

# How Quantum are the cosmological perturbations?

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# Why do we care?

Primordial perturbations have a **quantum** origin : correlations are quantum.

Our observations are **classical** : we got a set of classical *probability distribution functions* pdf

$$P_{cl} = \langle a_{lm} a_{lm}^* \rangle$$

*Is there a way to test for the quantum origin of perturbations?*

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**Example** : 2 entangled qubits store 2 bits of info.

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**David Tong's Nightmare** : The CMB has quantum correlations telling us the Secret of M-Theory but humanity stupidly built Planck and made measurements which loses this information.



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All is not lost! If a system has quantum correlations, then it doesn't obey classical correlation statistics -- we can check! (e.g. **Bell's Inequality**.)

General theorem (**CHSH inequality**) :

***Quantum information cannot be represented by a local joint prob. distribution function.***

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General theorem (**CHSH inequality**) :

***Quantum information cannot be represented by a local joint prob. distribution function.***

*Entanglement is **not** the only measure of quantumness!*

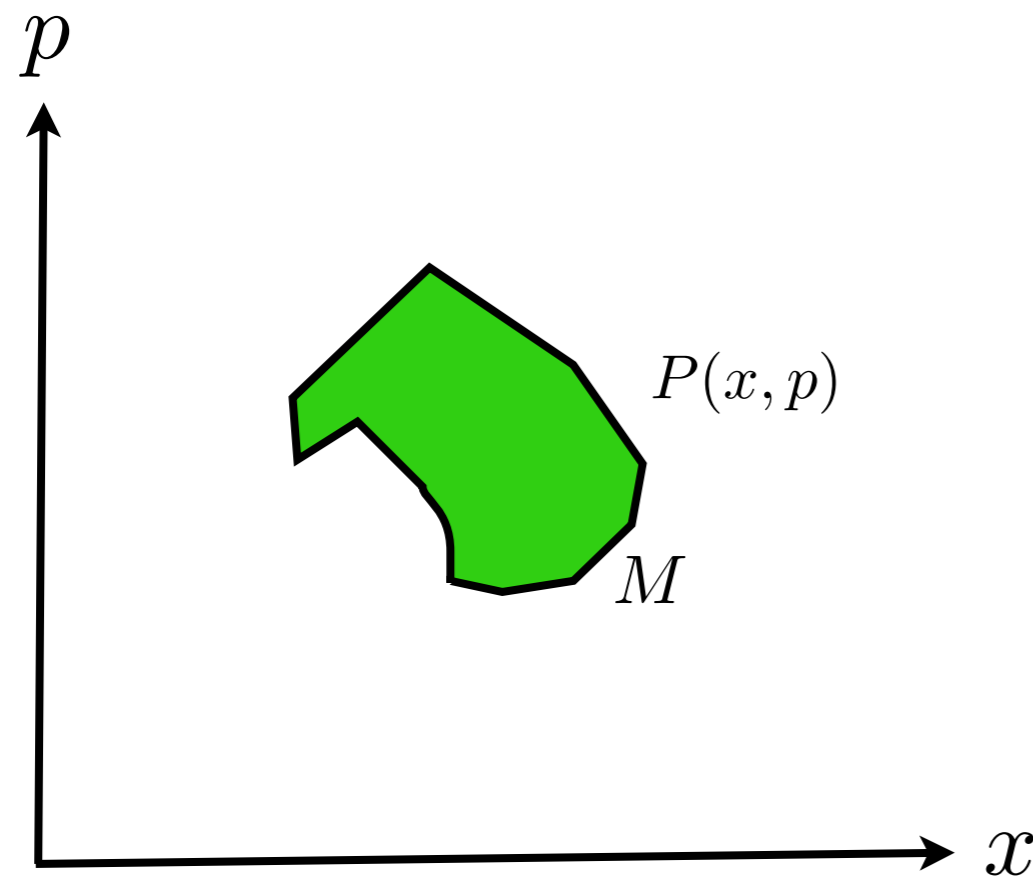
To construct such a statistic, we need to know the nature of the quantum correlations.

# Classical vs Quantum States

**Classical States** : described by *joint probability distribution functions* (pdf) of observables  $P(x, p)$

Probability of finding particle in region  $M$

$$\text{Prob} = \int_M dx dp P(x, p)$$



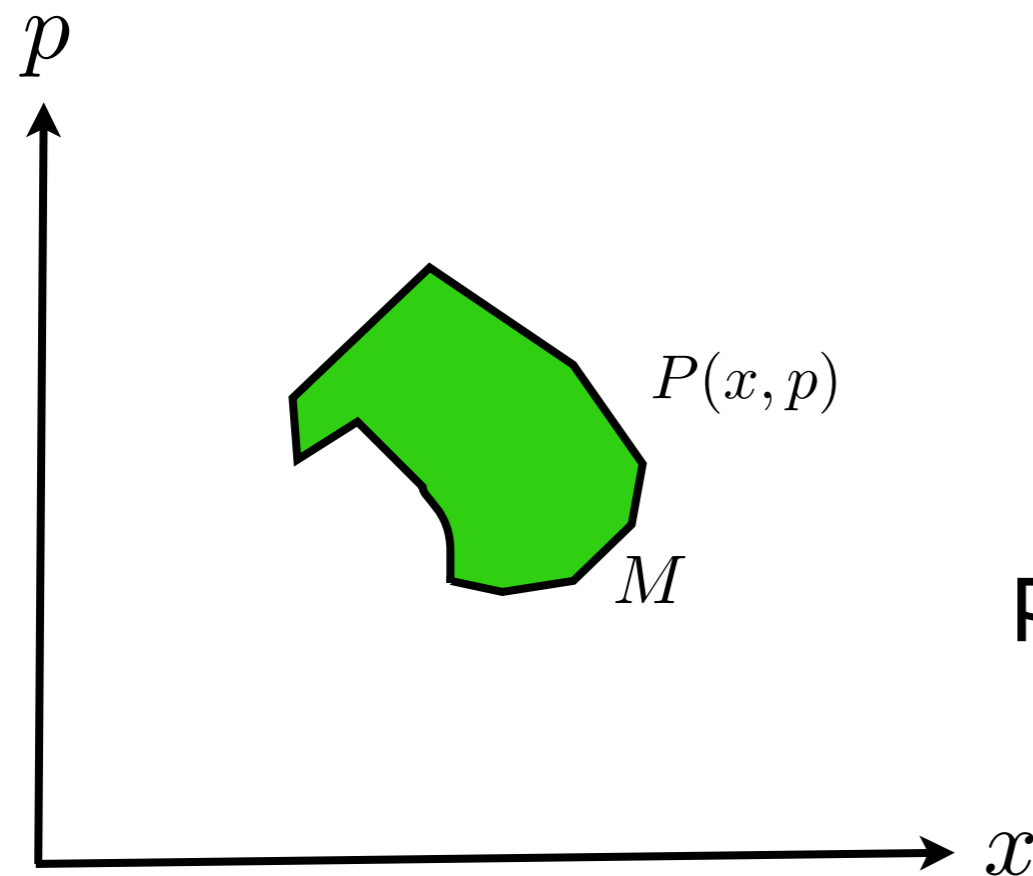
Phase space for single particle state  $(x, p)$

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Probability = not sure where the particle is = “ambiguity” = entropy

Boltzmann/Shannon Entropy

$$H(P) = - \int dx P(x) \log P(x)$$



# Classical vs Quantum States

**Quantum States** : described by *density matrices*  $\rho$

A state vector  $|u_i\rangle$  describes a *pure state*.

$$\rho = \sum_i p_i |u_i\rangle \langle u_i| \quad \sum_i p_i = 1 \quad \text{A mixture of pure states } |u_i\rangle$$

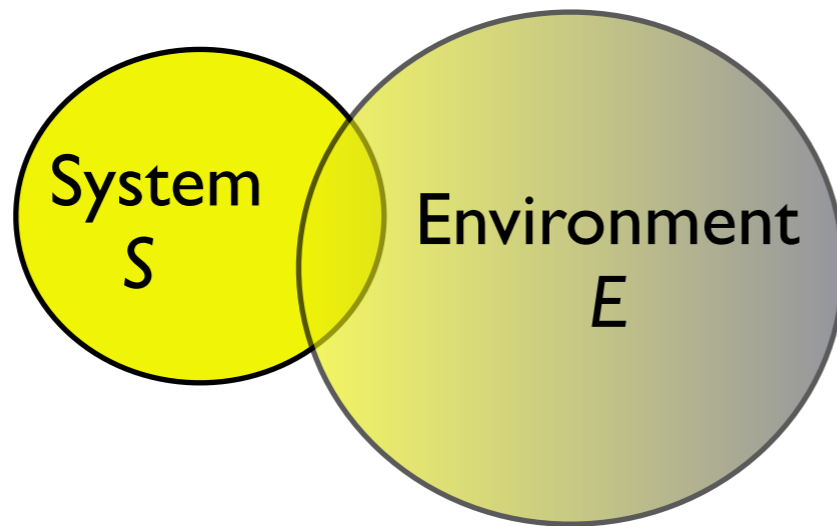
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Pure states  $\rho = |u\rangle \langle u|$  can evolve into *mixed states* under non-unitary operations in Open systems.



Combined “Bipartite”  $\rho_{SE}$

Access only to S :  $\rho_S = \text{Tr}_E \rho_{SE}$

“Ambiguity” = Von Neumann Entropy  $S(\rho_S) = -\text{Tr}(\rho_S \log \rho_S)$

# Classical vs Quantum States

**Quantum States** : described by *density matrices*  $\rho$

Given bipartite system, it is *separable* if

$$\rho = \sum_i p_i |u_i\rangle_S |e_i\rangle_E \langle e_i|_S \langle u_i|_E$$

example  $\rho = \frac{1}{2} |0_S\rangle |0_E\rangle \langle 0_S| \langle 0_E| + \frac{1}{2} |1_S\rangle |1_E\rangle \langle 1_S| \langle 1_E|$

**Pure states** : *separability* = non-entanglement  
= classical pdf.

**Mixed states** : *separability* = non-entanglement  
 $\neq$  classical pdf (*quantum discord*)

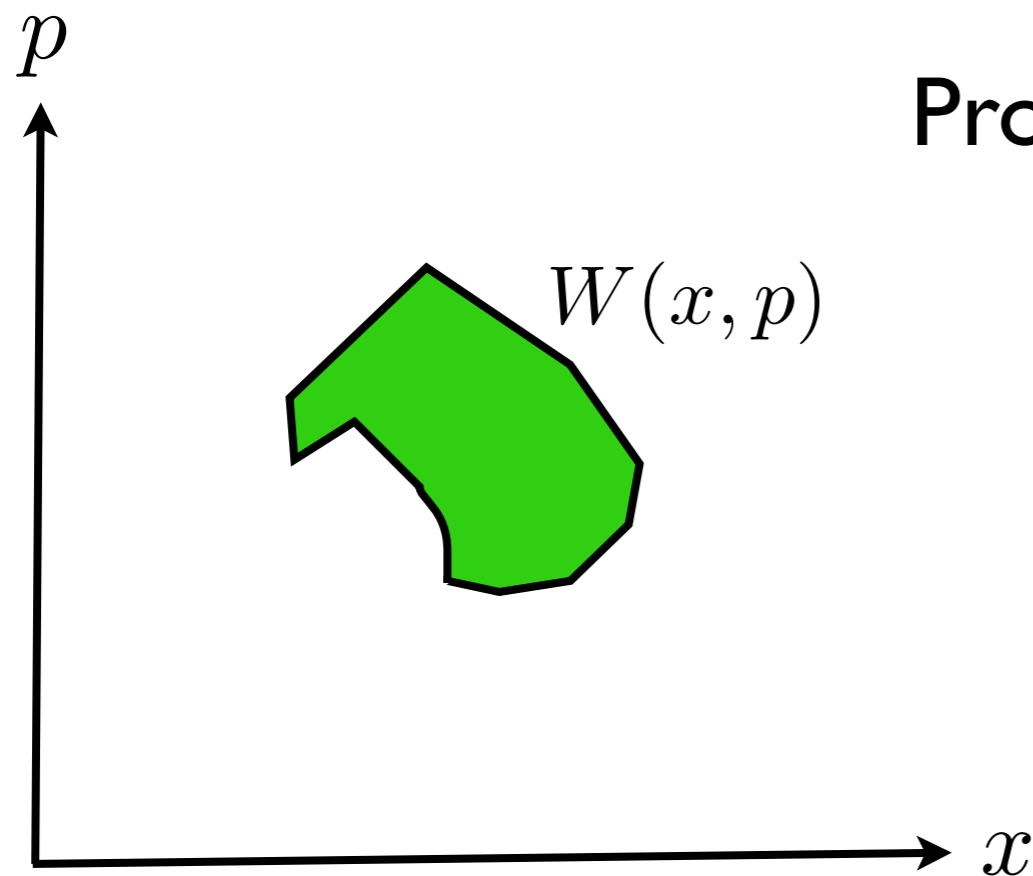
# Classical vs Quantum States

Equivalent “quasi-pdf” picture : **Wigner distribution**

$$W(x, p) = \frac{1}{\pi} \int_{-\infty}^{\infty} e^{2ipy} \langle x - \frac{y}{2} | \rho | x + \frac{y}{2} \rangle$$

Prob density of  $x$  is then

$$\langle x | \rho | x \rangle = \frac{1}{2\pi} \int_{-\infty}^{\infty} dp W(x, p)$$

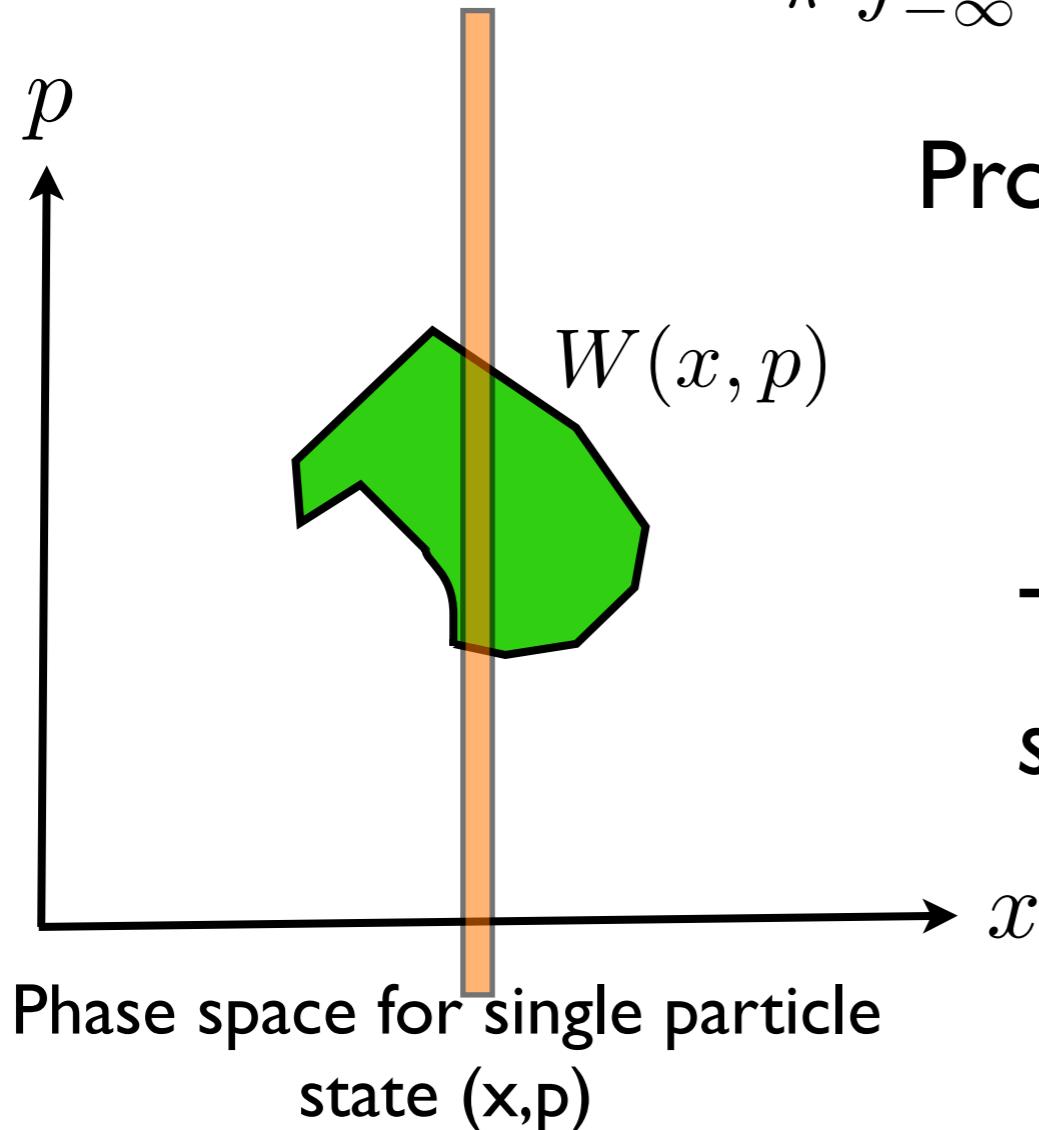


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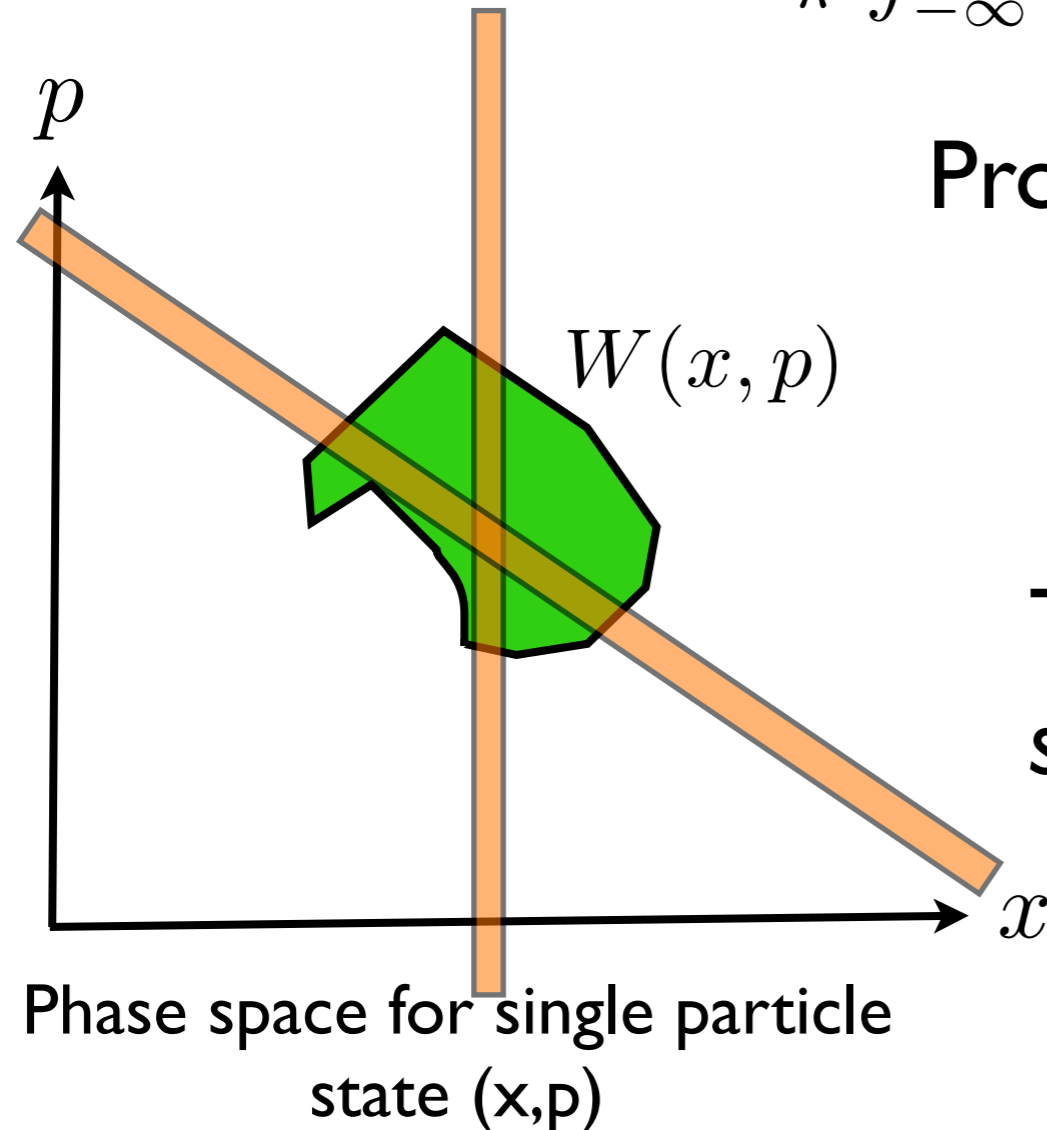
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This is an integration over *an infinite strip of  $p$*  (uncertainty principle).

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*Bertrand and Bertrand's theorem : any infinite strip would do.*

# Decoherence in a nutshell

Consider pure state  $|S\rangle = \alpha|0\rangle + \beta|1\rangle$

**Coherence** = quantum phase of  $\alpha$  and  $\beta$  preserved.

$$\begin{aligned}\rho = |S\rangle\langle S| &= |\alpha|^2|0\rangle\langle 0| + |\beta|^2|1\rangle\langle 1| + \alpha\beta^*|0\rangle\langle 1| + \alpha^*\beta|1\rangle\langle 0| \\ &= \begin{pmatrix} |\alpha|^2 & \alpha\beta^* \\ \alpha^*\beta & |\beta|^2 \end{pmatrix}\end{aligned}$$

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**Decoherence** : couple  $S$  to environment  $E$ .

$$|S\rangle \otimes |E\rangle = (\alpha|0\rangle + \beta|1\rangle)|E\rangle \xrightarrow{\text{couplings}} \alpha|0\rangle|E(0)\rangle + \beta|1\rangle|E(1)\rangle$$

If we have only access to  $S$ , then

$$\rho_S = \text{Tr}_E \rho_{SE} = \begin{pmatrix} \rho_{00} & \rho_{01} \rightarrow 0 \\ \rho_{10} \rightarrow 0 & \rho_{11} \end{pmatrix} \rightarrow \begin{pmatrix} |\alpha|^2 & 0 \\ 0 & |\beta|^2 \end{pmatrix}$$

Final matrix is *mixed* and phase info is lost.



# Decoherence basis is crucial

Secret assumption : decoherence occurred in  $\{|0\rangle, |1\rangle\}$  basis.

If decoherence occurs at rotated basis

$$\{\cos \theta |0\rangle + e^{i\phi} \sin \theta |1\rangle, -e^{-i\phi} \sin \theta |0\rangle + \cos \theta |1\rangle\}$$

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*Quantum nature of cosmological perturbations depends on how they interact with environment and how we measure them.*

# What about inflation?

Starobinsky, Polarski (1998)

Single mode Hamiltonian for cosmological perturbations

$$\hat{H}_k = \frac{1}{2} \left( p_k^2 + k^2 y_k^2 + \frac{2a'}{a'} y_k p_k \right)$$

$$\delta\phi_k \equiv y_k p_k = \frac{\partial L(y_k, y'_k)}{\partial y'_k} = y'_k - a'/ay_k$$

$\hat{H}_k$  is a unitary evolution operator.

Schrodinger's equation of wave function  $\psi(y, \eta)$

$$i\hbar \frac{\partial \psi(y, \eta)}{\partial \eta} = \hat{H}_k \psi(y, \eta).$$

with solution  $\psi(y, \eta) = \left( \frac{2\Omega_R(\eta)}{\pi} \right)^{1/4} \exp(-(\Omega_R + i\Omega_I)y^2)$

for inflation background  $\Omega_R \rightarrow ke^{-2r}$  ,  $\Omega_I \rightarrow -ke^{-r}$

$r = \# \text{efolds}$

# Cosmological “Squeezed states”

Construct density matrix

$$\rho_S = |\psi\rangle\langle\psi| = \frac{2\Omega_R}{\pi} \exp \left[ -\left(\frac{\Omega_R}{2}(y - y')\right)^2 - \frac{\Omega_R}{2}(y + y')^2 - i\Omega_I(y^2 - y'^2) \right]$$

**As  $\Omega_R/k = e^{-2r} \ll 1 \Rightarrow$  off-diagonal terms get killed off**

**So far unitary evolution : *no decoherence so still pure state.***

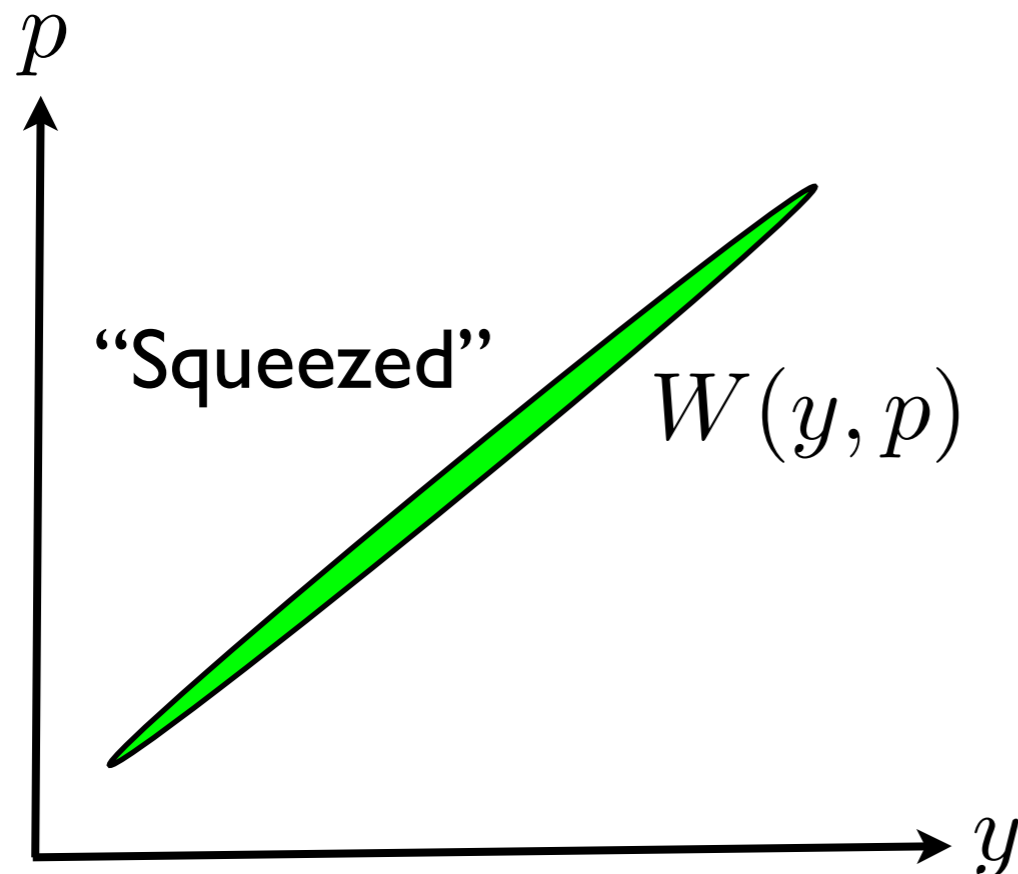
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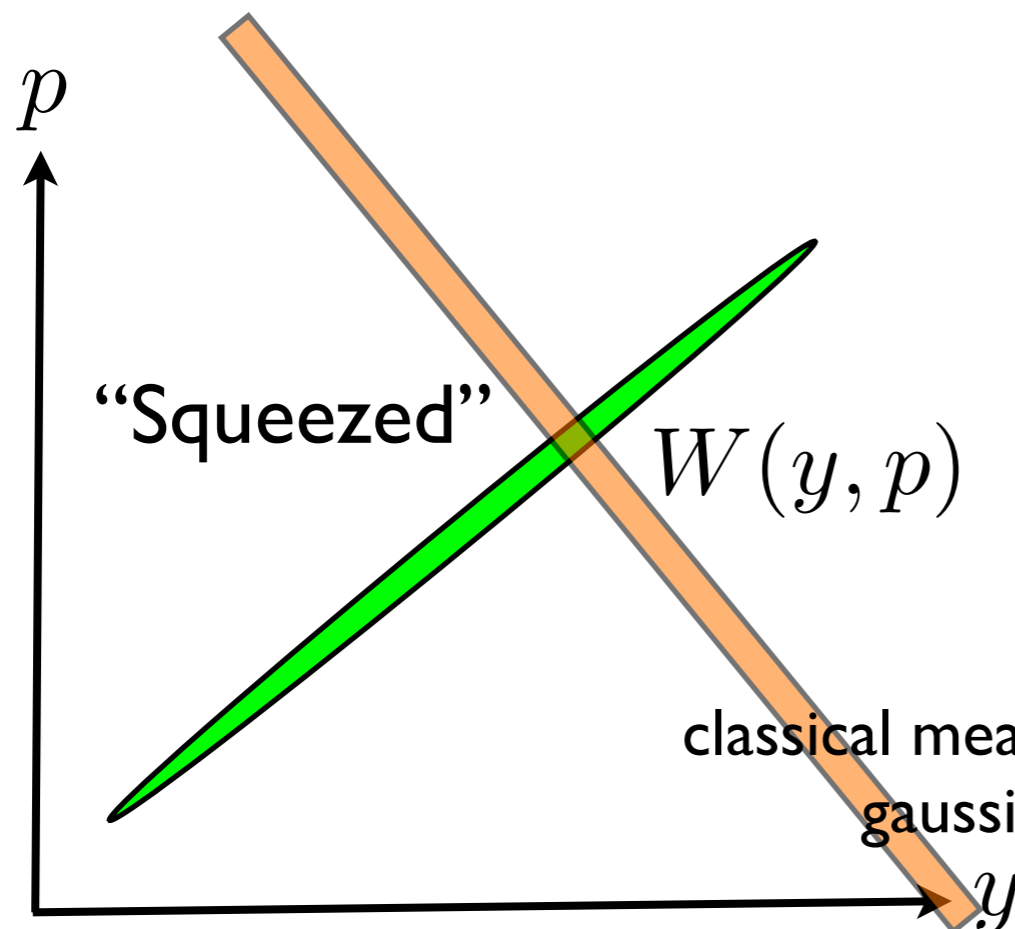
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classical measurements =  
gaussian pdf

# Cosmological “Squeezed states”

*Starobinsky, Kiefer and Polarski’s decoherence ansatz (1998)*

Couple to environment  $\rho_{SE}$ , we add a decoherence term

$$\rho'_S = \text{Tr}_E \rho_{SE} = \rho_S \times \exp\left[-\frac{\zeta}{2} (y - y')^2\right] \quad \zeta \gg \Omega_R$$

New *mixed state* is still a gaussian but with axes  $(e^r, \zeta)$

Question 1 : is the decoherence basis parallel to  $\{y_k\}$ ?

Question 2 : how to quantify “quantumness”?

We will model  $\rho_{SE}$  and use **quantum discord** to answer both questions.

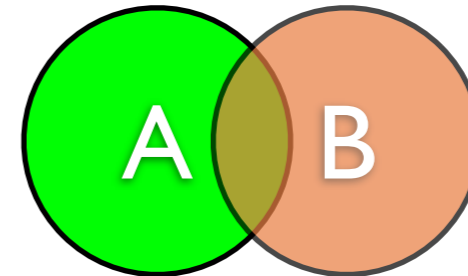
Ollivier and Zurek (2001)  
Henderson and Vedral (2001)

# Quantum Discord

Ollivier and Zurek (2001)

## Classical Mutual Information

$$J(A : B) = H(A) - H(A|B)$$



A and B correlated, mutual info is how much we learn more about A when B is found out.

Classical pdf, **Bayes Theorem**  $H(A|B) = H(A, B) - H(B)$

Get equivalent expression

$$I(A : B) = H(A) + H(B) - H(A, B)$$

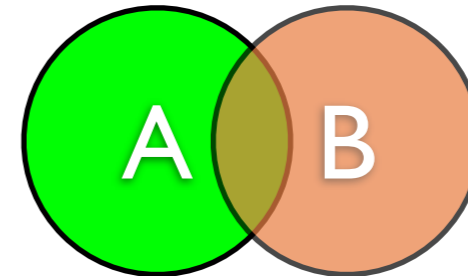


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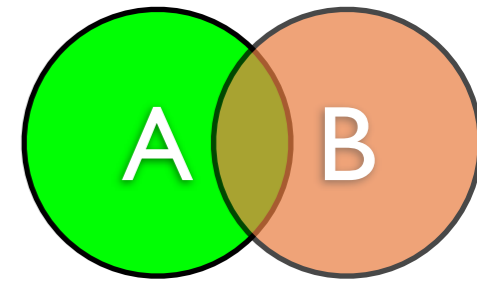
Quantum generalization: replace Shannon with Von

Neumann  $H(A) \rightarrow S(A)$

$$I(A : B) \rightarrow \mathcal{I}(A : B) = S(A) + S(B) - S(A, B)$$

$$J(A : B) \rightarrow \mathcal{J}(A : B) = S(A) - S(A|B)$$

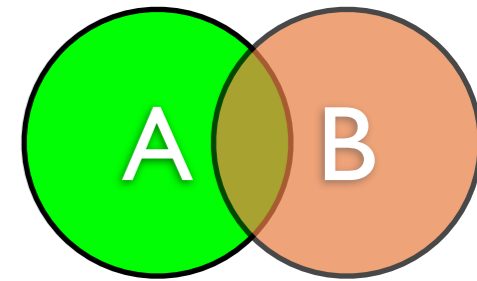
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Ollivier and Zurek (2001) propose the following:

1. Given some basis of measurements on B  $\{\Pi_k^B\}$
2. Each  $\Pi_k^B$  measurement occurs with prob.  $P_k$  and

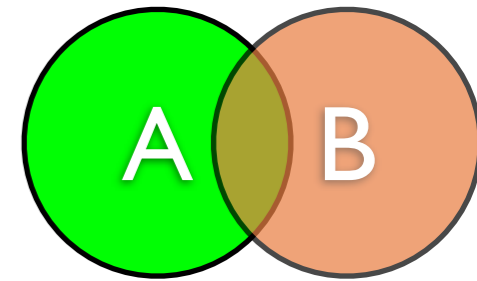
$$\rho_{AB} \rightarrow \rho_{A|B=\Pi_k^B} = \frac{\rho_{AB}\Pi_k^B}{P_k}$$

3. Define  $S(A|B = \{\Pi_k^B\}) = \sum_k P_k S(\rho_{A|B=\Pi_k^B})$

4. Quantum Discord is then

$$\delta(A : B)_{\Pi_k^B} = \mathcal{I}(A : B) - \mathcal{J}(A : B)_{\Pi_k^B} > 0$$

# Quantum Discord



Some facts on Discord :

1. Zero discord  $\delta(A : B)_{\Pi_k^B} = 0$  means decoherence occurred in “pointer basis” and no entanglement. Can define to be “Classical”.
2. Mixed Separable state can have *non-zero* discord. No entanglement  $\neq$  no quantum correlations!
3. Basis-independent discord : minimize over all possible decoherence basis.
4. Recently shown separable 2 qubits computers with discord is exponentially faster than classical computers.

Datta, Shaji, Caves (2007)

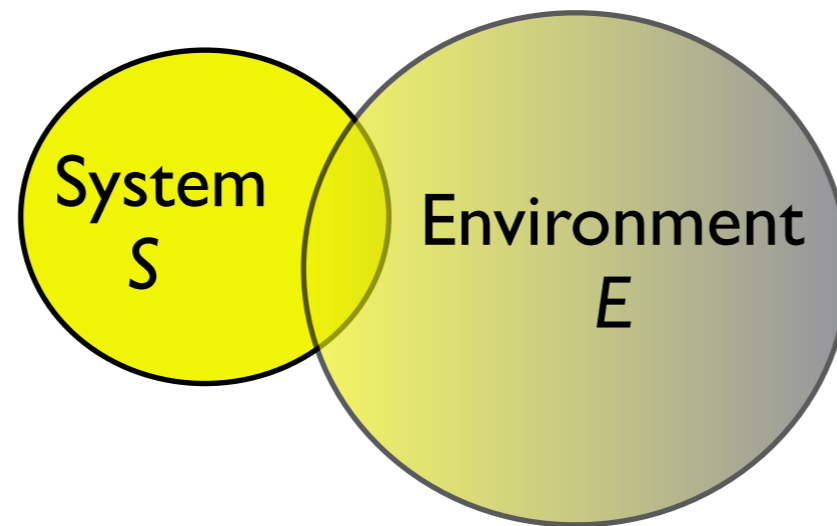
# How to model system-environment?

Start with Starobinsky-Polarski-Kiefer Gaussian ansatz

$$\rho'_S \rightarrow W(y, p) = \frac{1}{\pi} \exp \left[ -\frac{1}{2} \mathbf{x} \sigma_S^{-1} \mathbf{x}^T \right], \quad \sigma_S(\Omega_R, \Omega_I, \zeta)$$

We want to find a joint density matrix  $\rho_{SE}$  such that

$$\rho_S = \text{Tr}_E(\rho_{SE})$$



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Not unique and continuous states are HARD. A way forward is to assume that  $\rho_{SE}$  is also Gaussian.

Central Limit  
Theorem



# How to model system-environment?

A **unique** pure gaussian  $\rho_{SE}$  can be constructed (“gaussian purification” of  $\rho_S$ ).

$$W(y_1, y_2, p_1, p_2) = \frac{1}{\pi} \exp \left[ -\frac{1}{2} \mathbf{x} \sigma_{SE}^{-1} \mathbf{x}^T \right] \quad \mathbf{x} = (y_1, p_1, y_2, p_2)$$

$$\sigma_{SE} = \begin{pmatrix} \frac{1}{\lambda} & \frac{a}{\lambda} & \frac{\sqrt{\zeta}}{\lambda^{3/4} \sqrt[4]{\zeta+\lambda}} & 0 \\ \frac{a}{\lambda} & \frac{a^2}{\lambda} + \zeta + \lambda & \frac{a\sqrt{\zeta}}{\lambda^{3/4} \sqrt[4]{\zeta+\lambda}} & -\sqrt{\zeta} \sqrt[4]{\frac{\zeta+\lambda}{\lambda}} \\ \frac{\sqrt{\zeta}}{\lambda^{3/4} \sqrt[4]{\zeta+\lambda}} & \frac{a\sqrt{\zeta}}{\lambda^{3/4} \sqrt[4]{\zeta+\lambda}} & \sqrt{\frac{\zeta+\lambda}{\lambda}} & 0 \\ 0 & -\sqrt{\zeta} \sqrt[4]{\frac{\zeta+\lambda}{\lambda}} & 0 & \sqrt{\frac{\zeta+\lambda}{\lambda}} \end{pmatrix}$$

$$a = -\Omega_I/k, \quad \lambda = \Omega_R/k, \quad \xi = \zeta/k$$

# How to model system-environment?

A **unique** pure gaussian  $\rho_{SE}$  can be constructed (“gaussian purification” of  $\rho_S$ ).

Perturb around  $\rho_{SE}$  to obtain general mixed states.

Given pure state bipartite  $\rho_{SE}$  we can compute the discord (turns out to be basis independent)

$$\delta(A : B) = \frac{1 + \sqrt{1 + \chi}}{2} \log \left( \frac{1 + \sqrt{1 + \chi}}{2} \right) - \frac{\sqrt{1 + \chi} + 1}{2} \log \left( \frac{\sqrt{1 + \chi} - 1}{2} \right), \quad \chi = \frac{\zeta}{\Omega_R} \gg 1$$

Zero when  $\zeta = 0$  so the non-decohered perturbations has classical statistics!



# How to model system-environment?

This is actually equal to the Von Neumann entropy

$$\delta(A : B) = S(\rho_S)$$

(Kiefer, Starobinsky and Polarski 1999)

Reason :  $\rho_{SE}$  is pure, and discord captures *mixed* state quantum correlations beyond entanglement entropy.

Still work in progress : mixed  $\rho_{SE}$

# Summary

1. Environment picks out the basis of which we measure quantum cosmological correlations.
2. Even if cosmological perturbations are highly squeezed, measurements in off-basis may retain “quantum” correlations.
3. Propose a gaussian construction of joint  $\rho_{SE}$  perturbations-environment bipartite state
4. Propose **quantum discord** as robust measure of quantum correlations in the joint system.