

# Black hole entropy of gauge fields

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Black holes possess an entropy given by the Bekenstein-Hawking formula:

$$S_{\text{BH}} = \frac{Ac^3}{4G\hbar}.$$

Can be inferred **macroscopically**.

- From the Hawking temperature  $T_{\text{H}} = \frac{\hbar}{2\pi} \kappa$  and the first law of black hole mechanics  $\delta E = T_{\text{H}} \delta S_{\text{BH}}$ .<sup>1</sup>
- From a saddle point evaluation of the Euclidean partition function.<sup>2</sup>

But the statistical meaning of  $S_{\text{BH}}$  is not clear.

What is the statistical mechanics of black hole thermodynamics?

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<sup>1</sup>Hawking 1975

<sup>2</sup>Gibbons & Hawking 1977

# Entanglement entropy

The black hole horizon also has an entanglement entropy.<sup>3</sup>

We have a tensor product decomposition, and partial trace

$$\mathcal{H} = \mathcal{H}_{\text{in}} \otimes \mathcal{H}_{\text{out}}, \quad \rho_{\text{out}} = \text{tr}_{\text{in}} |\psi\rangle\langle\psi|.$$

The entropy of  $\rho_{\Omega}$  is the entanglement entropy:

$$S_{\text{ent}} = -\text{tr} \rho_{\text{out}} \ln \rho_{\text{out}}.$$

Statistical meaning: entropy comes from missing correlations due to inaccessible black hole interior.

How is this related to the macroscopic quantity  $S_{\text{BH}}$ ?

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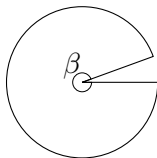
<sup>3</sup>Sorkin 1983; Bombelli, Koul, Lee, & Sorkin 1986; Srednicki 1993

# How are $S_{\text{BH}}$ and $S_{\text{ent}}$ related?

To relate  $S_{\text{BH}}$  and  $S_{\text{ent}}$ , we use the conical method.<sup>4</sup>

For simplicity we consider a Rindler horizon.

Let  $Z(\beta)$  be the Euclidean path integral on a cone of angle  $\beta$ , times  $D - 2$  flat and compact transverse directions.



$$Z = \int \mathcal{D}\phi e^{-S[\phi]}$$
$$S_{\text{cone}} \equiv \left(1 - \beta \frac{\partial}{\partial \beta}\right) \ln Z \Big|_{\beta=2\pi}$$

This **conical entropy** is equivalent to varying the period at  $\infty$ .

**Strategy:** Compute  $S_{\text{cone}}$  in effective theory, and in the quantum theory.

<sup>4</sup>Bañados, Teitelboim & Zanelli 1993

## Relating $S_{\text{cone}}$ and $S_{\text{BH}}$

Suppose we integrate out the matter fields, leading to an effective action

$$\int \mathcal{D}\phi e^{-S[\phi]} = e^{-\int \sqrt{g} L_{\text{eff}}}, \quad L_{\text{eff}} = \frac{1}{16\pi G_{\text{eff}}} (R - 2\Lambda_{\text{eff}} + \dots)$$

The conical entropy formula gives the Bekenstein-Hawking formula <sup>5</sup>

$$S_{\text{cone}} = \frac{A}{4G_{\text{eff}}} = S_{\text{BH}}.$$

In terms of the **effective** Newton's constant,  $G_{\text{eff}}$ .

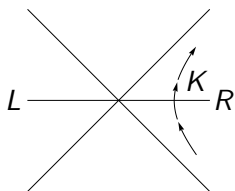
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<sup>5</sup>Susskind & Uglum 1994; Jacobson 1994

# Relating $S_{\text{cone}}$ and $S_{\text{ent}}$

Now consider evaluating the entropy in the quantum theory.

The Minkowski vacuum  $|0\rangle$  restricted to one Rindler wedge is thermal in the boost generator  $K$ :<sup>6</sup>



$$\rho_R \equiv \text{tr}_L |0\rangle\langle 0| = \frac{e^{-2\pi K}}{Z(2\pi)},$$
$$Z(\beta) \equiv \text{tr} e^{-\beta K}.$$

Varying  $\beta$  is equivalent to varying the temperature of a thermal state:

$$S_{\text{cone}} \equiv \left(1 - \beta \frac{\partial}{\partial \beta}\right) \ln Z \Big|_{\beta=2\pi} = -\text{tr} \rho_R \ln \rho_R = S_{\text{ent}}.$$

<sup>6</sup>Bisognano & Wichmann 1975

# Not so fast!

The conical geometry also has a singular curvature at the tip:<sup>7</sup>

$$R_{abcd}(x) = (2\pi - \beta)\epsilon_{ab}\epsilon_{cd}\delta_{\Sigma}(x).$$

Nonminimally coupled matter interacts with this curvature.

The contribution to the conical entropy coming from the tip is:

$$\langle S_{\text{Wald}} \rangle = -2\pi \int_{\Sigma} \sqrt{h} \left\langle \frac{\partial L}{\partial R_{abcd}} \right\rangle \epsilon_{ab}\epsilon_{cd}.$$

This term is the contribution of the matter fields to the Wald entropy.<sup>8</sup>

Thus for nonminimally coupled matter,<sup>9</sup>

$$S_{\text{BH}} = S_{\text{ent}} + \langle S_{\text{Wald}} \rangle.$$

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<sup>7</sup>Fursaev & Solodukhin 1995

<sup>8</sup>Wald 1993; Visser 1993; Jacobson, Kang & Myers 1993

<sup>9</sup>As suggested by arguments in Frolov & Fursaev 1997

# One loop results

The conical entropy has been calculated for free fields of spin  $\leq 2$ ,

$$S_{\text{cone}} = A c_1 \left( 2\pi \int_{\epsilon^2}^{\infty} ds \frac{e^{-m^2 s}}{(4\pi s)^{D/2}} \right).$$

$c_1$  depends on the field and  $N$ , the number of on-shell degrees of freedom<sup>10</sup>

Spin	Field	$N$	$c_1$
0	Nonminimally coupled scalar	1	$\frac{N}{6} - \xi$
$\frac{1}{2}$	Dirac spinor	$2^{\lfloor \frac{D}{2} \rfloor - 1}$	$\frac{N}{6}$
1	Maxwell field	$D - 2$	$\frac{N}{6} - 1$
$\frac{3}{2}$	Rarita-Schwinger field	$(D - 3)2^{\lfloor \frac{D}{2} \rfloor - 1}$	$\frac{N}{6}$
2	Graviton	$\frac{D(D-3)}{2}$	$\frac{N}{6} - \frac{D^2 - D + 4}{2}$

For gauge fields there is a mysterious **contact term**.<sup>11</sup>

<sup>10</sup>Solodukhin 2011

<sup>11</sup>Kabat 1995



# Electromagnetic field

For the electromagnetic field the Lagrangian is

$$L = \frac{1}{4} F^{ab} F_{ab}, \quad \Rightarrow \quad \text{expect } S_{\text{Wald}} = 0.$$

We add ghosts  $c$  and  $\bar{c}$ , a gauge fixing term, and integrate by parts:

$$L' = -\frac{1}{2} A^a (g_{ab} \nabla^2 - R_{ab}) A^b - \bar{c} \nabla^2 c.$$

The Wald entropy contribution from the gauge field is:

$$\langle S_{\text{Wald}} \rangle = -\pi \int_{\Sigma} \sqrt{h} g_{\perp}^{ab} \langle A_a A_b \rangle.$$

Evaluated using the heat-kernel regularization it gives  $c_1 = -1$ .

**Problem:** Gauge invariance? What about  $D = 2$ , where there are no local degrees of freedom?

# Compact spacetime

We now consider  $D = 2$  and compactify (e.g. 2D de Sitter).

In 2D, any vector field can be written as

$$A = d\phi + \delta\psi + B, \quad \Delta B = 0$$

The vector field cancels with the ghosts up to zero modes.

The number of zero modes (vector minus two ghosts) is  $2g - 2 = -\chi$ , where  $\chi$  is the Euler characteristic.

Using Gauss-Bonnet, we can write  $\chi = \frac{1}{4\pi} \int \sqrt{g} R$ .

Zero mode contribution to the effective action is proportional to  $\int \sqrt{g} R$

$$S_{\text{zero modes}} = - \left( 2\pi \int_{\epsilon^2}^{\infty} ds \frac{e^{-m^2 s}}{(4\pi s)^{D/2}} \right).$$

The  $c_1 = -1$  in the conical entropy comes from **zero modes**.

# Reduced phase space

Heat kernel method *does not* treat zero modes properly.

2D gauge theory has a huge symmetry group: area-preserving diffeomorphisms. It is “almost topological” and can be solved exactly:<sup>12</sup>

$$Z = \sum_{E \in q\mathbb{Z}} e^{-\frac{1}{2}VE^2}.$$

Since  $V \propto \beta$ , the conical entropy is

$$S_{\text{cone}} = \left(1 - V \frac{\partial}{\partial V}\right) \ln Z = - \sum_E p_E \ln p_E = S_{\text{ent}}.$$

This is finite, positive, and equal to the entanglement entropy.

One can show that  $S_{\text{cone}} = S_{\text{ent}}$  for 2D Yang-Mills as well.

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<sup>12</sup>Witten 1991

## Conclusions:

- Black hole entropy is closely related to entanglement entropy

$$S_{\text{BH}} = S_{\text{ent}} + \langle S_{\text{Wald}} \rangle .$$

- But the “contact term” in the entropy of gauge fields can't be explained this way; it is absent when the partition function is evaluated carefully (in  $D = 2$ ).

## Future work:

- Gauge theories in  $D > 2$ .
- Linearized gravity.

For details see arXiv:1206.5831.

**Thank you.**