### Black holes, baryons, axions, the Sagittarius dwarf galaxy, and dark matter

Marc Kamionkowski LSST DM Workshop, Chicago, 5 August 2019

### Dark matter!!

### What is it?

### The Party line (~1985—2015)

 Weakly Interacting Massiv Particles (WIMPS).
 e.g.,neutralinos





### Simplicity/Elegance

### **Direct/indirect** searches



Inelastic, Sommerfeld-enhanced, momentum-dependent, leptophilic,co-annihilating, dipolar, millicharged, resonant, superheavy, sub-GeV, self-interacting, atomic, dark-sector, Higgs portal,..... Inelastic, Sommerfeld-enhanced, momentum-dependent, leptophilic,co-annihilating, dipolar, millicharged, resonant, superheavy, sub-GeV, self-interacting, atomic, dark-sector.....



### The dark matter landscape now

### **Dark-matter candidates**



# How each 2020 Democratic presidential candidate could win dark-matter

Literally all of them.



#### I. ~30-Msun black hole dark matter?

- II. Millicharged DM and EDGES
- III. Intensity mapping and decaying dark matter
- IV. DM, the Sagittarius dwarf galaxy, and new long-range DM interactions

### Did LIGO detect dark matter?

(Bird, Cholis, Munoz, Ali-Haimoud, Kamionkowski, Kovetz, Raccanelli, Riess, 2016)

- highly speculative; not crazy
- Surprising coincidence: If black holes of 30 solar masses make up the dark matter, they merge with rate comparable to that inferred from the initial LIGO event! (Bird et al. 2016)

### Suppose DM = 30-Msun BHs

#### Gravitational radiative recombination

$$\begin{split} \sigma &= 2^{3/7} \, \pi \left( \frac{85 \, \pi}{6 \sqrt{2}} \right)^{2/7} R_s^2 \left( \frac{v_{\rm pbh}}{c} \right)^{-18/7} \\ &= 1.37 \times 10^{-14} \, M_{30}^2 \, v_{\rm pbh-200}^{-18/7} \, {\rm pc}^2, \end{split}$$



$$\mathcal{V} = 5 f (M_c / 500 \, M_\odot)^{-11/21} \, \mathrm{Gpc}^{-3} \, \mathrm{yr}^{-1}$$

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Since then....

### Scenario ruled out (??) by:

- CMB (Ricotti, Ostriker, Mack 2007)
- Dwarf-galaxy dynamics (Brandt, 2016; Koushiappas et al. 2016)
- Quasar lensing (Mediavilla 2017)
- X rays from accretion of ISM (Gaggero et al. 2017; Inoue & Kusenko 2017)
- SN dispersions (Zumalcarregui & Seljak 2017)
- Pulsar timing (Schutz & Liu 2017)
- Good taste [[Supergravity inflation (1606.07361,1612.02529); axion inflation (1610.03763; 1704.03464); broken scale invariance (1611.06130,1702.03901);non-thermal histories (1703.04825); trapped inflation (1606.00206); double inflation (1705.06225); axion stars (0609.04724); critical Higgs inflation (0705.04861); contracting Universe (0609.02556).... ]]

### **CMB** fluctuations

Ricotti, Ostriker, and Mack (2008): heating of primordial plasma due to accretion onto PBHs leads to unacceptable fluctuations in CMB (by ~3-4 OoMs!!)



### How does the CMB probe PBHs?

- PBHs accrete primordial plasma
- Accreted gas gets heated
- Heated gas radiates
- Radiation heats plasma  $\rightarrow$  spectral distortions
- Radiation also affects ionization balance
  - $\rightarrow$  changes recombination history
    - $\rightarrow$  affects CMB power spectra

### Our work

(Ali-Haimoud&MK 2017)

- first-principles calculation
- Given many uncertainties/complications, make simplest but most robust assumptions
- seek bound, not best estimate
- Self-consistently include DM-baryon relative velocities



#### Primordial PBH binaries (Ali-Haimoud, Kovetz, MK 2017)

#### Basic idea: Nakamura, Sasaki, Tanaka & Thorne 1997



On small enough scales, PBHs are randomly distributed (or maybe not quite!)

Some PBH pairs happen to be close enough that they decouple from the Hubble flow deep in the radiation era.

As they fall towards one another, torqued by other PBHs result in a non-zero (but small) angular momentum

Inspiral through GW radiation, some merge at the present time.

### Does LIGO rule out PBH-dark matter?

*Probably* but more checks are needed



# How to test PBH DM hypothesis?

- BBH mass spectrum
- BBH eccentricity No EM/neutrino counterparts!
- Clustering with DM
- Stochastic GW background
- Lensing echoes of fast radio bursts

Given current LIGO rate, expect perhaps ~20,000 more BBH mergers in next decade!!

#### PBH binaries have high initial eccentricities:



#### see many more modes of grav. waves

#### ~1 such event in LIGO; ~10 in Einstein Telescope



Cholis, Ali-Haimoud, Bird, Munoz, MK, Kovetz, and Raccanelli (2016) The BH binary mass distribution

#### The Black-Hole Mass Function from GWs

#### with 5 years of aLIGO:



#### The Black-Hole Mass Function from GWs

#### with 5 years of aLIGO data:

With Dark Matter PBHs:

Kovetz, Cholis, Breysse, MK 2017; Kovetz, 2017



### Lensing of Fast Radio Bursts by Compact Objects

Munoz, Kovetz, Dai, MK, 1605.00008

- FRBs = <millisecond ~GHz radio bursts
- ~10,000 on sky per day
- Large dispersion measures imply cosmological distances
- Forthcoming experiments (e.g., CHIME) should detect thousands



### FRB Lensing (Muñoz, Kovetz, Dai, Kamionkowski, PRL 117 (2016))



#### Images separation (~nano-arcsec) too small to be detected, but there can be a >ms time delay



Also tried echoes in gamma-ray-burst light curves, but not with current measurements (Ji, Kovetz, MK, 2019)

- Longer pathlengths  $\rightarrow$  more lensing
- Seek GRB echo superimposed on original light curve

### Baryon—DM interactions from cosmic dawn? (EDGES: Bowman et al. Nature 2018; Barkana, Nature 2018)





#### Lots of neutral hydrogen

If DM-baryon interactions exist, they can mediate heat exchange between baryons and DM (Dvorkin, Blum, MK 2014)

$$\frac{dT_{\chi}}{da} = -2\frac{T_{\chi}}{a} + \frac{2\dot{Q}_{\chi}}{3aH}, \qquad \qquad \frac{dT_b}{da} = -2\frac{T_b}{a} + \frac{\Gamma_C}{aH}(T_{\gamma} - T_b) + \frac{2\dot{Q}_b}{3aH}$$

Tashiro, Kadota, Silk 2014: DM can act as heat sink and cool neutral hydrogen in dark ages

Munoz, Kovetz, Ali-Haimoud 2015: included heating due to baryon-DM relative velocities and pointed out implications for global 21-cm signal

$$\dot{Q}_b = F(V_{\chi b})(T_{\chi} - T_b) - \frac{\rho_{\chi}}{\rho_m} \frac{m_{\chi} m_b}{m_{\chi} + m_b} \frac{d}{dt} \left(\frac{1}{2} V_{\chi b}^2\right)$$





Mill-charged—DM explanation very tightly constrained (Boddy et al., 2018; Kovetz et al., 2018)

# **New results** (Creque-Sarbinowski, Ji, Kovetz, MK 1903.09154)

Relic abundance fixed by charge&mass

→ rules out millicharged-DM parameter space for EDGES

Dark-matter decay and lineintensity mapping (Creque-Sarbinowski & MK, arXiv:1806.11119)

### Intensity mapping (review: Kovetz et al. 1709.09066)

Measure sky brightness of some emission line as function of angular position and frequency (a proxy for distance)
→ 3d distribution of emitters



#### Fig credit: Patrick Breysse



#### Fig. from Patrick Breysse and Ely Kovetz



### DM decay

• If DM decays to photon line, decay line will be correlated with large-scale structure



### 1. Dark Matter, the Equivalence Principle, and Dwarf Galaxies

Kesden & MK 2006

# Force

*inertial* acceleration mass

 $m_i a$ 

### Gravitational force



Question: does dark matter fall same way in a gravitational field as ordinary matter?

Conclusions/Summary: We have shown that it does, even though we don't even know what it is.

#### Here's how:

Sagittarius dwarf galaxy

- small galaxy (~billion stars plus lots of dark matter held together by own self-gravity)
- •A satellite of Milky Way (in orbit around Milky Way)
- •Has "tidal tails" streams of stars, stripped from the galaxy, running ahead and behind the galaxy

![](_page_57_Figure_0.jpeg)

#### Where do tidal streams come from?

![](_page_58_Figure_1.jpeg)

![](_page_59_Figure_0.jpeg)

![](_page_60_Figure_0.jpeg)

# What's going on?

![](_page_62_Figure_0.jpeg)

By-eye analysis: Dark matter falls same way as ordinary matter to better than 10%

#### **Conclusions:**

- Parameter space for canonical WIMP shrinking ---time to think anew?
- ~30-Msun PBHs face challenges: now guilty until prov en innocent
- EDGES signal is very intriguing, but cooling of hydrogen by scattering from DM hard to come by
- Intensity mapping provides one new astrophysical tool in arsenal of DM seekers
- Dark matter falls the same way as ordinary matter in g ravitational potential wells