

The role of hadronic cascades in GRB models of efficient neutrino production

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High-Energy Messengers: Connecting the Non-Thermal Extragalactic Backgrounds

Outline

- Introduction : UHECRs and neutrinos
- Motivation and goals
- Method
- Results & implications for GRB models

Intro

UHECRs & Neutrinos

GRBs as a candidate

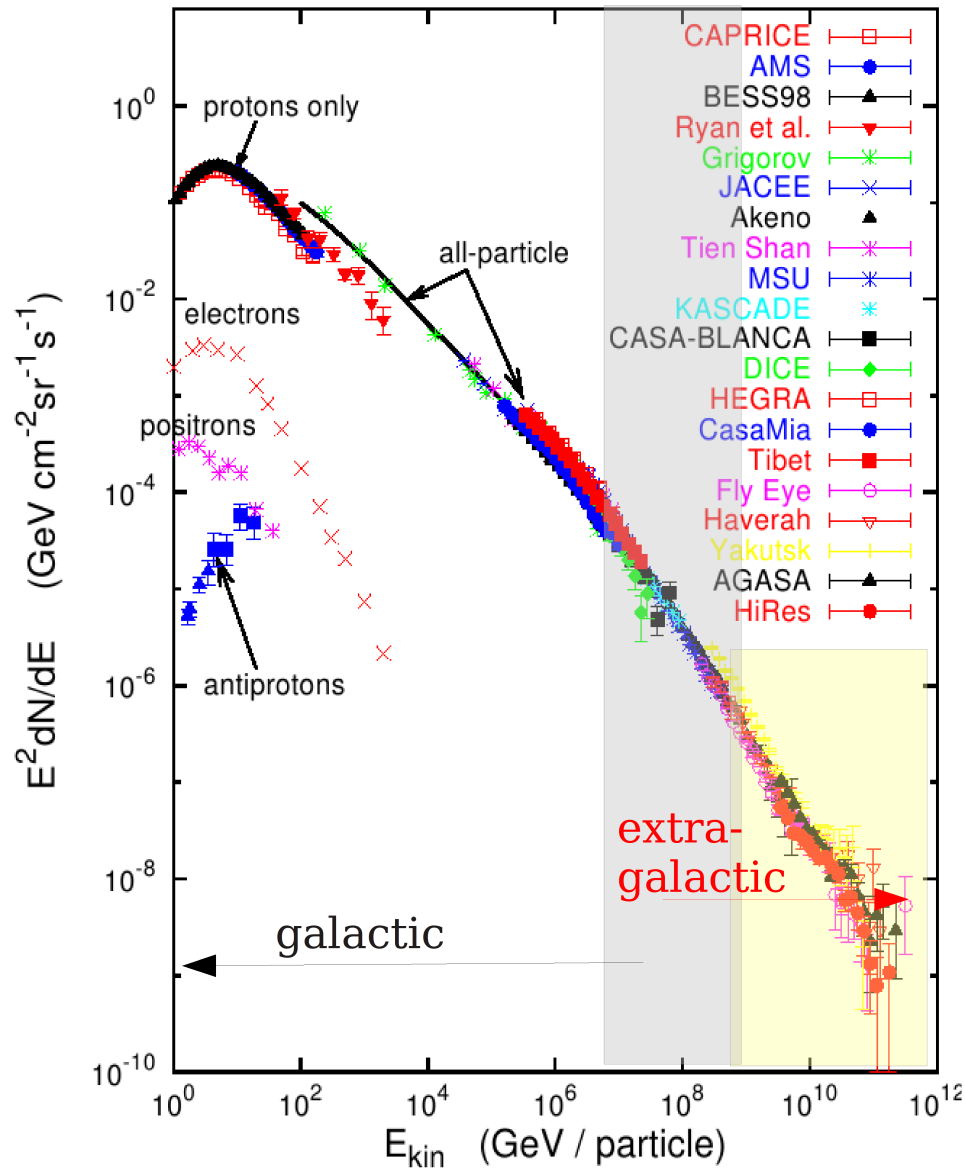
Motivation & Goals

Method

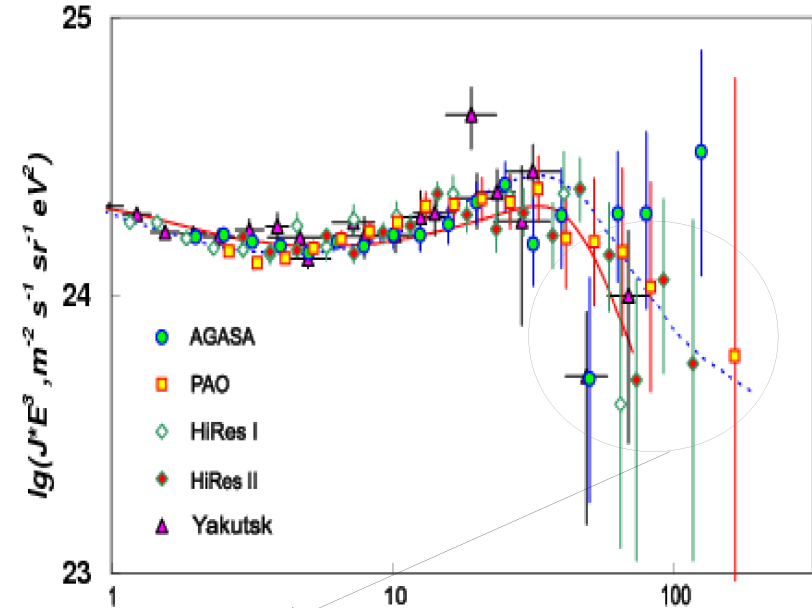
Results

Implication

Energies and rates of the cosmic-ray particles

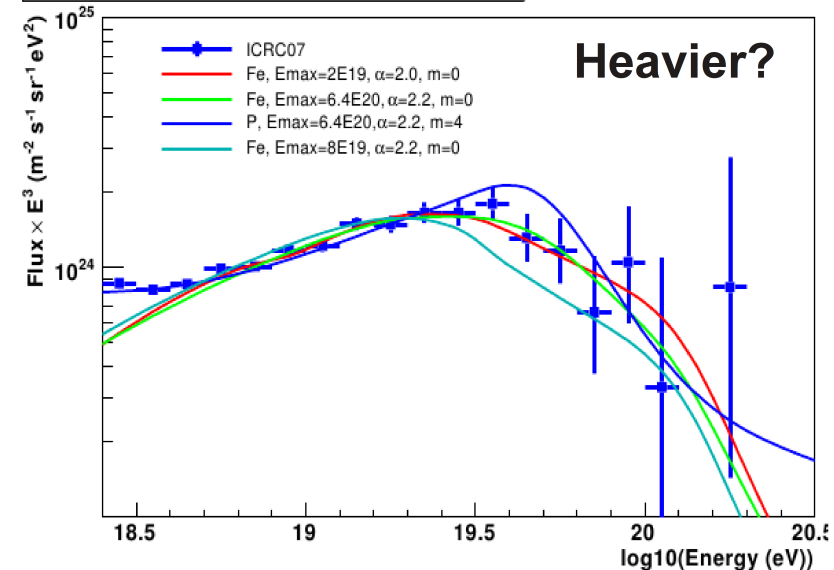


UHECRs



Evidence of GZK E, EeV cutoff

Energy Spectrum ($dF/dE \cdot E^3$ vs $\log(E)$)



Arisaka et al. 2007, JCAP, 12

e.g. *talks*: E. Kido, H. Takami, F. Oikonomou

Intro

Neutrino production connected with UHECRs

UHECRs & Neutrinos

UHECR source - a simple scenario:
only p+ only pγ interactions + neutron escape

At Δ-resonance $p + \gamma \rightarrow \pi^0 + p$ 2/3 cases

$p + \gamma \rightarrow \pi^+ + n$ 1/3 cases

$$\pi^0 \rightarrow \gamma + \gamma$$

$$\pi^+ \rightarrow \mu^+ + \nu_\mu$$

$$\mu^+ \rightarrow e^+ + \nu_e + \bar{\nu}_\mu$$

GRBs as a candidate

Neutron decay $n \rightarrow p + e^- + \bar{\nu}_e$

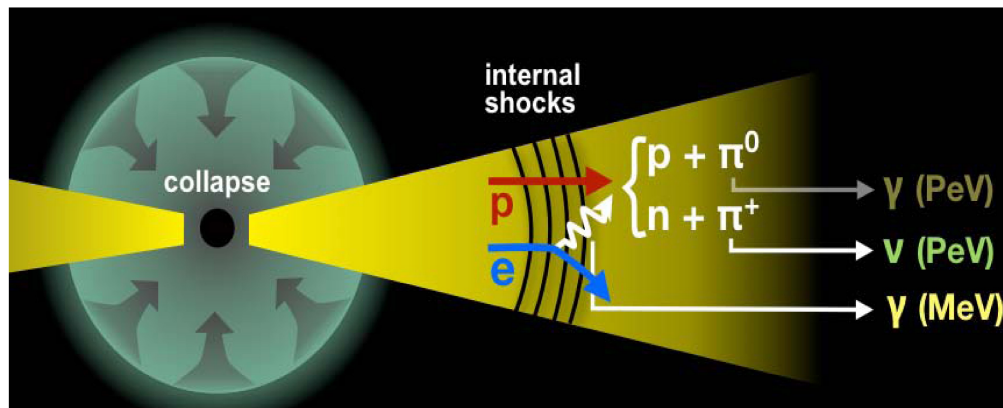
$p + CMB \rightarrow \dots$ cosmogenic ν

$$(\nu_e : \nu_\mu : \nu_\tau) = (1 : 2 : 0)$$

Motivation & Goals

GRBs: a good candidate

Method



Results

$$(1) \quad t_{acc} \leq \min(t_{loss}, t_{dyn})$$

$$t_{acc} \approx k E / ZeBc$$

(2) local injection rate at

$$10^{19} - 10^{21} \text{ eV}$$

$$E_{CR} \frac{dN_{CR}}{dE_{CR}} \approx 10^{44} \frac{\text{erg}}{\text{Mpc}^3 \text{ yr}}$$

Implications

Intro

Two main approaches

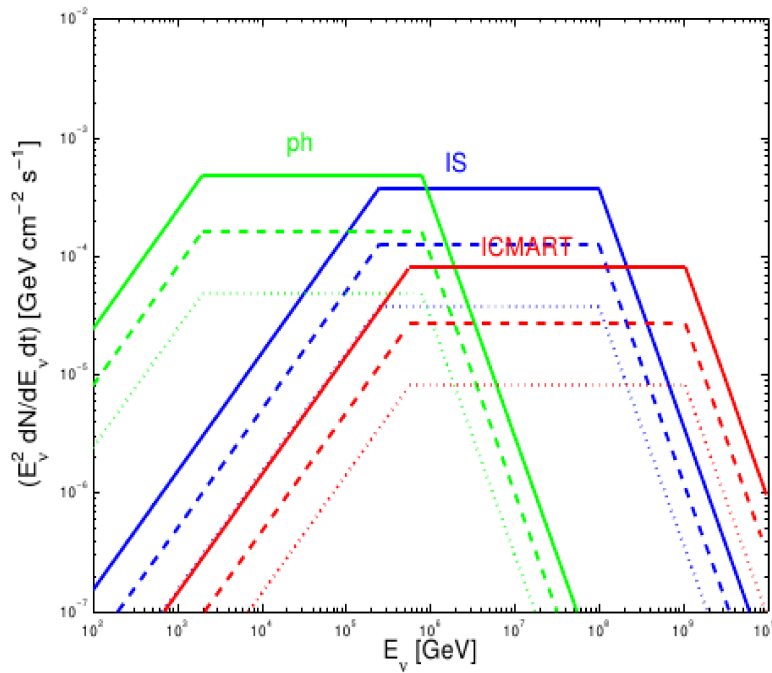
UHECRs & Neutrinos

- Use: "fixed" photon targets (MeV)
- Focus on: ν emission
- Neglect: emission from cascade (e.g. Guetta et al. 2004, Abbasi et al. 2010, Zhang & Kumar 2013)

GRBs as a candidate

- Use: "fixed" photon targets (MeV)
- Focus on: emission from the cascade (e.g. Boettcher & Dermer 1998, Dermer & Atoyan 1996, Asano & Inoue 2007, Asano & Meszaros 2014)

Motivation & Goals



Credit: Zhang & Kumar 2013, PhRvL, 110

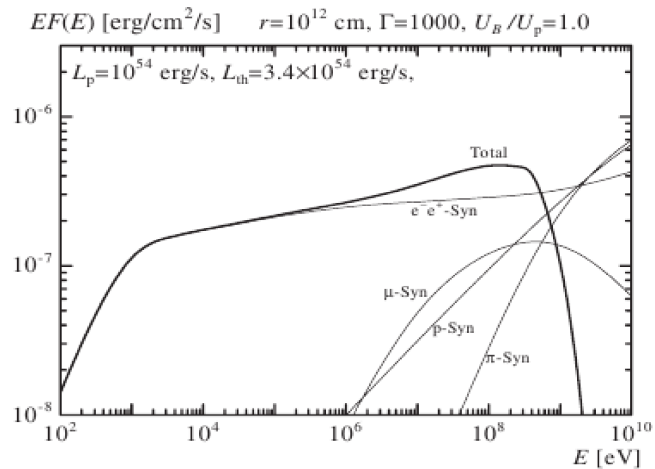
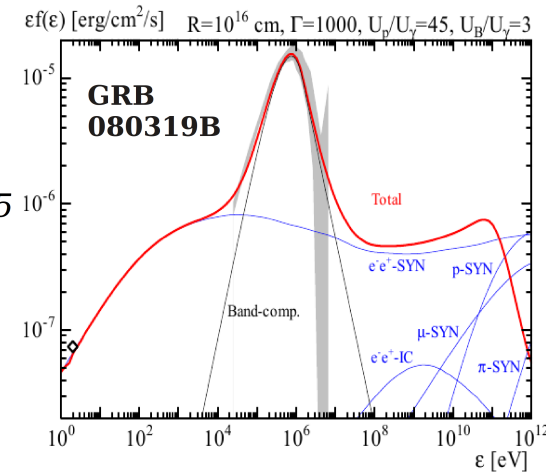
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(e.g. Boettcher & Dermer 1998, Dermer & Atoyan 1996, Asano & Inoue 2007, Asano & Meszaros 2014)

Right: Asano+2010, ApJ, 725
Bottom: Murase+ 2012, ApJ, 746



Intro

UHECRs &
Neutrinos

UHECR
sources

GRBs as a
candidate

Motivation & Goals

Motivation: Try to combine both approaches

Goals:

- Find a generic framework
- Find the parameter space where:
the contribution of the hadronic cascade (HC) to the “fixed”
photon field is important
- Derive a maximum value of the “ p/γ ” ratio (e.g. $\eta=L_p/L_\gamma$)
- Derive the ν production efficiency for this maximum η .
- Compare with results from the literature



Method

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What follows can be found in:

Petropoulou, 2014, accepted by MNRAS (arXiv:1405.7669)

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I assume:

- a “fixed” photon field for MeV emission (Band-function)
- proton acceleration to UHE with power-law distribution of slope $p=2$
- injection of protons in a region of size $R_b \sim r/\Gamma$ with magnetic field B

Point of differentiation:

Use as main parameters the photon and proton **compactnesses**

$$l_\gamma = \frac{\sigma_T L_\gamma}{4\pi R_b \Gamma^4 m_e c^3}$$
$$l_p = \frac{\sigma_T L_p}{4\pi R_b \Gamma^4 m_p c^3}$$

- Intrinsic quantities for the description of a leptohadronic system
- The efficiency, the cooling etc depend on them
- Different combinations of observables in GRB models lead to the same l_γ, l_p

Intro

Numerical method: Kinetic equation approach

(Dimitrakoudis et al. 2012)

UHECRs & Neutrinos

Protons:

$$\frac{\partial n_p}{\partial t} + L_p^{BH} + L_p^{photopion} + L_p^{psyn} + \frac{n_p}{t_{p,esc}} = Q_p^{inj} + Q_p^{photopion}$$

GRBs as a candidate

Electrons:

$$\frac{\partial n_e}{\partial t} + L_e^{syn} + L_e^{ics} + L_e^{ann} + L_e^{tpp} + \frac{n_e}{t_{e,esc}} = Q_e^{ext} + Q_e^{BH} + Q_e^{\gamma\gamma} + Q_e^{photopion} + Q_e^{tpp}$$

Motivation & Goals

Photons:

$$\frac{\partial n_\gamma}{\partial t} + \frac{n_\gamma}{t_{\gamma,esc}} + L_\gamma^{\gamma\gamma} + L_\gamma^{ssa} = Q_\gamma^{syn} + Q_\gamma^{psyn} + Q_\gamma^{ics} + Q_\gamma^{ann} + Q_\gamma^{photopion}$$

Method

Neutrinos:

$$\frac{\partial n_\nu}{\partial t} + \frac{n_\nu}{t_{esc}} = Q_\nu^{photopion}$$

Results

Neutrons:

$$\frac{\partial n_n}{\partial t} + L_n^{photopion} + \frac{n_n}{t_{esc}} = Q_n^{photopion}$$

Implications

Pion, muon & kaon decay is modeled using results of MC code SOPHIA (Muecke et al. 2000)

Synchrotron cooling of the above is also included.

Courtesy of S. Dimitrakoudis

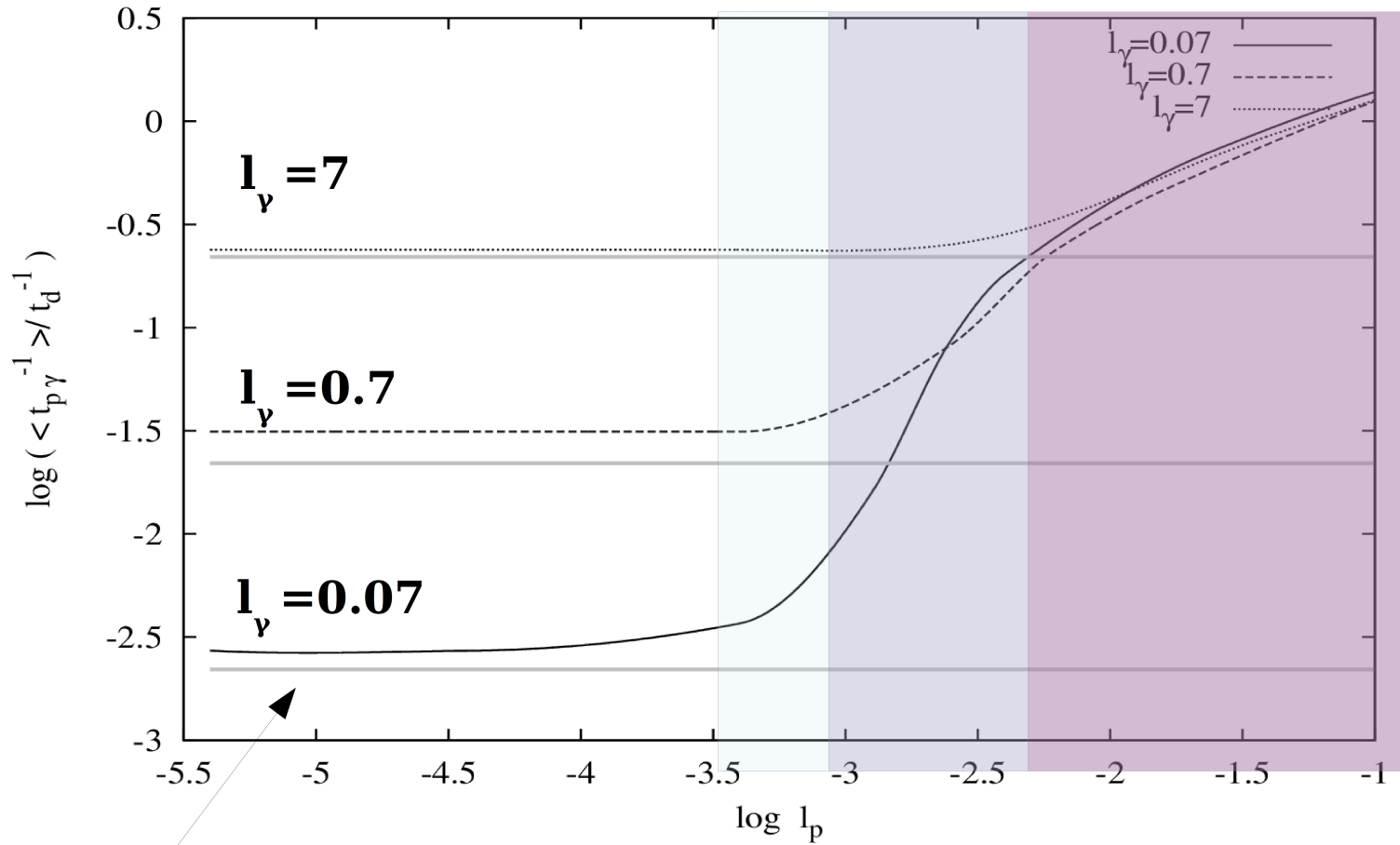
Intro

Indirect manifestation of hadronic cascade

UHECRs & Neutrinos

Proton cooling rate due to $p\gamma$ interactions

GRBs as a candidate



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Analytical expression - neglecting HC emission



Set up of cascade emission

Direct manifestation of hadronic cascade

Intro

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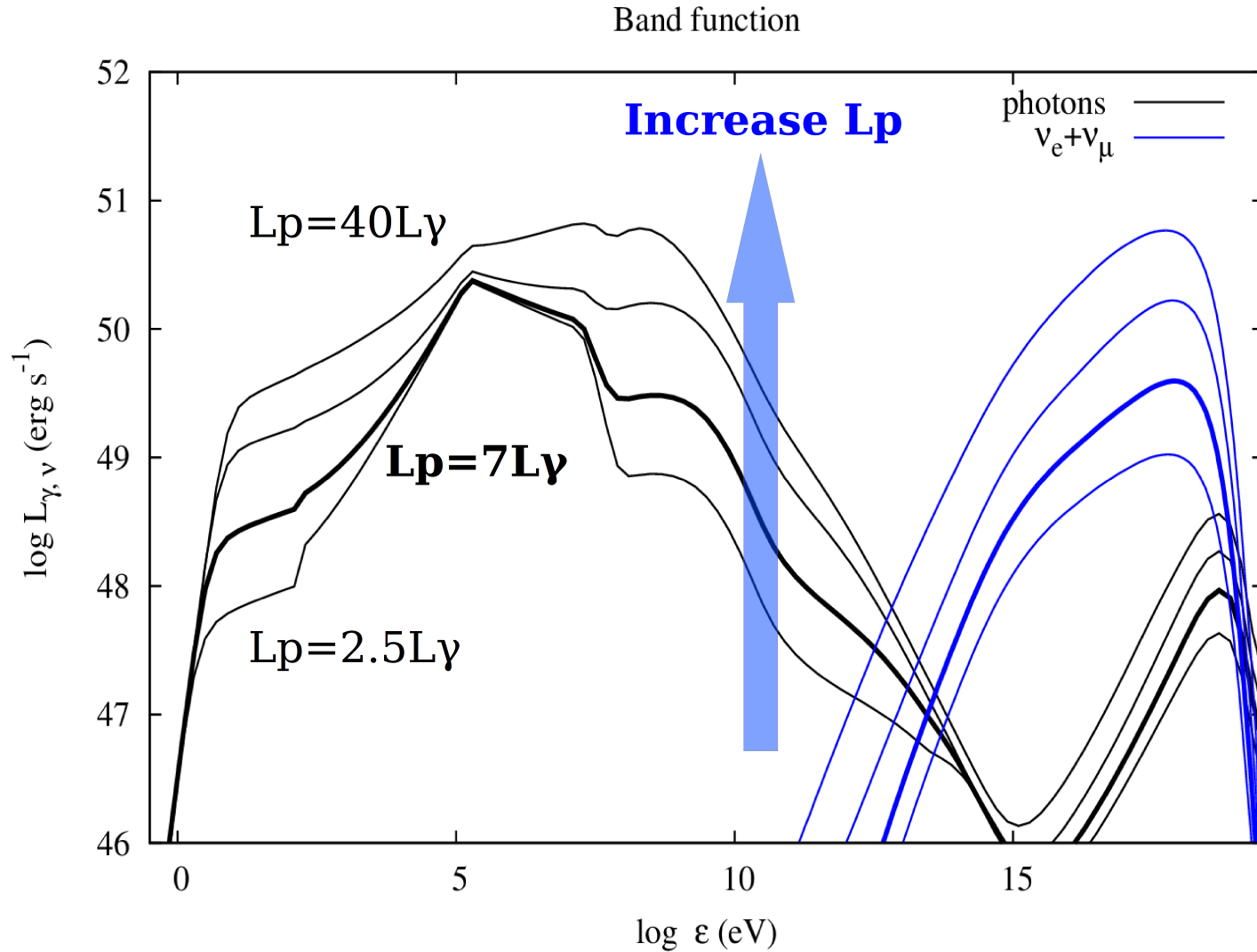
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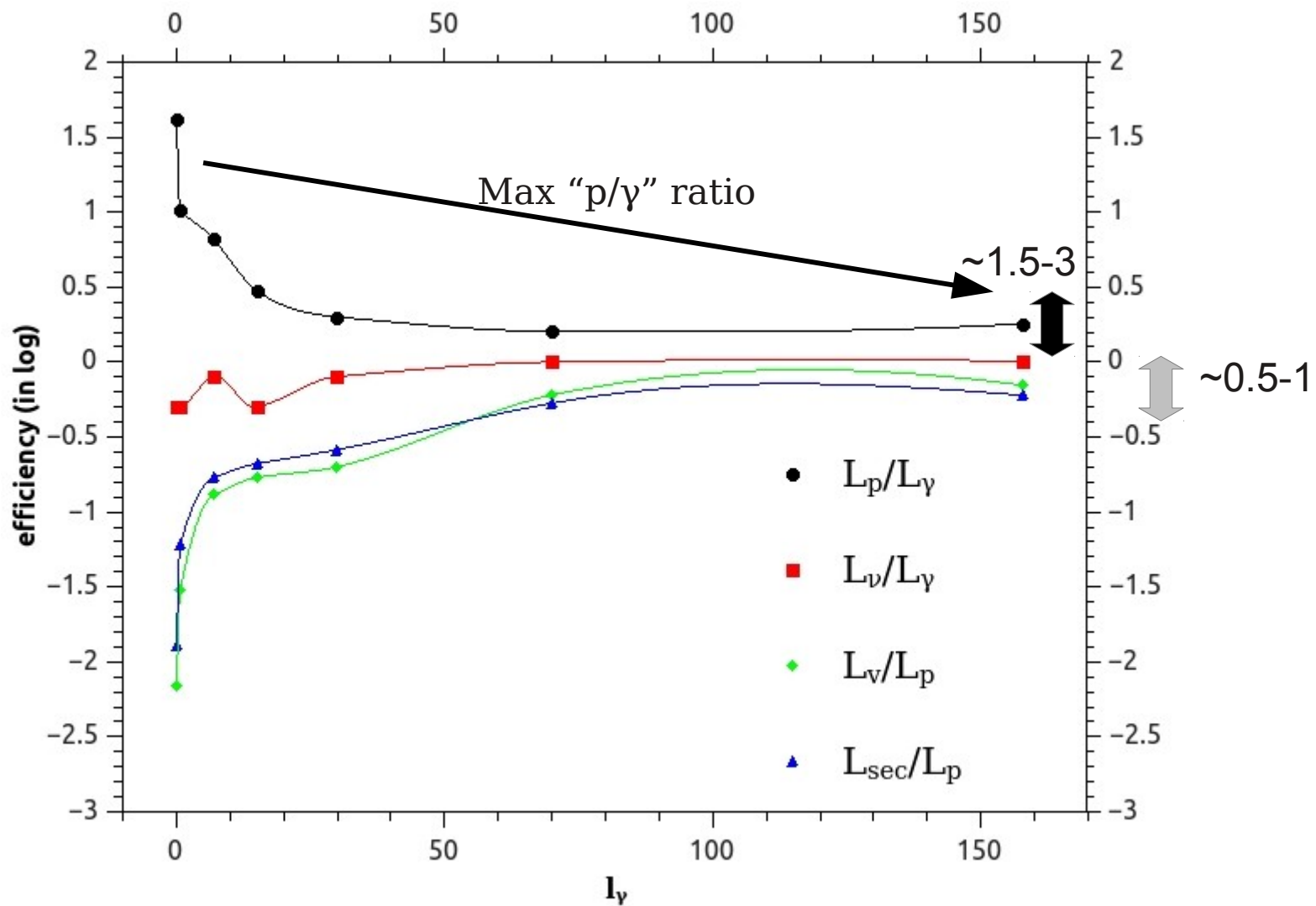
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Implications



$$l_\gamma = 0.7, L_\gamma = 1e51 \text{ erg/s } (\Gamma = 225, B = 960\text{G}, R_b = 1e12\text{cm})$$

Maximum Efficiencies: Just before the HC dominance



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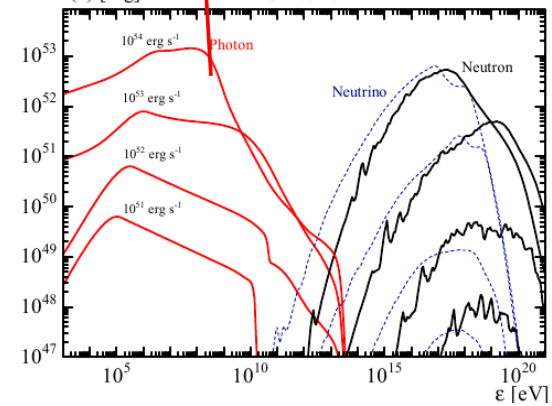
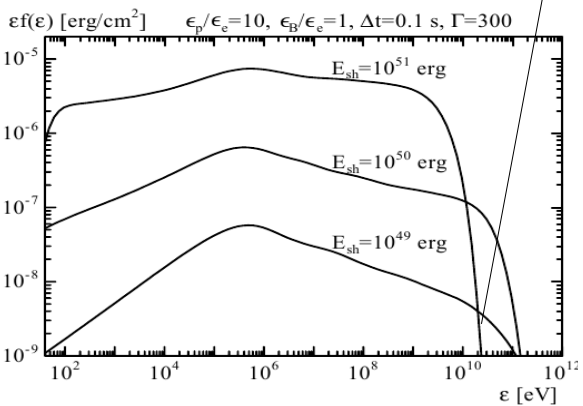
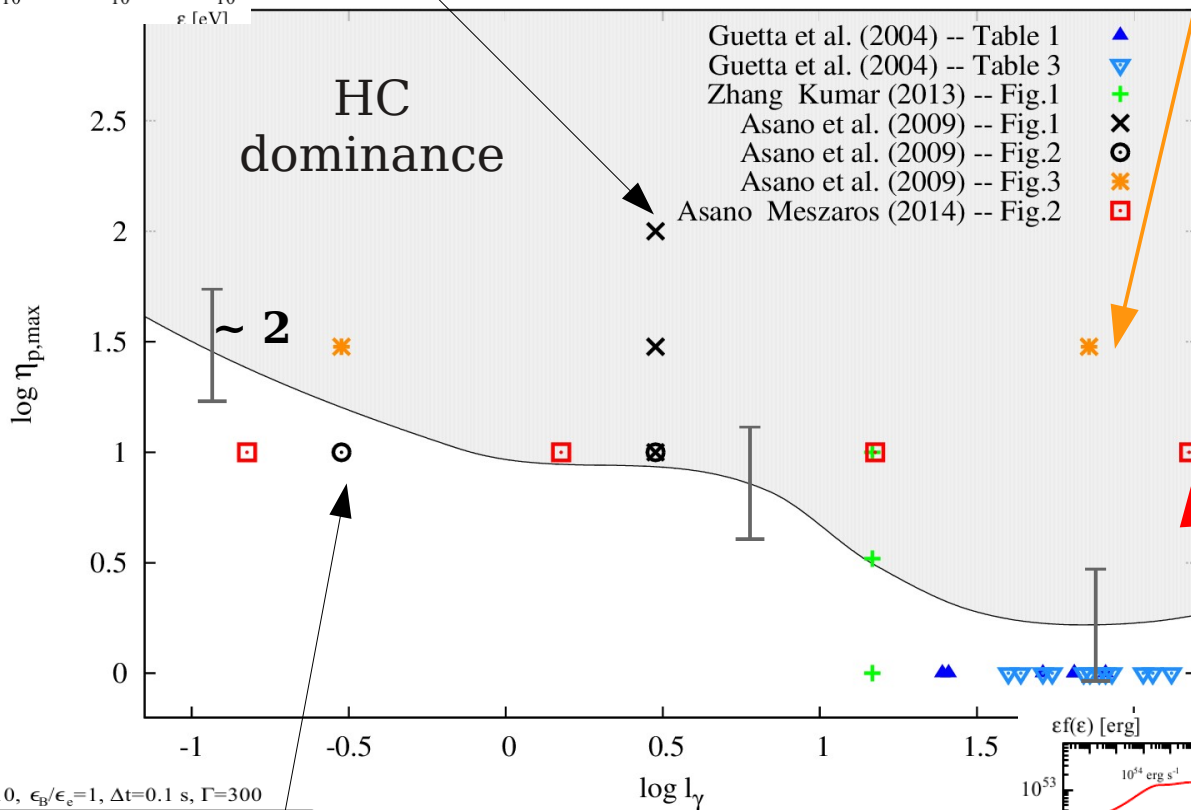
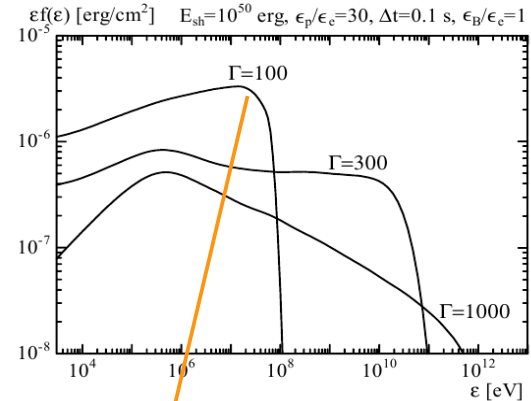
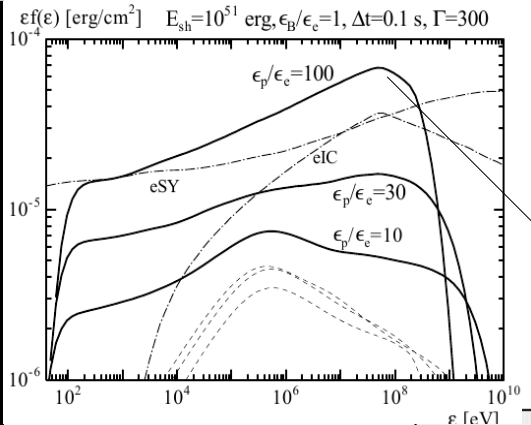
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The $\eta_{p,max} - l_\gamma$ plane comparison with other works



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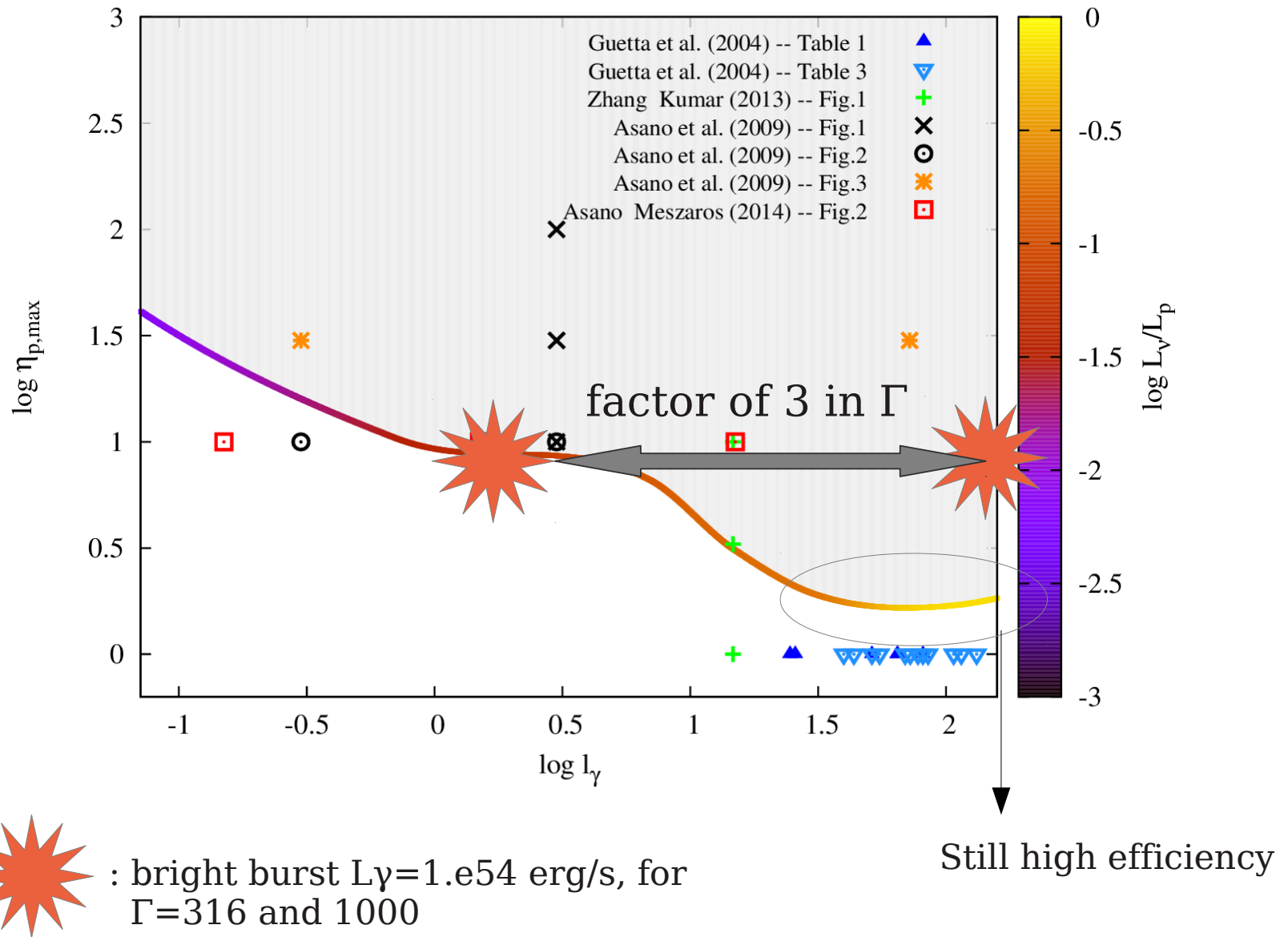
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The $\eta_{p,max} - l_{\gamma}$ plane ν efficiency



Summary

- Hadronic cascade (HC) becomes dominant for high enough L_p/L_γ and modifies the *pre-assumed* GRB photon emission
- The γ -ray and proton compactnesses are *key* parameters of the study
- Different combination of GRB observables lead to the same γ - and p -compactnesses
- The maximum “ p/γ ” ratio ($\eta_{p,\max}$) is a decreasing function of the γ -ray compactness
- Even for the maximum “ p/γ ” ratio the ν efficiency may reach $\sim 60\%$ (for high l_γ)
- The plane $\eta_{p,\max} - l_\gamma$ is a useful tool for studying HC in GRB emission

