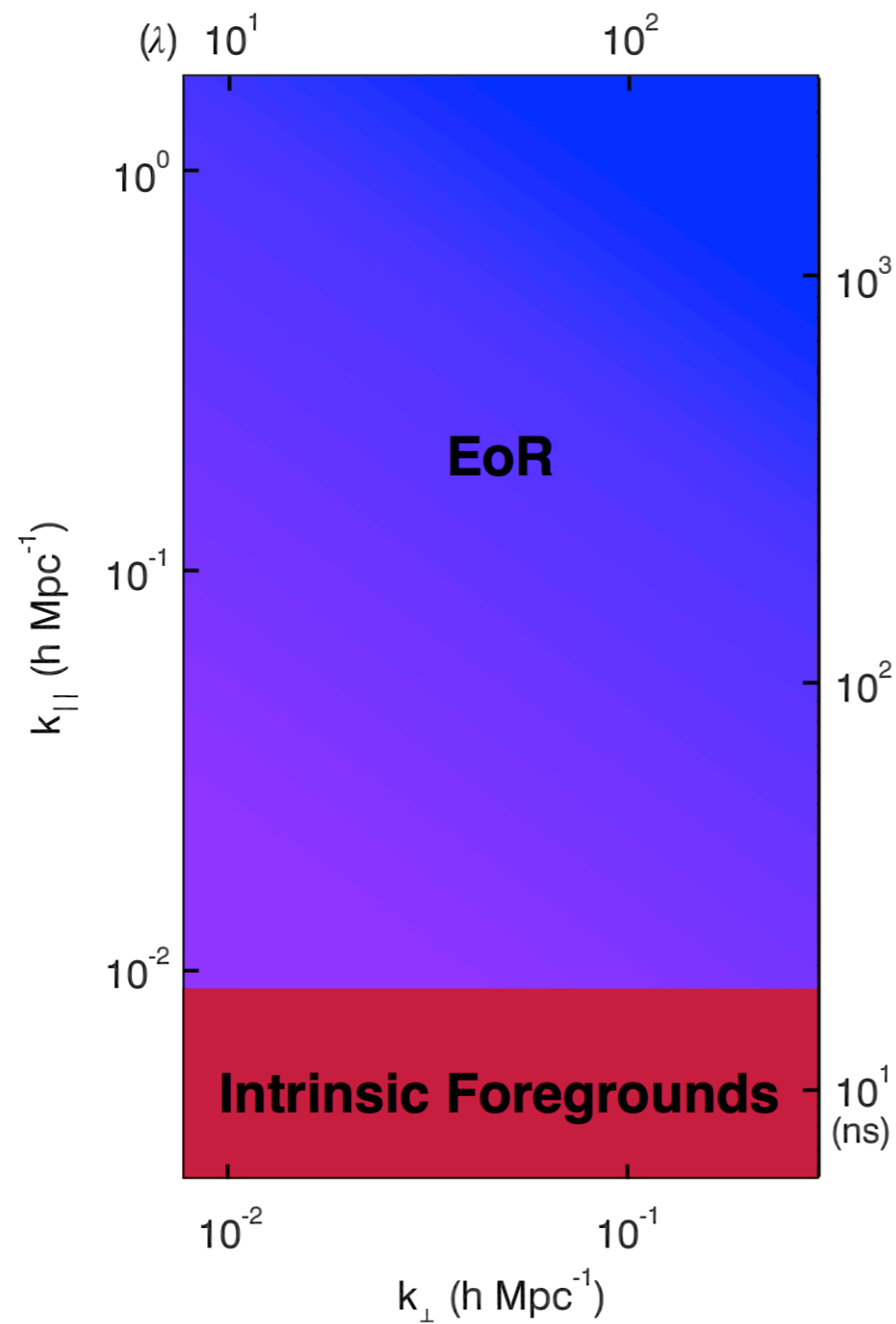
Morales & Hewitt (2004)

Moon to scale



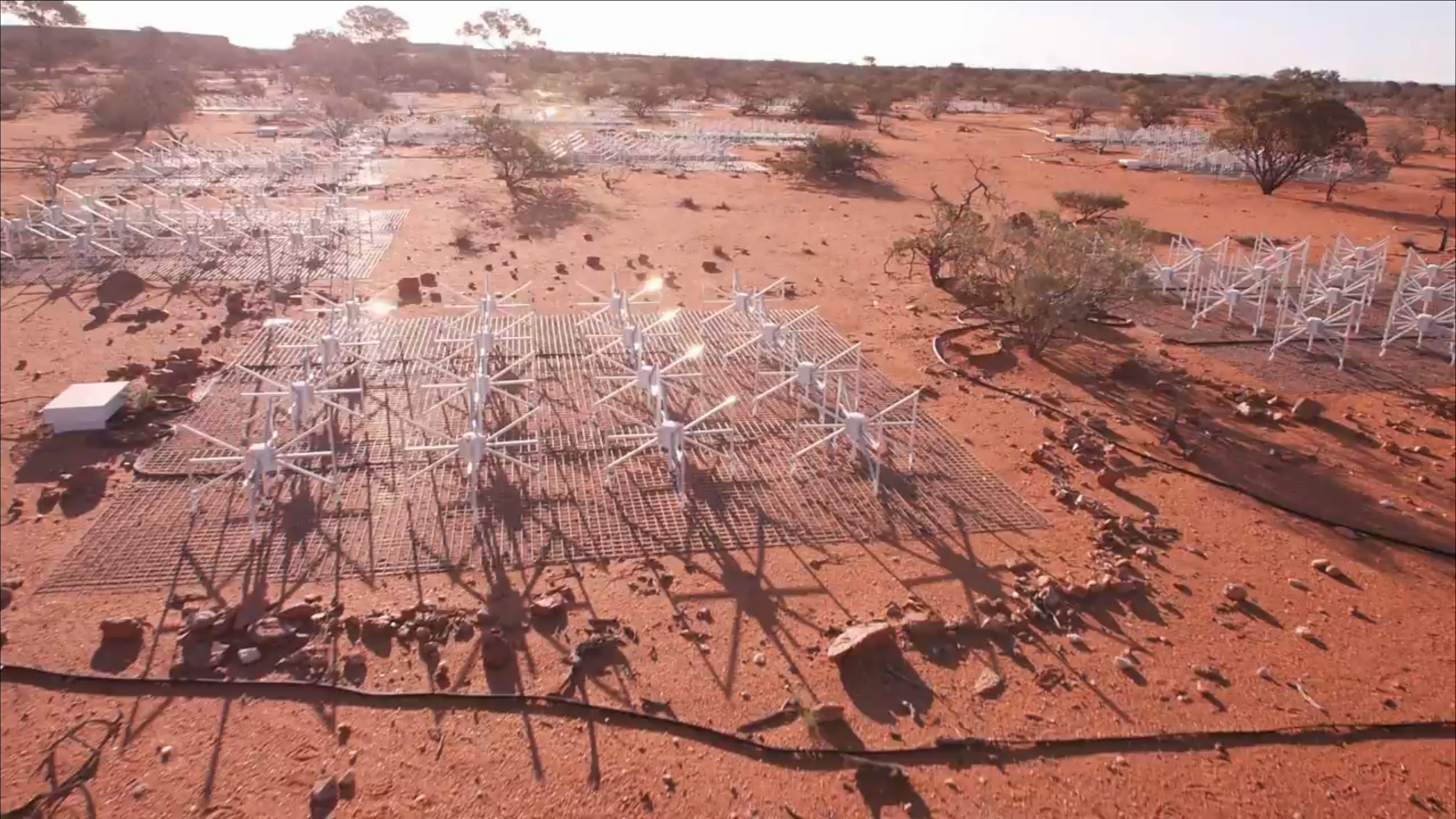
2 minutes of data, dark field, signal dominated

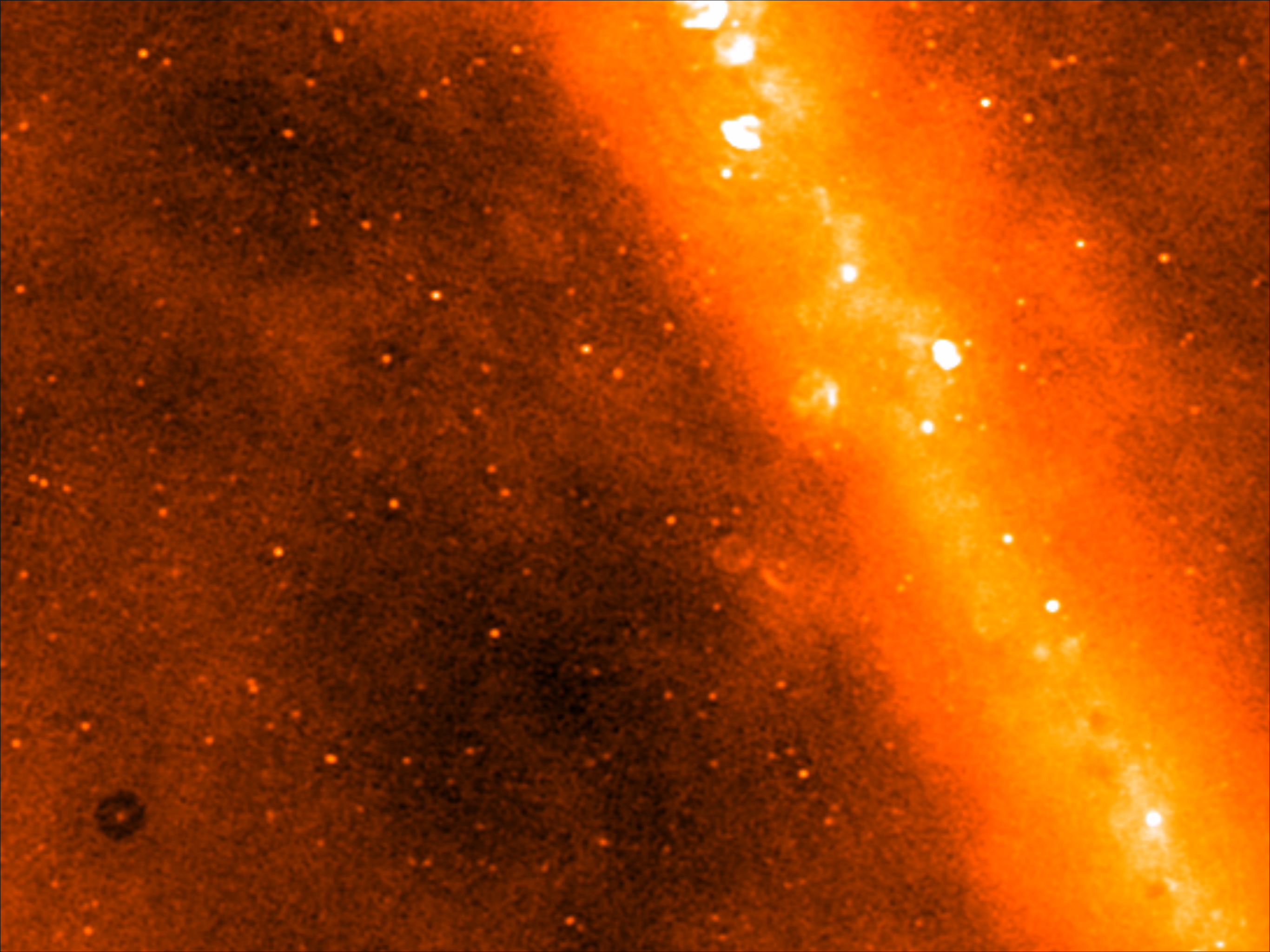
21 cm window



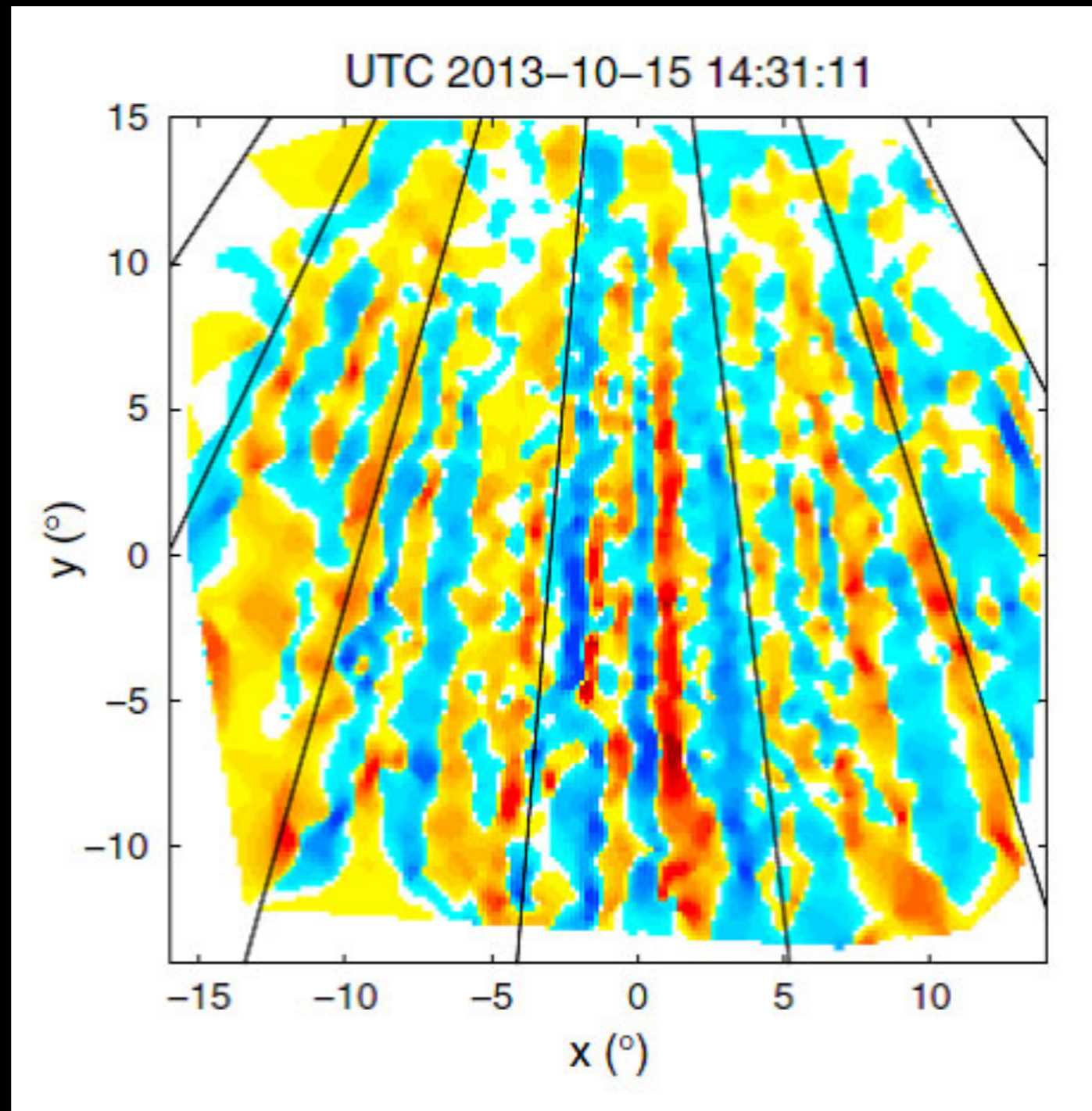
Looking at the data...



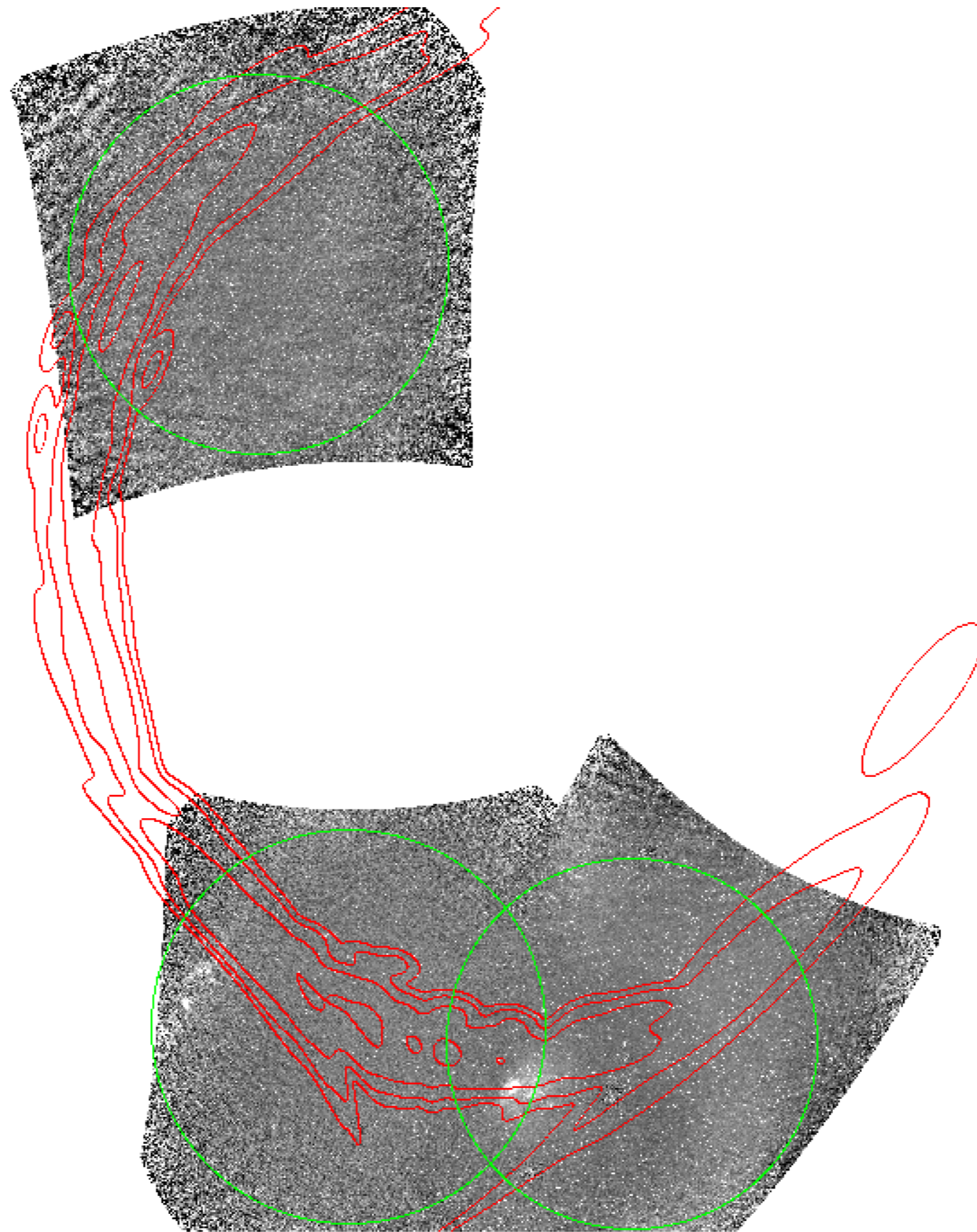




Ionosphere & Plasmasphere



Variable sky (LIGO follow up)





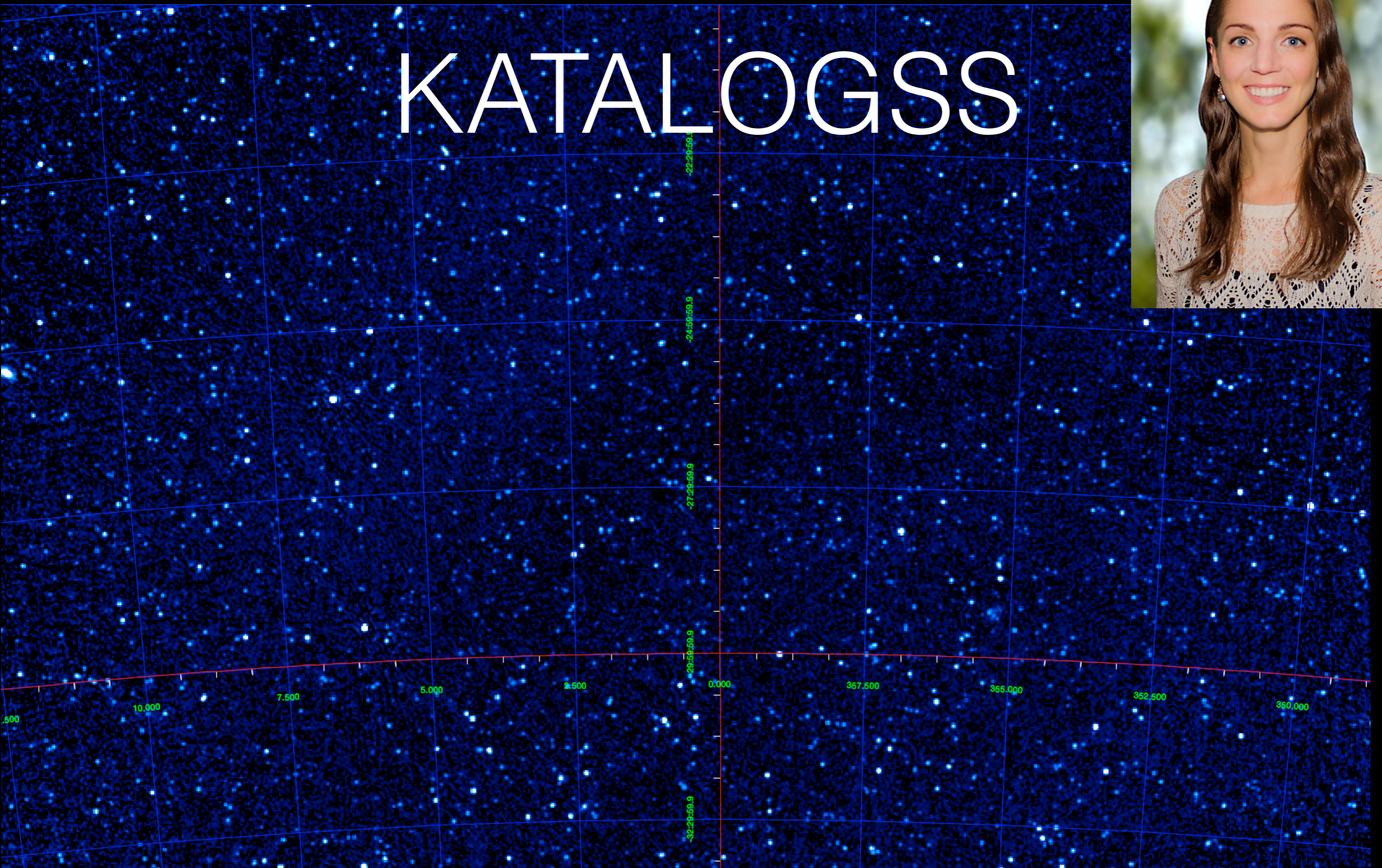








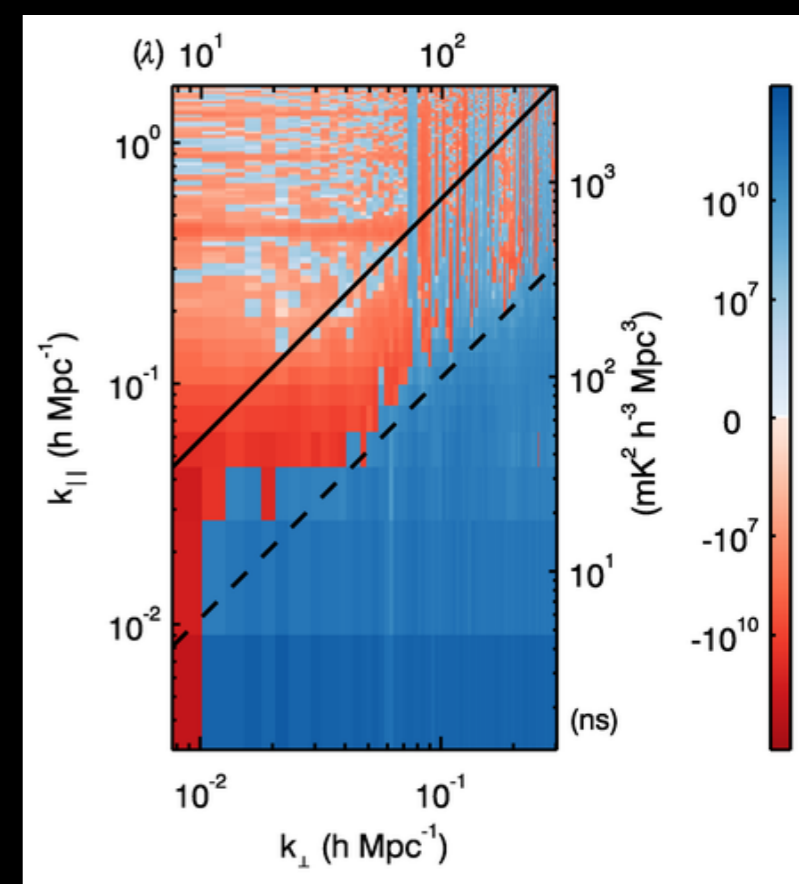
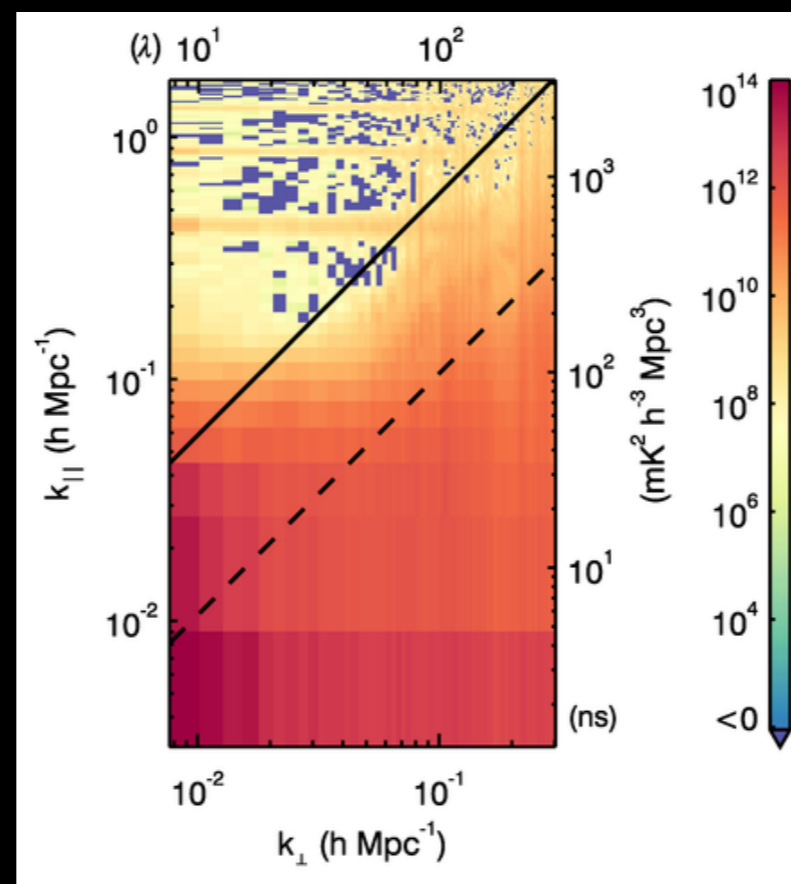
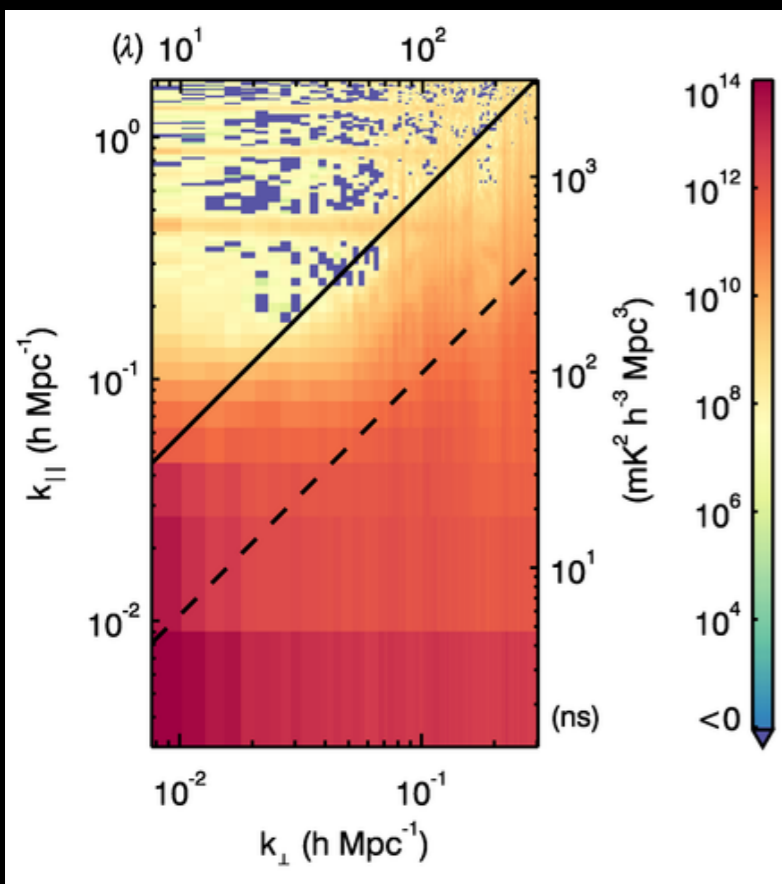
KATALOGSS



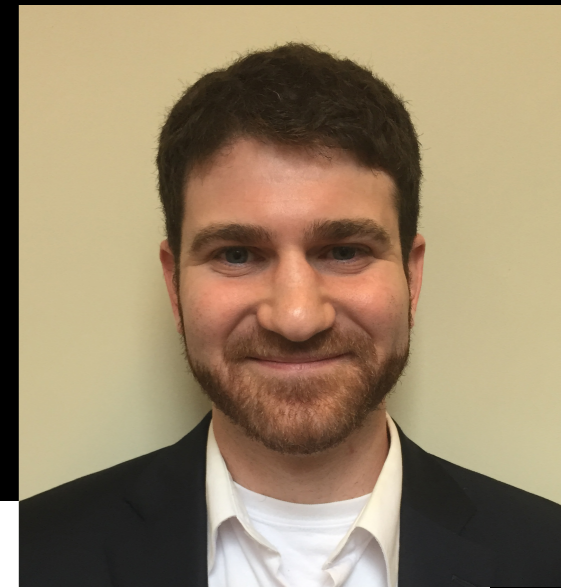
Precision foreground catalog

Carroll et al. 2016

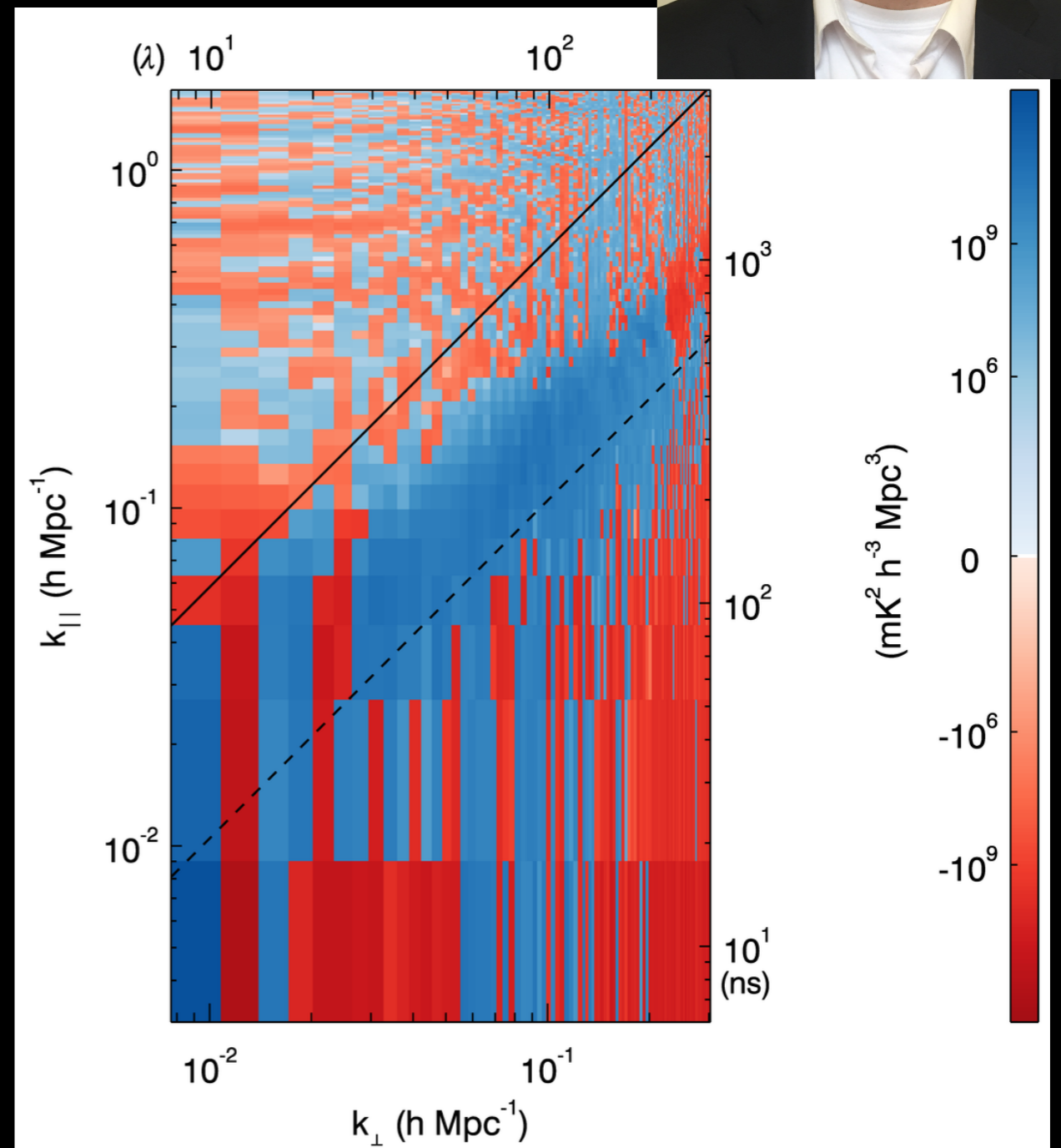
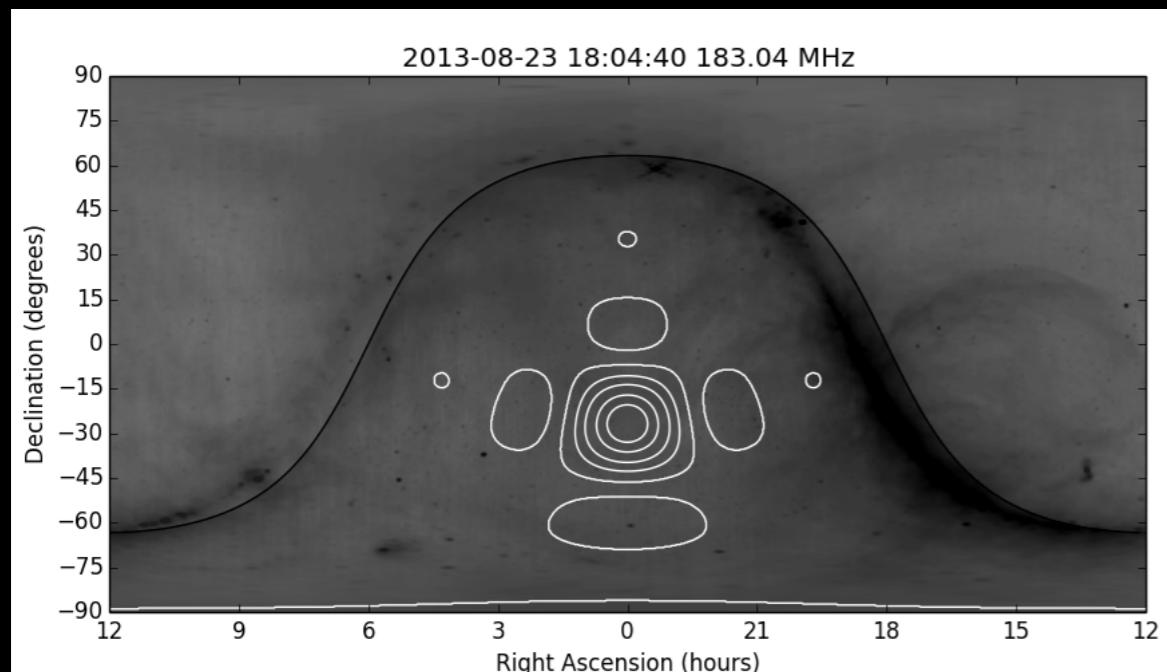
KATALOGSS



Sidelobe Sources



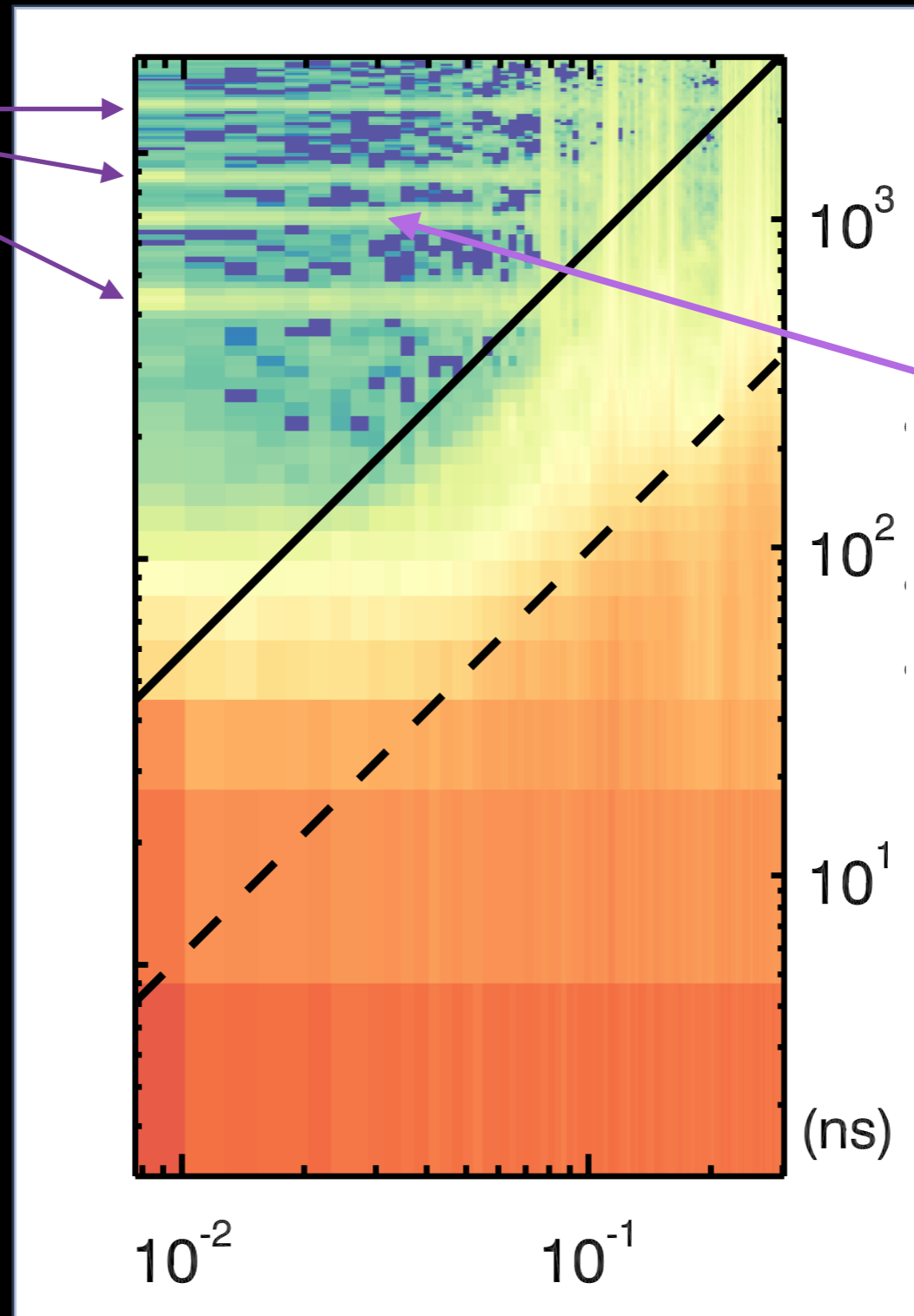
- Sidelobe sources bleed higher in k_{\parallel}
- Most critical for opening the EoR window



Precision Calibration



Frequency gaps

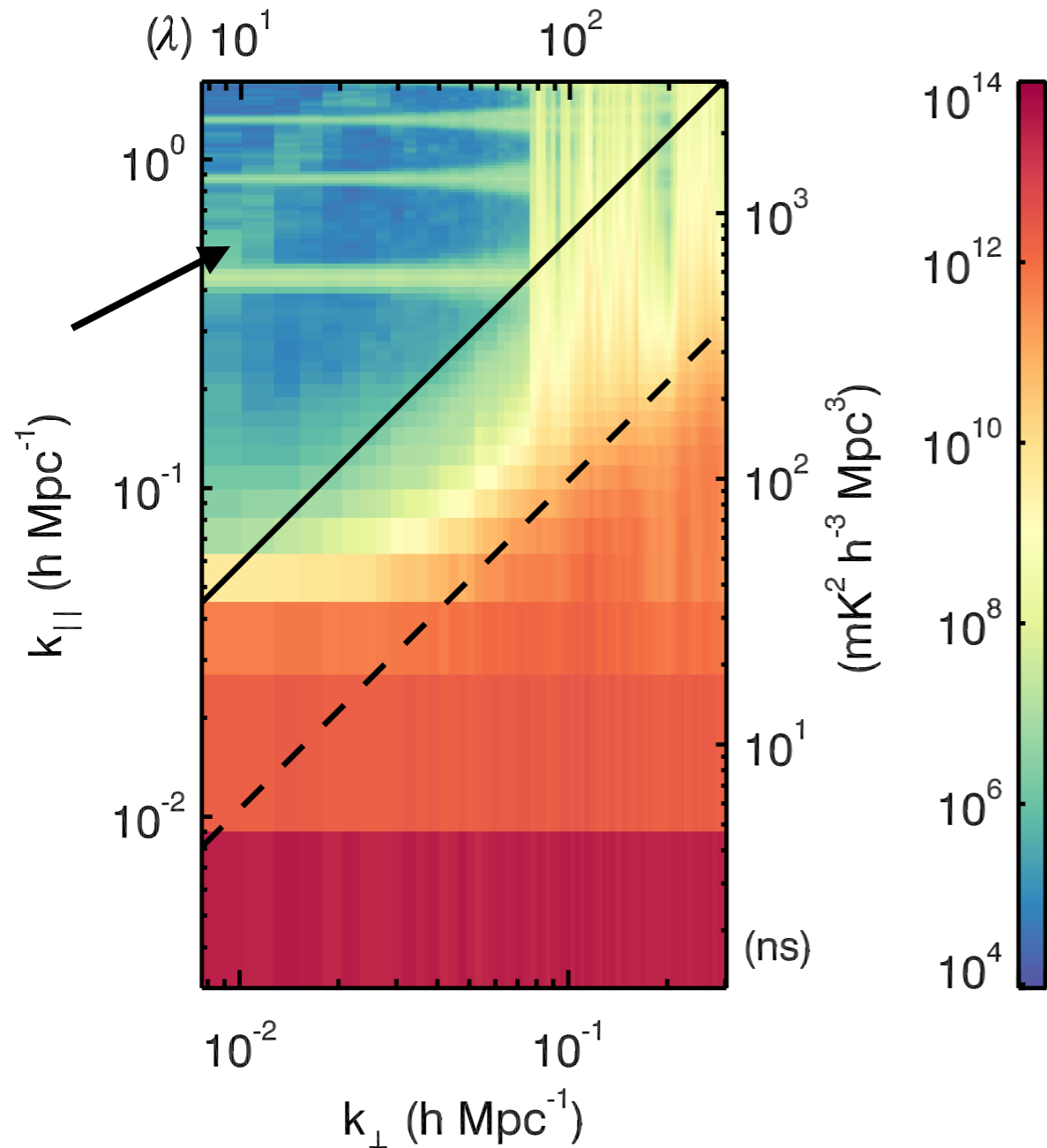


150 m cable reflection

Beardsley et al. 2016

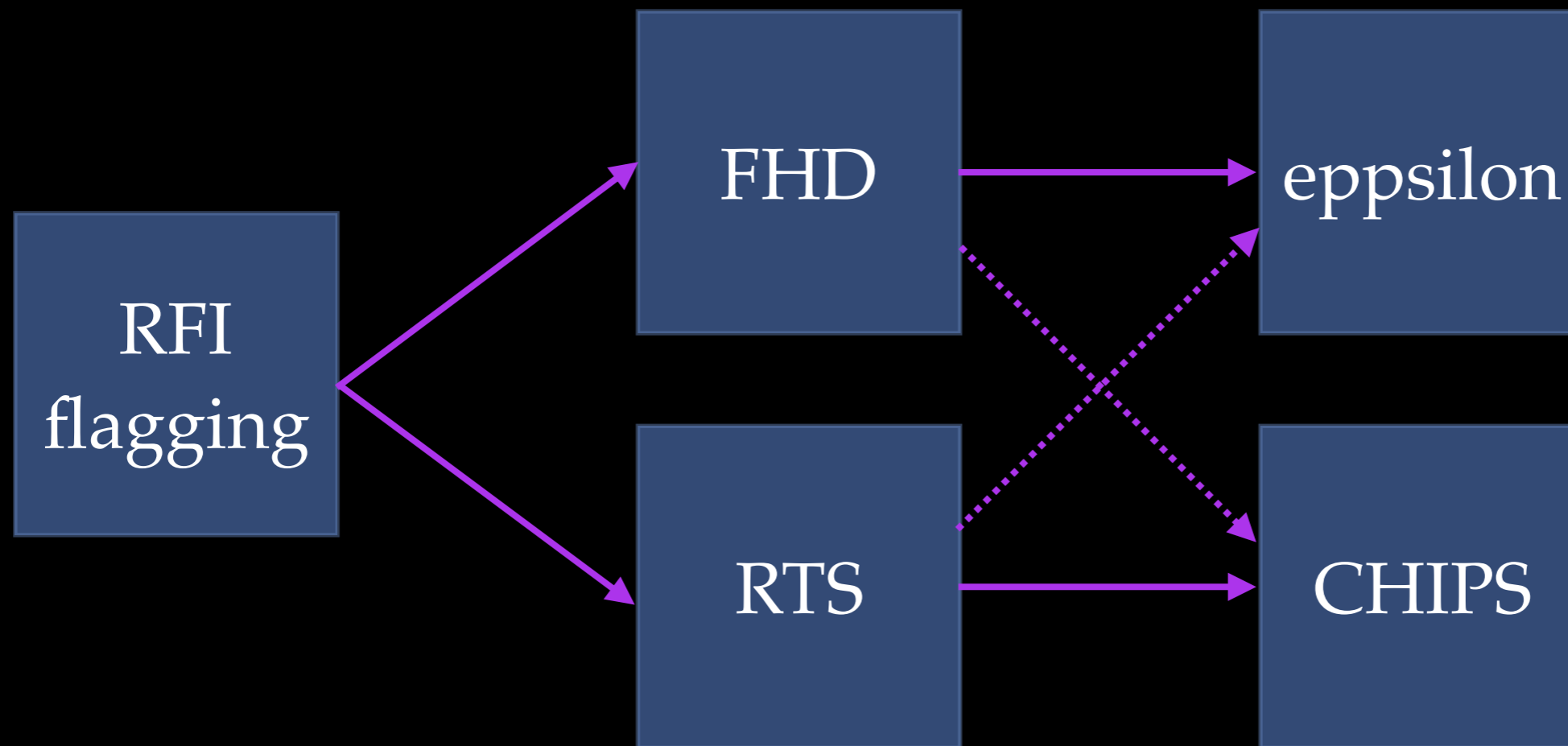
Barry et al. 2016

Precision Analysis



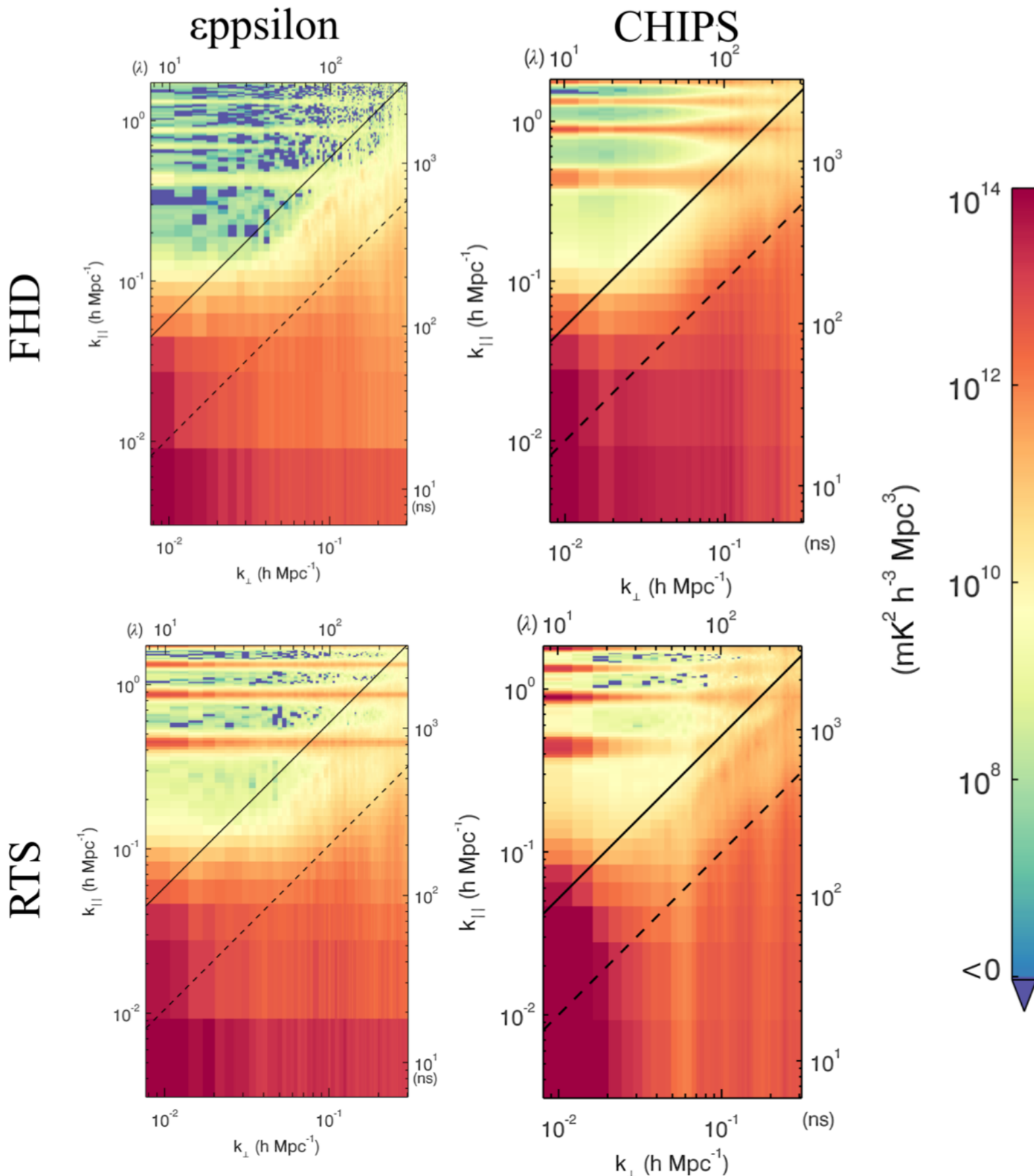
- Antenna gridding kernels must be calculated to better than $1/7000$ wavelength

MWA Community comparison framework



Foreground Removal

Power Spectrum Method



Where are we going...

- MWA phase II commissioning
- HERA funded and being built



08/03/2016









Conclusion

- Radio is a natural place for a physicist
- Many large telescopes and precursors under construction
- Intersection of big data and precision analysis



